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# On The Almost $\eta$-Ricci Solitons On Pseudosymmetric Lorentz Generalized Sasakian Space Forms 

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#### Abstract

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#### Abstract

In this paper, we consider Lorentz generalized Sasakian space forms admitting almost $\eta$-Ricci solitons in some curvature tensors. Ricci pseudosymmetry concepts of Lorentz generalized Sasakian space forms admitting $\eta$-Ricci soliton have introduced according to the choice of some special curvature tensors such as Riemann, concircular, projective, $\mathscr{M}$-projective, $W_{1}$ and $W_{2}$. Then, again according to the choice of the curvature tensor, necessary conditions are given for Lorentz generalized Sasakian space form admitting $\eta$-Ricci soliton to be Ricci semisymmetric. Then some characterizations are obtained and some classifications have made.


## 1. Introduction

The notion of Ricci flow was introduced by Hamilton in 1982. With the help of this concept, Hamilton found the canonical metric on a smooth manifold. Then Ricci flow has become a powerful tool for the study of Riemannian manifolds, especially for those manifolds with positive curvature. Perelman used Ricci flow and it surgery to prove Poincare conjecture in [1,2]. The Ricci flow is a flow is an evolution equation for metrics on a Riemannian manifold defined as follows:

$$
\frac{\partial}{\partial t} g(t)=-2 S(g(t))
$$

A Ricci soliton emerges as the limit of the solitons of the Ricci flow. A solution to the Ricci flow is called Ricci soliton if it moves only by a one parameter group of diffeomorphism and scaling.
During the last two decades, the geometry of Ricci solitons has been the focus of attention of many mathematicians. In particular, it has become more important after Perelman applied Ricci solitons to solve the long standing Poincare conjecture posed in 1904. In [3], Sharma studied the Ricci solitons in contact geometry. Thereafter Ricci solitons in contact metric manifolds have been studied by various authors such as Ashoka et al. in [4,5], Bagewadi et al. in [6], Ingalahalli in [7], Bejan and Crasmareanu in [8], Blaga in [9], Chandra et al. in [10], Chen and Deshmukh in [11], Deshmukh et al. in [12], He and Zhu [13], Atçeken et al. in [14], Nagaraja and Premalatta in [15], Tripathi in [16] and many others.
$\phi$-sectional curvature plays the important role for Sasakian manifold. If the $\phi$-sectional curvature of a Sasakian manifold is constant, then the manifold is a Sasakian-space-form [17]. P. Alegre and D. Blair described generalized Sasakian space forms [18]. P. Alegre and D. Blair obtained important properties of generalized Sasakian space forms in their studies and gave some examples. P. Alegre and A. Carriazo later discussed generalized indefinite Sasakian space forms [19]. Generalized indefinite Sasakian space forms are also called Lorentz-Sasakian space forms, and Lorentz manifolds are of great importance for Einstein's theory of Relativity.
In this paper, we consider Lorentz generalized Sasakian space forms admitting almost $\eta$-Ricci solitons in some curvature tensors. Ricci pseudosymmetry concepts of Lorentz generalized Sasakian space forms admitting $\eta$-Ricci soliton have introduced according to the choice of some special curvature tensors such as Riemann, concircular, projective, $\mathscr{M}$ - projective, $W_{1}$ and $W_{2}$. Then, again according to the choice of the curvature tensor, necessary conditions are given for Lorentz generalized Sasakian space form admitting $\eta$-Ricci soliton to be Ricci semisymmetric. Then some characterizations are obtained and some classifications have made.

## 2. Preliminaries

Let $\tilde{M}$ be a $(2 n+1)$-dimensional semi-Riemannian manifold. If the $\tilde{M}$ semi-Riemannian manifold with $(\phi, \xi, \eta, g)$ structure tensors satisfies the following conditions, this manifold is called $\varepsilon$ - almost contact metric manifold and $(\phi, \xi, \eta)$ triple is called almost contact structure.

$$
\begin{aligned}
& \phi \xi=0, \quad \eta \circ \phi=0, \quad \eta(\xi)=1, \quad \phi^{2}=-I d+\eta \otimes \xi \\
& g\left(Y_{1}, Y_{2}\right)=g\left(\phi Y_{1}, \phi Y_{2}\right)+\varepsilon \eta\left(Y_{1}\right) \eta\left(Y_{2}\right), \quad \eta\left(Y_{1}\right)=\varepsilon g\left(Y_{1}, \xi\right)
\end{aligned}
$$

where

$$
\varepsilon=g(\xi, \xi)= \pm 1
$$

If $d \eta$ and $g$ provide the relation

$$
d \eta\left(Y_{1}, Y_{2}\right)=g\left(Y_{1}, \phi Y_{2}\right)
$$

then $\tilde{M}$ is called a contact pseudometric manifold and the $(\phi, \xi, \eta)$ triple is called a contact structure.
Let be define a $\left(h\left(\frac{d}{d Y_{1}}\right), Y_{2}\right)$ vector field on $\mathbb{R} \times \tilde{M}$, where $Y_{1}$ is a coordinate on $\mathbb{R}$ and $h$ is a $C^{\infty}$ function on $\mathbb{R} \times \tilde{M}$. The structure defined as

$$
J\left(h \frac{d}{d Y_{1}}, Y_{2}\right)=\left(\eta\left(Y_{2}\right) \frac{d}{d Y_{1}}, \phi Y_{2}-h \xi\right)
$$

on $\mathbb{R} \times \tilde{M}$ is called a almost complex structure and $J^{2}=-i d$. If $J$ is integrable, the almost contact structure $(\phi, \xi, \eta)$ is said to be normal. If $Y_{1}$ is perpendicular to $\xi$, the plane spanned by $Y_{1}$ and $\phi Y_{1}$, is called the $\phi$-section. The curvature of the $\phi$-section is called the $\phi$-sectional curvature. The curvature of the indefinite Sasakian manifold defined in this way is precisely determined by the $\phi$-section curvature. If the $\phi-$ section curvature of the indefinite Sasakian manifold is equal to a constant $c$, the curvature tensor of this manifold is defined as

$$
\begin{aligned}
\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{3}= & \left(\frac{c+3 \varepsilon}{4}\right)\left\{g\left(Y_{2}, Y_{3}\right) Y_{1}-g\left(Y_{1}, Y_{3}\right) Y_{2}\right\}+\left(\frac{c-\varepsilon}{4}\right)\left\{g\left(Y_{1}, \phi Y_{3}\right) \phi Y_{2}-g\left(Y_{2}, \phi Y_{3}\right) \phi Y_{1}+2 g\left(Y_{1}, \phi Y_{2}\right) \phi Y_{3}\right\} \\
& +\left(\frac{c-\varepsilon}{4}\right)\left\{\eta\left(Y_{1}\right) \eta\left(Y_{3}\right) Y_{2}-\eta\left(Y_{2}\right) \eta\left(Y_{3}\right) Y_{1}+\varepsilon g\left(Y_{1}, Y_{3}\right) \eta\left(Y_{2}\right) \xi-\varepsilon g\left(Y_{2}, Y_{3}\right) \eta\left(Y_{1}\right) \xi\right\} .
\end{aligned}
$$

For an $\varepsilon$-almost contact metric manifold $\tilde{M}$, if there are $\digamma_{1}, \digamma_{2}, \digamma_{3} \in C^{\infty}(\tilde{M})$ functions such that

$$
\begin{aligned}
\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{3}= & \digamma_{1}\left\{g\left(Y_{2}, Y_{3}\right) Y_{1}-g\left(Y_{1}, Y_{3}\right) Y_{2}\right\}+\digamma_{2}\left\{g\left(Y_{1}, \phi Y_{3}\right) \phi Y_{2}-g\left(Y_{2}, \phi Y_{3}\right) \phi Y_{1}+2 g\left(Y_{1}, \phi Y_{2}\right) \phi Y_{3}\right\} \\
& +\digamma_{3}\left\{\eta\left(Y_{1}\right) \eta\left(Y_{3}\right) Y_{2}-\eta\left(Y_{2}\right) \eta\left(Y_{3}\right) Y_{1}+\varepsilon g\left(Y_{1}, Y_{3}\right) \eta\left(Y_{2}\right) \xi-\varepsilon g\left(Y_{2}, Y_{3}\right) \eta\left(Y_{1}\right) \xi\right\}
\end{aligned}
$$

then manifold $\tilde{M}$ is called a generalized indefinite Sasakian space form.
In this article, only the Lorentzian case, which corresponds to the $\varepsilon=-1$, where the index of the metric is 1 , will be discussed. Such manifolds are called Lorentz generalized Sasakian space forms and are denoted by $M^{2 n+1}\left(\digamma_{1}, \digamma_{2}, \digamma_{3}\right)$. Thus, the curvature tensor of a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form is defined as

$$
\begin{align*}
\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{3}= & \digamma_{1}\left\{g\left(Y_{2}, Y_{3}\right) Y_{1}-g\left(Y_{1}, Y_{3}\right) Y_{2}\right\}+\digamma_{2}\left\{g\left(Y_{1}, \phi Y_{3}\right) \phi Y_{2}-g\left(Y_{2}, \phi Y_{3}\right) \phi Y_{1}+2 g\left(Y_{1}, \phi Y_{2}\right) \phi Y_{3}\right\} \\
& +\digamma_{3}\left\{\eta\left(Y_{1}\right) \eta\left(Y_{3}\right) Y_{2}-\eta\left(Y_{2}\right) \eta\left(Y_{3}\right) Y_{1}-g\left(Y_{1}, Y_{3}\right) \eta\left(Y_{2}\right) \xi+g\left(Y_{2}, Y_{3}\right) \eta\left(Y_{1}\right) \xi\right\} . \tag{2.1}
\end{align*}
$$

Lemma 2.1. Let $M^{2 n+1}\left(\digamma_{1}, \digamma_{2}, \digamma_{3}\right)$ be the $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. The following relations are provided for $M^{2 n+1}\left(\digamma_{1}, \digamma_{2}, \digamma_{3}\right)$.

$$
\begin{align*}
\tilde{\nabla}_{Y_{1}} \xi & =\left(\digamma_{1}+\digamma_{3}\right) \phi Y_{1},  \tag{2.2}\\
\tilde{R}\left(Y_{1}, \xi\right) Y_{3} & =-\left(\digamma_{1}+\digamma_{3}\right)\left[g\left(Y_{1}, Y_{3}\right) \xi+\eta\left(Y_{3}\right) Y_{1}\right],  \tag{2.3}\\
\tilde{R}\left(\xi, Y_{2}\right) Y_{3} & =\left(\digamma_{1}+\digamma_{3}\right)\left[g\left(Y_{2}, Y_{3}\right) \xi+\eta\left(Y_{3}\right) Y_{2}\right],  \tag{2.4}\\
\tilde{R}\left(Y_{1}, Y_{2}\right) \xi & =\left(\digamma_{1}+\digamma_{3}\right)\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right],  \tag{2.5}\\
\eta\left(\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{3}\right) & =\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{3}\right),  \tag{2.6}\\
S\left(Y_{1}, Y_{2}\right) & =\left(2 n \digamma_{1}+3 \digamma_{2}+\digamma_{3}\right) g\left(Y_{1}, Y_{2}\right)+\left(3 \digamma_{2}-(2 n-1) \digamma_{3}\right) \eta\left(Y_{1}\right) \eta\left(Y_{2}\right),  \tag{2.7}\\
S\left(Y_{1}, \xi\right) & =-2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(Y_{1}\right),  \tag{2.8}\\
Q Y_{1} & =\left(2 n \digamma_{1}+3 \digamma_{2}+\digamma_{3}\right) Y_{1}+\left((2 n-1) \digamma_{3}-3 \digamma_{2}\right),  \tag{2.9}\\
Q \xi & =2 n\left(\digamma_{1}+\digamma_{3}\right) \xi, \tag{2.10}
\end{align*}
$$

where $\tilde{R}, S$ and $Q$ are the Riemann curvature tensor, Ricci curvature tensor and Ricci operator of $M^{2 n+1}\left(\digamma_{1}, \digamma_{2}, \digamma_{3}\right)$, respectively.
Let $M$ be a Riemannian manifold, $T$ is $(0, k)$-type tensor field and $A$ is $(0,2)$-type tensor field. In this case, Tachibana tensor field $Q(A, T)$ is defined as

$$
\begin{equation*}
Q(A, T)\left(X_{1}, \ldots, X_{K} ; Y_{1}, Y_{2}\right)=-T\left(\left(Y_{1} \wedge_{A} Y_{2}\right) X_{1}, \ldots, X_{k}\right)-\ldots-T\left(X_{1}, \ldots, X_{k-1},\left(Y_{1} \wedge_{A} Y_{2}\right) X_{k}\right) \tag{2.11}
\end{equation*}
$$

where

$$
\begin{equation*}
\left(Y_{1} \wedge_{A} Y_{2}\right) Y_{3}=A\left(Y_{2}, Y_{3}\right) Y_{1}-A\left(Y_{1}, Y_{3}\right) Y_{2}, \tag{2.12}
\end{equation*}
$$

$k \geq 1, X_{1}, X_{2}, \ldots, X_{k}, Y_{1}, Y_{2} \in \Gamma(T M)$.
Precisely, a Ricci soliton on a Riemannian manifold $(\tilde{M}, g)$ is defined as a triple $(g, \xi, \lambda)$ on $\tilde{M}$ satisfying

$$
\begin{equation*}
L_{\xi} g+2 S+2 \lambda g=0 \tag{2.13}
\end{equation*}
$$

where $L_{\xi}$ is the Lie derivative operator along the vector field $\xi$ and $\lambda$ is a real constant. We note that if $\xi$ is a Killing vector field, then the Ricci soliton reduces to an Einstein metric $(g, \lambda)$. Futhermore, in [20], generalization is the notion of $\eta$-Ricci soliton defined by J.T. Cho and M. Kimura as a quadruple $(g, \xi, \lambda, \mu)$ satisfying

$$
\begin{equation*}
L_{\xi} g+2 S+2 \lambda g+2 \mu \eta \oplus \eta=0 \tag{2.14}
\end{equation*}
$$

where $\lambda$ and $\mu$ are real constants and $\eta$ is the dual of $\xi$ and $S$ denotes the Ricci tensor of $\tilde{M}$. Furthermore if $\lambda$ and $\mu$ are smooth functions on $\tilde{M}$, then it called almost $\eta$-Ricci soliton on $\tilde{M}$ [20].
Suppose the quartet $(g, \xi, \lambda, \mu)$ is almost $\eta$-Ricci soliton on manifold $\tilde{M}$. Then,

- If $\lambda<0$, then $\tilde{M}$ is shriking.
- If $\lambda=0$, then $\tilde{M}$ is steady.
- If $\lambda>0$, then $\tilde{M}$ is expanding.


## 3. Almost $\eta$-Ricci Solitons on Ricci Pseudosymmetric and Ricci Semisymmetric Lorentz Generalized Sasakian Space Forms

Now let $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form. Then we have

$$
\begin{aligned}
\left(L_{\xi} g\right)\left(Y_{1}, Y_{2}\right) & =L_{\xi} g\left(Y_{1}, Y_{2}\right)-g\left(L_{\xi} Y_{1}, Y_{2}\right)-g\left(Y_{1}, L_{\xi} Y_{2}\right) \\
& =\xi g\left(Y_{1}, Y_{2}\right)-g\left(\left[\xi, Y_{1}\right], Y_{2}\right)-g\left(Y_{1},\left[\xi, Y_{2}\right]\right) \\
& =g\left(\nabla_{\xi} Y_{1}, Y_{2}\right)+g\left(Y_{1}, \nabla_{\xi} Y_{2}\right)-g\left(\nabla_{\xi} Y_{1}, Y_{2}\right)+g\left(\nabla_{Y_{1}} \xi, Y_{2}\right)-g\left(\nabla_{\xi} Y_{2}, Y_{1}\right)+g\left(Y_{1}, \nabla_{Y_{2}} \xi\right),
\end{aligned}
$$

for all $Y_{1}, Y_{2} \in \Gamma(T M)$. By using $\phi$ is anti-symmetric, we have

$$
\begin{equation*}
\left(L_{\xi} g\right)\left(Y_{1}, Y_{2}\right)=0 \tag{3.1}
\end{equation*}
$$

Thus, in a Lorentz generalized Sasakian space form, from (2.14) and (3.1), we have

$$
\begin{equation*}
S\left(Y_{1}, Y_{2}\right)+\lambda g\left(Y_{1}, Y_{2}\right)+\mu \eta\left(Y_{1}\right) \eta\left(Y_{2}\right)=0 \tag{3.2}
\end{equation*}
$$

It is clear from (16) that the $(2 n+1)$-dimensional Lorentz generalized Sasakian admitting almost $\eta$-Ricci soliton $\left(M^{2 n+1}, g, \xi, \lambda, \mu\right)$ is an $\eta$-Einstein manifold.
For $Y_{2}=\xi$ in (3.2), this implies that

$$
\begin{equation*}
S\left(\xi, Y_{1}\right)=(\lambda-\mu) \eta\left(Y_{1}\right) \tag{3.3}
\end{equation*}
$$

Taking into account of (3.3), we conclude that

$$
\mu-\lambda=2 n\left(\digamma_{1}+\digamma_{3}\right) .
$$

Definition 3.1. Let $M^{2 n+1}$ be an $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. If $\tilde{R} \cdot S$ and $Q(g, S)$ are linearly dependent, then the $M^{2 n+1}$ is said to be Ricci pseudosymmetric.
In this case, there exists a function $L_{1}$ on $M^{2 n+1}$ such that

$$
\tilde{R} \cdot S=L_{1} Q(g, S) .
$$

In particular, if $L_{1}=0$, the manifold $M^{2 n+1}$ is said to be Ricci semisymmetric.
Let us now investigate the Ricci pseudosymmetric case of the ( $2 n+1$ ) -dimensional Lorentz generalized Sasakian space forms.
Theorem 3.2. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space forms and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a Ricci pseudosymmetric, then

$$
L_{1}=\frac{\left(\digamma_{1}+\digamma_{3}\right)\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]}{\mu}
$$

provided $\mu \neq 0$.
Proof. Let be assume that Lorentz generalized Sasakian space form $M^{2 n+1}$ be Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space forms $M^{2 n+1}$. That is mean

$$
\left(\tilde{R}\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{1} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right),
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, \tilde{R}\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{1}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} \tag{3.4}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.4), we get

$$
\begin{equation*}
S\left(\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, \tilde{R}\left(Y_{1}, Y_{2}\right) \xi\right)=L_{1}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.5}
\end{equation*}
$$

If we make use of (2.5) and (2.8) in (3.5), we have

$$
\begin{align*}
& S\left(Y_{4},\left(\digamma_{1}+\digamma_{3}\right)\left[\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}\right]\right)-2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(\tilde{R}\left(Y_{1}, Y_{2}\right) Y_{4}\right) \\
& =L_{1}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.6}
\end{align*}
$$

If we use (2.6) in the (3.6), we get

$$
\begin{aligned}
& -2 n\left(\digamma_{1}+\digamma_{3}\right)^{2} g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)+\left(\digamma_{1}+\digamma_{3}\right) S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right) \\
& =L_{1}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\}
\end{aligned}
$$

If we use (3.2) in (3.5), we can write

$$
\begin{equation*}
\left[\left(\digamma_{1}+\digamma_{3}\right)\left[2 n\left(\digamma_{1}+\digamma_{3}\right)-\lambda\right]+\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right] L_{1}\right] \times g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)=0 \tag{3.7}
\end{equation*}
$$

It is clear from (3.7)

$$
L_{1}=\frac{\left(\digamma_{1}+\digamma_{3}\right)\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]}{\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)}
$$

This completes the proof.
We can give the results obtained from this theorem as follows.
Corollary 3.3. Let $M^{2 n+1}$ be a Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a Ricci semisymmetric, then $\lambda=2 n\left(\digamma_{1}+\digamma_{3}\right)$ and $\mu=4 n\left(\digamma_{1}+\digamma_{3}\right)$.
Corollary 3.4. Let $M^{2 n+1}$ be a Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a Ricci semisymmetric, then $M^{2 n+1}$ is an $\eta$-Einstein manifold.

Corollary 3.5. Let $M^{2 n+1}$ be a Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a Ricci semisymmetric, then we observe that:
(i) $M^{2 n+1}$ is expanding, if $\digamma_{1}+\digamma_{3}>0$.
(ii) $M^{2 n+1}$ is shriking, if $\digamma_{1}+\digamma_{3}<0$.

For a $(2 n+1)$-dimensional semi-Riemann manifold $M$, the concircular curvature tensor is defined as

$$
\begin{equation*}
C\left(Y_{1}, Y_{2}\right) Y_{3}=R\left(Y_{1}, Y_{2}\right) Y_{3}-\frac{r}{2 n(2 n+1)}\left[g\left(Y_{2}, Y_{3}\right) Y_{1}-g\left(Y_{1}, Y_{3}\right) Y_{2}\right] \tag{3.8}
\end{equation*}
$$

For a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form, if we choose $Y_{3}=\xi$ in (3.8), we can write

$$
\begin{equation*}
C\left(Y_{1}, Y_{2}\right) \xi=\left[\left(\digamma_{1}+\digamma_{3}\right)-\frac{r}{2 n(2 n+1)}\right]\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right] \tag{3.9}
\end{equation*}
$$

and similarly if we take the inner product of both sides of (24) by $\xi$, we get

$$
\begin{equation*}
\eta\left(C\left(Y_{1}, Y_{2}\right) Y_{3}\right)=\left[\left(\digamma_{1}+\digamma_{3}\right)-\frac{r}{2 n(2 n+1)}\right] g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{3}\right) \tag{3.10}
\end{equation*}
$$

Definition 3.6. Let $M^{2 n+1}$ be a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. If $C \cdot S$ and $Q(g, S)$ are linearly dependent, then the manifold is said to be concircular Ricci pseudosymmetric.
In this case, there exists a function $L_{2}$ on $M^{2 n+1}$ such that

$$
C \cdot S=L_{2} Q(g, S)
$$

In particular, if $L_{2}=0$, the manifold $M^{2 n+1}$ is said to be concircular Ricci semisymmetric.
Let us now investigate the concircular Ricci pseudosymmetric case of the Lorentz generalized Sasakian space form.
Theorem 3.7. Let $M^{2 n+1}$ be a Lorentz generalized Sasakian space forms and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a concircular Ricci pseudosymmetric, then

$$
L_{2}=\frac{\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]\left[2 n(2 n+1)\left(\digamma_{1}+\digamma_{3}\right)-r\right]}{2 n(2 n+1) \mu}
$$

provided $\mu \neq 0$.

Proof. Let be assume that Lorentz generalized Sasakian space form $M^{2 n+1}$ be concircular Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form $M^{2 n+1}$. That is mean

$$
\left(C\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{2} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right),
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(C\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, C\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{2}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} . \tag{3.11}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.11), we get

$$
\begin{equation*}
S\left(C\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, C\left(Y_{1}, Y_{2}\right) \xi\right)=L_{2}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.12}
\end{equation*}
$$

By using of (2.8) and (3.9) in (3.12), we have

$$
\begin{align*}
& S\left(Y_{4}, A\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right]\right)-2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(C\left(Y_{1}, Y_{2}\right) Y_{4}\right)  \tag{3.13}\\
& =L_{2}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\},
\end{align*}
$$

where $A=\left(\digamma_{1}+\digamma_{3}\right)-\frac{r}{2 n(2 n+1)}$. Substituting (3.10) into (3.13), we have

$$
\begin{align*}
& -2 n\left(\digamma_{1}+\digamma_{3}\right) A g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)+A S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right) \\
& =L_{2}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)\right\} . \tag{3.14}
\end{align*}
$$

If we use (3.2) in the (3.14), we can write

$$
\begin{equation*}
\left\{A\left[2 n\left(\digamma_{1}+\digamma_{3}\right)-\lambda\right]+\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right] L_{2}\right\} \times g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)=0 . \tag{3.15}
\end{equation*}
$$

It is clear from (3.15),

$$
L_{2}=\frac{\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]\left[2 n(2 n+1)\left(\digamma_{1}+\digamma_{3}\right)-r\right]}{2 n(2 n+1) \mu} .
$$

This completes the proof.
We can give the results obtained from this theorem as follows.
Corollary 3.8. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a concircular Ricci semisymmetric, then $M^{2 n+1}$ is either manifold with scalar curvature $r=2 n(2 n+1)\left(\digamma_{1}+\digamma_{3}\right)$ or $\lambda=2 n\left(\digamma_{1}+\digamma_{3}\right)$.
Corollary 3.9. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a concircular Ricci semisymmetric, then we observe that:
(i) The soliton $M^{2 n+1}$ is expanding, if $\left(\digamma_{1}+\digamma_{3}\right)>0$.
(ii) The soliton $M^{2 n+1}$ is shriking, if $\left(\digamma_{1}+\digamma_{3}\right)<0$.

For a $(2 n+1)$-dimensional semi-Riemann manifold $M$, the projective curvature tensor is defined as

$$
\begin{equation*}
P\left(Y_{1}, Y_{2}\right) Y_{3}=R\left(Y_{1}, Y_{2}\right) Y_{3}-\frac{1}{2 n}\left[S\left(Y_{2}, Y_{3}\right) Y_{1}-S\left(Y_{1}, Y_{3}\right) Y_{2}\right] . \tag{3.16}
\end{equation*}
$$

For a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form, if we choose $Y_{3}=\xi$ in (3.16) we can write

$$
\begin{equation*}
P\left(Y_{1}, Y_{2}\right) \xi=0, \tag{3.17}
\end{equation*}
$$

and similarly if we take the inner product of both sides of (3.16) by $\xi$, we get

$$
\begin{equation*}
\eta\left(P\left(Y_{1}, Y_{2}\right) Y_{3}\right)=0 . \tag{3.18}
\end{equation*}
$$

Definition 3.10. Let $M^{2 n+1}$ be an $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. If $P \cdot S$ and $Q(g, S)$ are linearly dependent, then the manifold is said to be projective Ricci pseudosymmetric.

In this case, there exists a function $L_{3}$ on $M^{2 n+1}$ such that

$$
P \cdot S=L_{3} Q(g, S) .
$$

In particular, if $L_{3}=0$, the manifold $M^{2 n+1}$ is said to be projective Ricci semisymmetric.
Let us now investigate the projective Ricci pseudosymmetry case of the Lorentz generalized Sasakian space form.
Theorem 3.11. Let $M^{2 n+1}$ be Lorentz Sasakian space forms and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a projective Ricci pseudosymmetric, then $M^{2 n+1}$ is either projective Ricci semisymmetric or almost $\eta$-Ricci soliton $(g, \xi, \lambda, \mu)$ reduces almost Ricci soliton ( $\mathrm{g}, \boldsymbol{\xi}, \lambda$ ).

Proof. Let be assume that Lorentz generalized Sasakian space form $M^{2 n+1}$ be projective Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form $M^{2 n+1}$. Then we have

$$
\left(P\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{3} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right),
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(P\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, P\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{3}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} \tag{3.19}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.19), we get

$$
\begin{equation*}
S\left(P\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, P\left(Y_{1}, Y_{2}\right) \xi\right)=L_{3}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.20}
\end{equation*}
$$

If we make use of (2.8) and (3.17) in (3.20) we have

$$
\begin{equation*}
-2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(P\left(Y_{1}, Y_{2}\right) Y_{4}\right)=L_{3}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.21}
\end{equation*}
$$

If we use (3.18) in the (3.21), we get

$$
\begin{equation*}
L_{3}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)\right\}=0 . \tag{3.22}
\end{equation*}
$$

If we use (3.2) in the (3.22), we can write

$$
\begin{equation*}
L_{3}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right] g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)=0 . \tag{3.23}
\end{equation*}
$$

It is clear from (3.23),

$$
\mu L_{3}=0 .
$$

This completes the proof.
We can give the results obtained from this theorem as follows.
Corollary 3.12. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a projective Ricci pseudosymmetric, then $M^{2 n+1}$ is either projective Ricci semisymmetric or we observe that:
(i) The soliton $M^{2 n+1}$ is expanding, if $\digamma_{1}+\digamma_{3}<0$.
(ii) The soliton $M^{2 n+1}$ is shriking, if $\digamma_{1}+\digamma_{3}>0$.

For a $(2 n+1)$-dimensional semi-Riemann manifold $M$, the $\mathscr{M}$-projective curvature tensor is defined as

$$
\begin{equation*}
\mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{3}=R\left(Y_{1}, Y_{2}\right) Y_{3}-\frac{1}{2 n}\left[S\left(Y_{2}, Y_{3}\right) Y_{1}-S\left(Y_{1}, Y_{3}\right) Y_{2}+g\left(Y_{2}, Y_{3}\right) Q Y_{1}-g\left(Y_{1}, Y_{3}\right) Q Y_{2}\right] \tag{3.24}
\end{equation*}
$$

For a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form, if we choose $Y_{3}=\xi$ in (3.24), we obtain

$$
\begin{equation*}
\mathscr{M}\left(Y_{1}, Y_{2}\right) \xi=\frac{1}{2 n}\left[\eta\left(Y_{2}\right) Q Y_{1}-\eta\left(Y_{1}\right) Q Y_{2}\right] \tag{3.25}
\end{equation*}
$$

and similarly if we take the inner product of both of sides of (3.24) by $\xi$, we get

$$
\begin{equation*}
\eta\left(\mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{3}\right)=\frac{1}{2 n} S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{3}\right) . \tag{3.26}
\end{equation*}
$$

Definition 3.13. Let $M^{2 n+1}$ be an $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. If $\mathscr{M} \cdot S$ and $Q(g, S)$ are linearly dependent, then the manifold is said to be $\mathscr{M}$-projective Ricci pseudosymmetric.

In this case, there exists a function $L_{4}$ on $M^{2 n+1}$ such that

$$
\mathscr{M} \cdot S=L_{4} Q(g, S) .
$$

In particular, if $L_{4}=0$, the manifold $M^{2 n+1}$ is said to be $\mathscr{M}$-projective Ricci semisymmetric.
Let us now investigate the $\mathscr{M}$-projective Ricci pseudosymmetric case of the Lorentz generalized Sasakian space form.
Theorem 3.14. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space forms and $(g, \xi, \lambda, \mu)$ be almost $\eta-$ Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $\mathscr{M}$-projective Ricci pseudosymmetric, then

$$
L_{4}=\frac{\lambda^{2}-2 n\left(\digamma_{1}+\digamma_{3}\right) \lambda}{2 n \mu},
$$

provided $\mu \neq 0$.

Proof. Let be assume that Lorentz generalized Sasakian space form $M^{2 n+1}$ be projective $\mathscr{M}$-projective Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form $M^{2 n+1}$. That is mean

$$
\left(\mathscr{M}\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{4} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right)
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(\mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, \mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{4}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} \tag{3.27}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.27) we get

$$
\begin{equation*}
S\left(\mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, \mathscr{M}\left(Y_{1}, Y_{2}\right) \xi\right)=L_{4}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.28}
\end{equation*}
$$

If we make use of (2.8) and (3.25) in (3.28), we have

$$
\begin{align*}
& -2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(\mathscr{M}\left(Y_{1}, Y_{2}\right) Y_{4}\right)+S\left(Y_{4}, \frac{1}{2 n}\left[\eta\left(Y_{2}\right) Q Y_{1}-\eta\left(Y_{1}\right) Q Y_{2}\right]\right)  \tag{3.29}\\
& =L_{4}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\}
\end{align*}
$$

By using (3.26) in the (3.29), we get

$$
\begin{align*}
& -\left(\digamma_{1}+\digamma_{3}\right) S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+\frac{1}{2 n} S\left(\eta\left(Y_{2}\right) Q Y_{1}-\eta\left(Y_{1}\right) Q Y_{2}, Y_{4}\right)  \tag{3.30}\\
& =L_{4}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)\right\}
\end{align*}
$$

If we put (3.2) in (3.30), we can write

$$
\begin{align*}
& \lambda\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)-\frac{\lambda}{2 n} S\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)  \tag{3.31}\\
& =L_{4}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right] g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)
\end{align*}
$$

Again, if we use (3.2) in the (3.31), we obtain

$$
\begin{equation*}
\left\{\frac{\lambda^{2}}{2 n}-\left(\digamma_{1}+\digamma_{3}\right) \lambda-L_{4}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right]\right\} \times g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)=0 \tag{3.32}
\end{equation*}
$$

It is clear from (3.32),

$$
L_{4}=\frac{\lambda^{2}-2 n\left(\digamma_{1}+\digamma_{3}\right) \lambda}{2 n\left[2 n\left(\digamma_{1}+\digamma_{3}\right)+\lambda\right]}
$$

This completes the proof.
We can give the following corollaries.
Corollary 3.15. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is $a \mathscr{M}$-projective Ricci semisymmetric, then $M^{2 n+1}$ is either steady or $\eta$-Einstein manifold.
Corollary 3.16. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $\mathscr{M}$-projective Ricci semisymmetric, then $M^{2 n+1}$ is either steady or we observe that:
(i) The soliton $M^{2 n+1}$ is shriking if $\lambda$ is between 0 and $2 n\left(\digamma_{1}+\digamma_{3}\right)$.
(ii) The soliton $M^{2 n+1}$ is steady if $\lambda=0$.
(iii) The soliton $M^{2 n+1}$ is expanding for other cases of $\lambda$.

For a $(2 n+1)$-dimensional semi-Riemann manifold $M$, the $W_{1}$-curvature tensor is defined as

$$
\begin{equation*}
W_{1}\left(Y_{1}, Y_{2}\right) Y_{3}=R\left(Y_{1}, Y_{2}\right) Y_{3}+\frac{1}{2 n}\left[S\left(Y_{2}, Y_{3}\right) Y_{1}-S\left(Y_{1}, Y_{3}\right) Y_{2}\right] \tag{3.33}
\end{equation*}
$$

For a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form, if we choose $Y_{3}=\xi$ in (3.33), we can write

$$
\begin{equation*}
W_{1}\left(Y_{1}, Y_{2}\right) \xi=2\left(\digamma_{1}+\digamma_{3}\right)\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right] \tag{3.34}
\end{equation*}
$$

and similarly if we take the inner product of both of sides of (3.33) by $\xi$, we get

$$
\begin{equation*}
\eta\left(W_{1}\left(Y_{1}, Y_{2}\right) Y_{3}\right)=2\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{3}\right) \tag{3.35}
\end{equation*}
$$

Definition 3.17. Let $M^{2 n+1}$ be a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form. If $W_{1} \cdot S$ and $Q(g, S)$ are linearly dependent, then the manifold is said to be $W_{1}-$ Ricci pseudosymmetric.

In this case, there exists a function $L_{5}$ on $M^{2 n+1}$ such that

$$
W_{1} \cdot S=L_{5} Q(g, S)
$$

In particular, if $L_{5}=0$, the manifold $M^{2 n+1}$ is said to be $W_{1}$-Ricci semisymmetric.
Let us now investigate the $W_{1}$-Ricci pseudosymmetric case of the Lorentz generalized Sasakian space form.

Theorem 3.18. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space forms and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $W_{1}$-Ricci pseudosymmetric, then

$$
L_{5}=\frac{2\left(\digamma_{1}+\digamma_{3}\right)\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]}{\mu}
$$

provided $\mu \neq 0$.
Proof. Let be assume that Lorentz Sasakian space form $M^{2 n+1}$ be $W_{1}-$ Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form $M^{2 n+1}$. That is mean

$$
\left(W_{1}\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{5} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right)
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(W_{1}\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, W_{1}\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{5}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} \tag{3.36}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.36), we get

$$
\begin{equation*}
S\left(W_{1}\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, W_{1}\left(Y_{1}, Y_{2}\right) \xi\right)=L_{5}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} \tag{3.37}
\end{equation*}
$$

If we make use of (2.8) and (3.34) in (3.37) we have

$$
\begin{align*}
& 2\left(\digamma_{1}+\digamma_{3}\right) S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)-2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(W_{1}\left(Y_{1}, Y_{2}\right) Y_{4}\right)  \tag{3.38}\\
& =L_{5}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\}
\end{align*}
$$

If we use (3.35) in the (3.38), we get

$$
\begin{align*}
& -4 n\left(\digamma_{1}+\digamma_{3}\right)^{2} g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)+2\left(\digamma_{1}+\digamma_{3}\right) S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)  \tag{3.39}\\
& =L_{5}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)\right\}
\end{align*}
$$

If we use (3.2) in the (3.39), we can write

$$
\begin{equation*}
\left\{2\left(\digamma_{1}+\digamma_{3}\right)\left[2 n\left(\digamma_{1}+\digamma_{3}-\lambda\right)\right]+L_{5}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right]\right\} \times g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)=0 \tag{3.40}
\end{equation*}
$$

It is clear from (3.40),

$$
L_{5}=\frac{2\left(\digamma_{1}+\digamma_{3}\right)\left[\lambda-2 n\left(\digamma_{1}+\digamma_{3}\right)\right]}{\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)}
$$

This completes the proof.
We can give the results obtained from this theorem as follows.
Corollary 3.19. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $W_{1}$-Ricci semisymmetric, then $\lambda=2 n\left(\digamma_{1}+\digamma_{3}\right)$ provided $\mu \neq 0$.

Corollary 3.20. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $W_{1}$-Ricci semisymmetric, then we observe that:
(i) The soliton $M^{2 n+1}$ is expanding, if $\left(\digamma_{1}+\digamma_{3}\right)>0$.
(ii) The soliton $M^{2 n+1}$ is shriking, if $\left(\digamma_{1}+\digamma_{3}\right)<0$.

For a $(2 n+1)$-dimensional semi-Riemann manifold $M$, the $W_{2}$-curvature tensor is defined as

$$
\begin{equation*}
W_{2}\left(Y_{1}, Y_{2}\right) Y_{3}=R\left(Y_{1}, Y_{2}\right) Y_{3}-\frac{1}{2 n}\left[g\left(Y_{2}, Y_{3}\right) Q Y_{1}-g\left(Y_{1}, Y_{3}\right) Q Y_{2}\right] \tag{3.41}
\end{equation*}
$$

For a $(2 n+1)$-dimensional Lorentz generalized Sasakian space form, if we choose $Y_{3}=\xi$ in (3.41), we can write

$$
\begin{equation*}
W_{2}\left(Y_{1}, Y_{2}\right) \xi=\left(\digamma_{1}+\digamma_{3}\right)\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right]-\frac{1}{2 n}\left[\eta\left(Y_{1}\right) Q Y_{2}-\eta\left(Y_{2}\right) Q Y_{1}\right] \tag{3.42}
\end{equation*}
$$

and similarly if we take the inner product of both sides of (3.41) by $\xi$, we get

$$
\begin{equation*}
\eta\left(W_{2}\left(Y_{1}, Y_{2}\right) Y_{3}\right)=\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{3}\right)+\frac{1}{2 n} S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{3}\right) \tag{3.43}
\end{equation*}
$$

Definition 3.21. Let $M^{2 n+1}$ be an $(2 n+1)$-dimensional Lorentz, generalized Sasakian space form. If $W_{2} \cdot S$ and $Q(g, S)$ are linearly dependent, then the manifold is said to be $W_{2}-$ Ricci pseudosymmetric.
In this case, there exists a function $L_{6}$ on $M^{2 n+1}$ such that

$$
W_{2} \cdot S=L_{6} Q(g, S)
$$

In particular, if $L_{6}=0$, the manifold $M^{2 n+1}$ is said to be $W_{2}$-Ricci semisymmetric.
Let us now investigate the $W_{2}-$ Ricci pseudosymmetric case of the Lorentz generalized Sasakian space form.

Theorem 3.22. Let $M^{2 n+1}$ be Lorentz generalized Sasakian space form and $\left(g, \xi, \kappa_{1}, \kappa_{2}\right)$ be almost $\eta$-Ricci soliton on $M^{2 n+1}$. If $M^{2 n+1}$ is a $W_{2}$-Ricci pseudosymmetric, then

$$
L_{6}=-\frac{\lambda^{2}+4 n^{2}\left(\digamma_{1}+\digamma_{3}\right)^{2}}{2 n \mu}
$$

provided $\mu \neq 0$.
Proof. Let be assume that Lorentz generalized Sasakian space form be $W_{2}-$ Ricci pseudosymmetric and $(g, \xi, \lambda, \mu)$ be almost $\eta$-Ricci soliton on Lorentz generalized Sasakian space form. That is mean

$$
\left(W_{2}\left(Y_{1}, Y_{2}\right) \cdot S\right)\left(Y_{4}, Y_{5}\right)=L_{6} Q(g, S)\left(Y_{4}, Y_{5} ; Y_{1}, Y_{2}\right)
$$

for all $Y_{1}, Y_{2}, Y_{4}, Y_{5} \in \Gamma\left(T M^{2 n+1}\right)$. From the last equation, we can easily write

$$
\begin{equation*}
S\left(W_{2}\left(Y_{1}, Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4}, W_{2}\left(Y_{1}, Y_{2}\right) Y_{5}\right)=L_{6}\left\{S\left(\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{4}, Y_{5}\right)+S\left(Y_{4},\left(Y_{1} \wedge_{g} Y_{2}\right) Y_{5}\right)\right\} \tag{3.44}
\end{equation*}
$$

If we choose $Y_{5}=\xi$ in (3.44), we get

$$
\begin{equation*}
S\left(W_{2}\left(Y_{1}, Y_{2}\right) Y_{4}, \xi\right)+S\left(Y_{4}, W_{2}\left(Y_{1}, Y_{2}\right) \xi\right)=L_{6}\left\{S\left(g\left(Y_{2}, Y_{4}\right) Y_{1}-g\left(Y_{1}, Y_{4}\right) Y_{2}, \xi\right)+S\left(Y_{4}, \eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}\right)\right\} \tag{3.45}
\end{equation*}
$$

If we make use of (2.8) and (3.42) in (3.45), we have

$$
\begin{align*}
& -2 n\left(\digamma_{1}+\digamma_{3}\right) \eta\left(W_{2}\left(Y_{1}, Y_{2}\right) Y_{4}\right)+S\left(Y_{4},\left(\digamma_{1}+\digamma_{3}\right)\left[\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right]-\frac{1}{2 n}\left[\eta\left(Y_{1}\right) Q Y_{2}-\eta\left(Y_{2}\right) Q Y_{1}\right]\right)  \tag{3.46}\\
& =L_{6}\left\{-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)+S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)\right\} .
\end{align*}
$$

If we use (3.43) in the (3.46), we get

$$
\begin{align*}
& -2 n\left(\digamma_{1}+\digamma_{3}\right)^{2} g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)-\frac{1}{2 n} S\left(Y_{4}, \eta\left(Y_{1}\right) Q Y_{2}-\eta\left(Y_{2}\right) Q Y_{1}\right)  \tag{3.47}\\
& =L_{6}\left\{S\left(Y_{4}, \eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}\right)-2 n\left(\digamma_{1}+\digamma_{3}\right) g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)\right\}
\end{align*}
$$

If we use (3.2) in the (3.47), we have

$$
\begin{align*}
& -2 n\left(\digamma_{1}+\digamma_{3}\right)^{2} g\left(\eta\left(Y_{2}\right) Y_{1}-\eta\left(Y_{1}\right) Y_{2}, Y_{4}\right)-\frac{\lambda}{2 n} S\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)  \tag{3.48}\\
& =-L_{6}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right] g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)
\end{align*}
$$

Again, if we use (3.2) in (3.48), we obtain

$$
\begin{equation*}
\left\{\frac{\lambda^{2}}{2 n}+2 n\left(\digamma_{1}+\digamma_{3}\right)^{2}+L_{6}\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right]\right\} \times g\left(\eta\left(Y_{1}\right) Y_{2}-\eta\left(Y_{2}\right) Y_{1}, Y_{4}\right)=0 \tag{3.49}
\end{equation*}
$$

It is clear from (3.49),

$$
L_{6}=-\frac{\lambda^{2}+4 n^{2}\left(\digamma_{1}+\digamma_{3}\right)^{2}}{2 n\left[\lambda+2 n\left(\digamma_{1}+\digamma_{3}\right)\right]} .
$$

This completes the proof.

## 4. Conclusion

In this paper, we consider Lorentz generalized Sasakian space forms admitting almost $\eta$-Ricci solitons in some curvature tensors. Ricci pseudosymmetry concepts of Lorentz generalized Sasakian space forms admitting $\eta$-Ricci soliton have introduced according to the choice of some special curvature tensors such as Riemann, concircular, projective, $\mathscr{M}$ - projective, $W_{1}$ and $W_{2}$. Then, again according to the choice of the curvature tensor, necessary conditions are given for Lorentz generalized Sasakian space form admitting $\eta$-Ricci soliton to be Ricci semisymmetric. Then some characterizations are obtained and some classifications have made.

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# Analytical Rational Solitons of the Modified Lakshmanan-Porsezian-Daniel Equation 

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#### Abstract

In this paper, the Lakshmanan-Porsezian-Daniel (LPD) equation is studied. New analytical rational solitons to the LPD equation are presented by an ansatz method. Wave solutions of three types, such as parabolic, trigonometric and hyperbolic function solutions have been retrieved. All solutions are plotted in 3D to enhance the understanding of their physical characteristics. These simulations, which represent the behavior of the resulting hyperbolic, parabolic and trigonometric solitons, are provided by choosing different appropriate values of the parameters.


## 1. Introduction

Many researchers in fields such as mathematics, physics, engineering and more are very interested in nonlinear partial equations, since most physical systems are not linear in nature. The Lakshmanan-Porsezian-Daniel (LPD) equation is a widely known nonlinear partial differential equation. It is a generalization of the nonlinear Schrödinger equation that includes higher order nonlinear and dispersed terms. In recent years, it has attracted great attention from mathematicians and physicists. The LPD equation has also been generalized and extended in many ways, including the addition of external forcing, the inclusion of damping effects, and the consideration of higher-dimensional versions of the equation. The LPD equation and its variants have been used to model a variety of physical systems in many areas of physics and engineering, including plasma physics, fluid dynamics, and nonlinear optics, and has been studied extensively from different perspectives such as integrability, symmetry analysis, solution methods and applications: Ricatti equation [1], $\tan \left(\frac{\psi(\eta)}{2}\right)$-expansion technique [2], collective variable [3], modified simple equation method [4], method of undetermined coefficients [5], Darboux transformations [6,7], Rogue wave equation [8], the modified auxiliary equation method [9] etc.

This paper investigates the Lakshmanan-Porsezian-Daniel equation [10, 11], a well known partial differential equation that describes the pulse propagation in an optical fiber which is in the form

$$
\begin{equation*}
i q_{t}+a q_{x x}+b q_{x t}+\zeta\left|q^{2}\right| q=\sigma q_{x x x x}+\beta\left|q_{x}\right|^{2} q+\gamma|q|^{2} q_{x x}+\delta|q|^{4} q \tag{1.1}
\end{equation*}
$$

where the complex valued function $q(x, t)$ depends on space $x$ and time $t$. The term $i q_{t}$ denotes the temporal evolution of pulse. The group velocity dispersion and spatio-temporal dispersion are given by $a$ and $b$, respectively. The fourth-order dispersion and two-photon absorption are represented by constants $\sigma$ and $\delta$, respectively. The parameters $\beta$ and $\gamma$ indicate the non-linear forms of dispersion.

In this paper, our aim is finding solutions in the form of parabolic, trigonometric and hyperbolic solitons of the LPD equation. First we start by using traveling wave variables to find a solution for Eq. (1.1). After analyzing the resulting system of equations to find the condition of its compatibility, it turned out that the structure for the system of equations. At the second stage, a special logarithmic transformation is applied. At the last stage, three different methods are applied to retrieved equation. The solutions obtained by appropriate selection of some parameters affecting the shape and velocity of the solitons are observed with three-dimensional plots.

[^0]

## 2. System of Differential Equations Corresponding to Eq. (1.1)

In order to find soliton solutions of Eq. (1.1), we use the traveling wave reduction in the form

$$
\begin{equation*}
q(x, t)=y(z) e^{i \theta}, \quad z=x-c t, \quad \theta=k x-w t \tag{2.1}
\end{equation*}
$$

where $y(z)$ is a complex-valued function that represents the structure of the pulse, $k, w, c, \theta$ are parameters of solution. The phase component of the soliton is $\theta, k$ is represents the frequency of the soliton, $c$ is the velocity of the soliton, while $w$ is the wave number.
Substituting solution (2.1) into Eq. (1.1) and equating the real and imaginary parts of expression to zero, respectively the following equations are obtained

$$
\begin{equation*}
\sigma y_{z z z z}+\left(-a+b c-6 \sigma k^{2}\right) y_{z z}+\gamma y^{2} y_{z z}+\beta y y_{z}^{2}+\left(\beta k^{2}-\gamma k^{2}-\zeta\right) y^{3}+\delta y^{5}+\left(\sigma k^{4}-w+a k^{2}-b w k\right) y=0 \tag{2.2}
\end{equation*}
$$

and

$$
\begin{equation*}
4 \sigma k y_{z z z}+\left(-4 \sigma k^{3}+c-2 a k+b(c k+w)\right) y_{z}+2 \gamma k y^{2} y_{z}=0 \tag{2.3}
\end{equation*}
$$

The $y(z)$ function must satify both of the above third and fourth order differential equations obtained, respectively (2.2) and (2.3).
For the purpose of evaluating the solution of the Eq. (1.1), we implement the following logaritmic transformation from [12] on $y$,

$$
\begin{equation*}
y=2(\ln f)_{z} \tag{2.4}
\end{equation*}
$$

If transformation (2.4) is substituted into in (2.2) and (2.3), then the resulting expressions obtained as:

$$
\begin{gather*}
\left(\begin{array}{c}
-4 \zeta f^{2} f_{z}^{3}-60 \sigma f f_{z z} f_{z}^{3}-12 \gamma f f_{z z} f_{z}^{3}-8 \beta f f_{z z} f_{z}^{3}+4 \gamma f^{2} f_{z}^{2} f_{z z z}+20 \sigma f^{2} f_{z}^{2} f_{z z z}+30 \sigma f^{2} f_{z z z}^{2} f_{z}+4 \beta f^{2} f_{z z}^{2} f_{z} \\
-10 \sigma f^{3} f_{z z} f_{z z z}+3 a f^{3} f_{z z} f_{z}+b c f^{4} f_{z z z}-5 \sigma f^{3} f_{z} f_{z z z z}-6 \sigma k^{2} f^{4} f_{z z z}+8 \gamma f_{z}^{5}+4 \beta f_{z}^{5}+16 \delta f_{z}^{5}+24 \sigma f_{z}^{5}-w f^{4} f_{z} \\
-2 a f^{2} f_{z}^{3}-b k w f^{4} f_{z}+k^{4} \sigma f^{4} f_{z}+a k^{2} f^{4} f_{z}-4 \gamma k^{2} f^{2} f_{z}^{3}+4 \beta k^{2} f^{2} f_{z}^{3}-12 \sigma k^{2} f^{2} f_{z}^{3}+2 b c f^{2} f_{z}^{3} \\
-a f^{4} f_{z z z}+\sigma f_{z z z z z} f^{4}-3 b c f^{3} f_{z z} f_{z}+18 \sigma k^{2} f^{3} f_{z z} f_{z} \\
-8 \gamma k f_{z}^{4}+4 \sigma f_{z z z} f^{3}+c f^{3} f_{z z}-16 \sigma f^{2} f_{z z z} f_{z}-c f^{2} f_{z}^{2}-12 \sigma f^{2} f_{z z}^{2}+48 \sigma f f_{z}^{2} f_{z z}-24 \sigma f_{z}^{4}
\end{array}\right)=0,  \tag{2.5}\\
\binom{-4 \sigma k^{3} f^{3} f_{z z}+4 \sigma f^{3} f^{2} f_{z}^{2}+b c k f^{3} f_{z z}-b x k f^{2} f_{z}^{2}-2 a k f^{3} f_{z z}+b w f^{3} f_{z z}+2 a k f_{z}^{2}-b w f_{z}^{2}+8 \gamma k f f_{z}}{-8}=0
\end{gather*}
$$

Now, we use this form to evaluate various rational solitons.

## 3. Hyperbolic Solitons of Differential Equations Corresponding to Eq. (1.1)

To get hyperbolic solitons of (2.5), we use the following transformation:

$$
\begin{equation*}
f=b_{0} \cosh z+b_{1} \sinh z \tag{3.1}
\end{equation*}
$$

where $b_{0}, b_{1}$ are any constants to be determined. Substituting (3.1) into (2.5) and equating the coefficient terms that are containing independent combinations of cosh and sinh functions to zero, we obtain a system of algebraic equations:

$$
\left(b_{0}-b_{1}\right)\left(b_{0}+b_{1}\right)\left(b_{0}^{2}+b_{1}^{2}\right)\left(-4 k^{3} \sigma+b c k-2 a k+b w+8 \gamma k+c+16 \sigma\right)=0
$$

$$
2 b_{0} b_{1}\left(b_{0}-b_{1}\right)\left(b_{0}+b_{1}\right)\left(-4 k^{3} \sigma+b c k-2 a k+b w+8 \gamma k+c+16 \sigma\right)=0
$$

$$
-\left(b_{0}-b_{1}\right)\left(b_{0}+b_{1}\right)\left(-4 b_{1}^{2} k^{3} \sigma+b b_{1}^{2} c k-2 a b_{1}^{2} k+b b_{1}^{2} w+8 b_{0}^{2} \gamma k+24 b_{0}^{2} \sigma+b_{1}^{2} c-8 b_{1}^{2} \sigma\right)=0
$$

$$
\begin{align*}
& -b_{1}\left(5 b_{0}^{4}+10 b_{0}^{2} b_{1}^{2}+b_{1}^{4}\right)\left(-k^{4} \sigma-a k^{2}+b k w-4 \beta k^{2}+4 \gamma k^{2}+4 \zeta-16 \delta+w\right)=0, \\
& -b_{0}\left(b_{0}^{4}+10 b_{0}^{2} b_{1}^{2}+5 b_{1}^{4}\right)\left(-k^{4} \sigma-a k^{2}+b k w-4 \beta k^{2}+4 \gamma k^{2}+4 \zeta-16 \delta+w\right)=0, \\
& 2 b_{1}\left(\begin{array}{c}
\left(-2 k^{4} \sigma-2 a k^{2}+2 b k w-14 \beta k^{2}+14 \gamma k^{2}+18 k^{2} \sigma-3 b c+14 \zeta+3 a-80 \delta-12 \gamma-12 \sigma+2 w\right) b_{0}^{4} \\
+\left(-7 k^{4} \sigma-7 a k^{2}+7 b k w-24 \beta k^{2}+24 \gamma k^{2}-12 k^{2} \sigma+2 b c+24 \zeta-2 a-80 \delta+8 \gamma+8 \sigma+7 w\right) b_{1}^{2} b_{0}^{2} \\
+\left(-k^{4} \sigma-a k^{2}+b k w-2 \beta k^{2}+2 \gamma k^{2}-6 k^{2} \sigma+b c+2 \zeta-a+4 \gamma+4 \sigma+w\right) b_{1}^{4}
\end{array}\right)=0, \\
& 2 b_{0}\left(\begin{array}{c}
\left(-2 \beta k^{2}+2 \gamma k^{2}+6 k^{2} \sigma-b c+2 \zeta+a-16 \delta-4 \gamma-4 \sigma\right) b_{0}^{4} \\
+\left(-3 k^{4} \sigma-3 a k^{2}+3 b k w-16 \beta k^{2}+16 \gamma k^{2}+12 k^{2} \sigma-2 b c+16 \zeta+2 a-80 \delta-8 \gamma-8 \sigma+3 w\right) b_{0}^{2} b_{1}^{2} \\
+\left(-3 k^{4} \sigma-3 a k^{2}+3 b k w-6 \beta k^{2}+6 \gamma k^{2}-18 k^{2} \sigma+3 b c+6 \zeta-3 a+12 \gamma+12 \sigma+3 w\right) b_{1}^{4}
\end{array}\right)=0, \\
& -b_{1}\left(\begin{array}{c}
\left(-8 \beta k^{2}+8 \gamma k^{2}+24 k^{2} \sigma-4 b c+8 \zeta+4 a-4 \beta-80 \delta-24 \gamma-40 \sigma\right) b_{0}^{4} \\
+\left(-4 k^{4} \sigma-4 a k^{2}+4 b k w-12 \beta k^{2}+12 \gamma k^{2}-12 k^{2} \sigma+2 b c+12 \zeta-2 a+8 \beta+24 \gamma+56 \sigma+4 w\right) b_{0}^{2} b_{1}^{2} \\
+\left(-k^{4} \sigma-a k^{2}+b k w-12 k^{2} \sigma+2 b c-2 a-4 \beta-16 \sigma+w\right) b_{1}^{4}
\end{array}\right)=0, \\
& b_{0}\left(\begin{array}{c}
(8 \gamma+4 \beta+16 \delta+24 \sigma) b_{0}^{4} \\
+\left(4 \beta k^{2}-4 \gamma k^{2}-12 k^{2} \sigma+2 b c-4 \zeta-2 a-8 \beta-8 \gamma-40 \sigma\right) b_{0}^{2} b_{1}^{2} \\
+\left(k^{4} \sigma+a k^{2}-b k w+12 k^{2} \sigma-2 b c+2 a+4 \beta+16 \sigma-w\right) b_{1}^{4}
\end{array}\right)=0, \tag{3.2}
\end{align*}
$$

After solving the system (3.2) with the help of Maple software, three cases of parametric values are obtained as follows:
Case 1:

$$
\begin{equation*}
b_{1}=b_{0}, \quad \sigma=\frac{-a k^{2}+b k w-4 \beta k^{2}+4 \gamma k^{2}+4 \zeta-16 \delta+w}{k^{4}} \tag{3.3}
\end{equation*}
$$

Via the parametric values in (3.3), we have

$$
\begin{equation*}
f=b_{0}(\cosh z+\sinh z) \tag{3.4}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=2 \tag{3.5}
\end{equation*}
$$

By using Eq. (3.5) into Eq. (2.1), we obtain a first type of rational hyperbolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=2 e^{i(k x-w t)} \tag{3.6}
\end{equation*}
$$

(See Figure 3.1)


Figure 3.1: 3D plots of the rational solution (3.6) in Case 1 with the values of $k=2, w=1, c=1$, (a) Real, (b) Imaginary and (c) Complex.

Case 2:

$$
\zeta=\frac{-\left(\begin{array}{c}
b c k^{8}+2 a k^{8}-3 b k^{7} w+2 b c k^{6}+c k^{7}-20 a k^{6}+18 b k^{5} w-4 k^{6} w+8 a k^{5}+12 b c k^{4} \\
-8 b k^{4} w+10 c k^{5}-16 a k^{4}-28 b c k^{3}+4 b k^{3} w+8 k^{4} w+24 a k^{3}+8 b c k^{2}+4 b k^{2} w+4 c k^{3} \\
-8 k^{3} w-16 a k^{2}-8 b c k+8 b k w-12 c k^{2}+32 a k-24 b w+8 c k+16 k w-24 c
\end{array}\right)}{16 k\left(k^{3}+2\right)}
$$

$$
\delta=\frac{\begin{array}{c}
(b c+2 a) k^{6}+(-3 b w+c) k^{5}+4(b c-4 a-w) k^{4}+4(3 b w+2 a+3 c) k^{3}  \tag{3.7}\\
+8(-(c+w) b+2 a) k^{2}+8((c-w) b-4 a-c-w) k+24(b w+c)
\end{array}}{64 k\left(k^{3}+2\right)}
$$

$$
\gamma=\frac{-3((b c-2 a) k+b w+c)}{4 k\left(k^{3}+2\right)}
$$

$$
\beta=\frac{\begin{array}{c}
-(b c+2 a) k^{5}+(3 b w-c) k^{4}+4(-b c+4 a+w) k^{3}-4(3 b w+2 a+3 c) k^{2} \\
+8((-2 c+w) b+4 a) k+8(2(c-w) b-2(a+c)+w)
\end{array}}{16 k^{3}+32}
$$

$$
\sigma=\frac{(b c-2 a) k+b w+c}{4 k^{3}+8}
$$

By using values in (3.7) into (3.1), we have

$$
\begin{equation*}
f=b_{0} \cosh z+b_{1} \sinh z \tag{3.8}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{2\left(b_{0} \sinh (z)+b_{1} \cosh (z)\right)}{b_{0} \cosh (z)+b_{1} \sinh (z)} \tag{3.9}
\end{equation*}
$$

By using Eq. (3.9) into Eq. (2.1), we obtain a second type of rational hyperbolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{2\left(b_{0} \sinh (z)+b_{1} \cosh (z)\right) e^{i(k x-w t)}}{b_{0} \cosh (z)+b_{1} \sinh (z)} \tag{3.10}
\end{equation*}
$$

(See Figure 3.2)


Figure 3.2: 3D plots of the rational solution (3.10) in Case 2 with the values of $k=2, w=1, c=0.1, b_{0}=-5, b_{1}=0.1$, (a) Real, (b) Imaginary and (c) Complex.

## 4. Parabolic Solutions of Differential Equations Corresponding to Eq. (1.1)

To get parabolic solution of (2.5), we choose $f$ as following:

$$
\begin{equation*}
f=b_{2} z^{2}+b_{1} z+b_{0} \tag{4.1}
\end{equation*}
$$

where $b_{0}, b_{1}$ and $b_{2}$ represent any constant parameters. By substituting (4.1) into (2.5) and equating the various coefficient terms of $z$, we then solve the following system of algebraic equations to find the values of the parameters:

$$
\left.\begin{array}{l}
2 b_{2}^{4}\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right)=0 \\
6 b_{1} b_{2}^{3}\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right)=0, \\
b_{2}^{2}\left(\left(8 k^{3} \sigma-2 b c k+4 a k-2 b w-2 c\right) b_{2} b 0+\left(28 k^{3} \sigma-7 b c k+14 a k-7 b w-7 c\right) b_{1}^{2}+(-64 \gamma k-48 \sigma) b_{2}^{2}\right)=0, \\
4 b_{1} b_{2}\left(\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right) b_{2} b 0+\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right) b_{1}^{2}+(-32 \gamma k-24 \sigma) b_{2}^{2}\right)=0, \\
{\left[\begin{array}{r}
+\left(-8 k^{3} \sigma+2 b c k-4 a k+2 b w+2 c\right) b_{2}^{2} b 0^{2}+\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right) b_{1}^{4}+(-112 \gamma k-144 \sigma) b_{2}^{2} b_{1}^{2}
\end{array}\right]=0,} \\
\quad\left(\left(16 k^{3} \sigma-4 b c k+8 a k-4 b w-4 c\right) b_{2} b_{1}^{2}+(64 \gamma k+288 \sigma) b_{2}^{3}\right) b_{0}, \\
-2 b_{1}\left(\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right) b_{2} b 0^{2}+\left(\left(-4 k^{3} \sigma+b c k-2 a k+b w+c\right) b_{1}^{2}+(-32 \gamma k-144 \sigma) b_{2}^{2}\right) b 0+(24 \gamma k+48 \sigma) b_{2} b_{1}^{2}\right)=0,  \tag{4.2}\\
{\left[\begin{array}{r}
\left(-8 k^{3} \sigma+2 b c k-4 a k+2 b w+2 c\right) b_{2} b_{0}^{3}+(16 \gamma k+96 \sigma) b_{2} b_{1}^{2} b 0+(-8 \gamma k-24 \sigma) b_{1}^{4}
\end{array}\right]=0,} \\
\quad\left(4 k^{3} \sigma-b c k+2 a k-b w-c\right) b_{1}^{2}-48 b_{2}^{2} \sigma
\end{array}\right] \begin{aligned}
& (4.2)
\end{aligned}
$$

$$
\begin{aligned}
& 2 b_{2}^{5}\left(k^{4} \sigma+a k^{2}-b k w-w\right)=0, \\
& 9 b_{1} b_{2}^{4}\left(k^{4} \sigma+a k^{2}-b k w-w\right)=0, \\
& -4 b_{2}^{3}\left(\left(-2 k^{4} \sigma-2 a k^{2}+2 b k w+2 w\right) b_{2} b_{0}+\left(-4 k^{4} \sigma-4 a k^{2}+4 b k w+4 w\right) b_{1}^{2}+\left(-8 \beta k^{2}+8 \gamma k^{2}+6 k^{2} \sigma-b c+8 \zeta+a\right) b_{2}^{2}\right)=0, \\
& -14 b_{1} b_{2}^{2}\left(\left(-2 k^{4} \sigma-2 a k^{2}+2 b k w+2 w\right) b_{2} b_{0}+\left(-k^{4} \sigma-a k^{2}+b k w+w\right) b_{1}^{2}+\left(-8 \beta k^{2}+8 \gamma k^{2}+6 k^{2} \sigma-b c+8 \zeta+a\right) b_{2}^{2}\right)=0, \\
& -2 b_{2}\left[\begin{array}{c}
\left(-18 k^{4} \sigma-18 a k^{2}+18 b k w+18 w\right) b_{2} b_{1}^{2}+\left(-32 \beta k^{2}+32 \gamma k^{2}-12 k^{2} \sigma+2 b c+32 \zeta-2 a\right) b_{2}^{3} \\
\left(-6 k^{4} \sigma-6 a k^{2}+6 b k w+6 w\right) b_{2}^{2} b_{0}^{2}+\left(-3 k^{4} \sigma-3 a k^{2}+3 b k w+3 w\right) b_{1}^{4} \\
+\left(-76 \beta k^{2}+76 \gamma k^{2}+66 k^{2} \sigma-11 b c+76 \zeta+11 a\right) b_{2}^{2} b_{1}^{2}+(-16 \beta-256 \delta-32 \gamma-24 \sigma) b_{2}^{4}
\end{array}\right]=0, \\
& -b_{1}\left[\begin{array}{c}
\left(\left(-20 k^{4} \sigma-20 a k^{2}+20 b k w+20 w\right) b_{2} b_{1}^{2}+\left(-160 \beta k^{2}+160 \gamma k^{2}-60 k^{2} \sigma+10 b c+160 \zeta-10 a\right) b_{2}^{3}\right) b_{0} \\
+\left(-30 k^{4} \sigma-30 a k^{2}+30 b k w+30 w\right) b_{2}^{2} b_{0}^{2}+\left(-k^{4} \sigma-a k^{2}+b k w+w\right) b_{1}^{4} \\
\left(-100 \beta k^{2}+100 \gamma k^{2}+120 k^{2} \sigma-20 b c+100 \zeta+20 a\right) b_{2}^{2} b_{1}^{2}+(-80 \beta-1280 \delta-160 \gamma-120 \sigma) b_{2}^{4}
\end{array}\right]=0, \\
& {\left[\begin{array}{c}
\left(\left(24 k^{4} \sigma+24 a k^{2}-24 b k w-24 w\right) b_{2} b_{1}^{2}+\left(32 \beta k^{2}-32 \gamma k^{2}+120 k^{2} \sigma-20 b c-32 \zeta+20 a\right) b_{2}^{3}\right) b_{0}^{2} \\
+\left(8 k^{4} \sigma+8 a k^{2}-8 b k w-8 w\right) b_{2}^{2} b_{0}^{3}+\left(32 \beta k^{2}-32 \gamma k^{2}-60 k^{2} \sigma+10 b c-32 \zeta-10 a\right) b_{2} b_{1}^{4} \\
+(96 \beta+1280 \delta+208 \gamma+240 \sigma) b_{2}^{3} b_{1}^{2} \\
+\left(\left(4 k^{4} \sigma+4 a k^{2}-4 b k w-4 w\right) b_{1}^{4}+\left(144 \beta k^{2}-144 \gamma k^{2}-144 \zeta\right) b_{2}^{2} b_{1}^{2}+(-64 \beta-192 \gamma-480 \sigma) b_{2}^{4}\right) b_{0}
\end{array}\right]=0,} \\
& -2 b_{1}\left[\begin{array}{c}
\left(\left(-3 k^{4} \sigma-3 a k^{2}+3 b k w+3 w\right) b_{1}^{2}+\left(-24 \beta k^{2}+24 \gamma k^{2}-90 k^{2} \sigma+15 b c+24 \zeta-15 a\right) b_{2}^{2}\right) b_{0}^{2} \\
+\left(-6 k^{4} \sigma-6 a k^{2}+6 b k w+6 w\right) b_{2} b_{0}^{3}+\left(-2 \beta k^{2}+2 \gamma k^{2}+6 k^{2} \sigma-b c+2 \zeta+a\right) b_{1}^{4} \\
+(-32 \beta-320 \delta-76 g-120 \sigma) b_{2}^{2} b_{1}^{2} \\
+\left(\left(-28 \beta k^{2}+28 \gamma k^{2}+30 k^{2} \sigma-5 b c+28 \zeta+5 a\right) b_{2} b_{1}^{2}+(48 \beta+144 \gamma+360 \sigma) b_{2}^{3}\right) b_{0}
\end{array}\right]=0, \\
& {\left[\begin{array}{c}
\left(2 k^{4} \sigma+2 a k^{2}-2 b k w-2 w\right) b_{2} b_{0}^{4}+\left(\left(4 k^{4} \sigma+4 a k^{2}-4 b k w-4 w\right) b_{1}^{2}+\left(72 k^{2} \sigma-12 b c+12 a\right) b_{2}^{2}\right) b_{0}^{3} \\
+\left(\left(24 \beta k^{2}-24 \gamma k^{2}+36 k^{2} \sigma-6 b c-24 \zeta+6 a\right) b_{2} b_{1}^{2}+(32 \beta+240 \sigma) b_{2}^{3}\right) b_{0}^{2} \\
+\left(\left(8 \beta k^{2}-8 \gamma k^{2}-24 k^{2} \sigma+4 b c-8 \zeta-4 a\right) b_{1}^{4}+(-64 \beta-144 \gamma-480 \sigma) b_{2}^{2} b_{1}^{2}\right) b_{0}+(24 \beta+160 \delta+56 \gamma+120 \sigma) b_{2} b_{1}^{4}
\end{array}\right]=0,} \\
& -b_{1}\left[\begin{array}{c}
\left(-k^{4} \sigma-a k^{2}+b k w+w\right) b_{0}^{4}+\left(-36 k^{2} \sigma+6 b c-6 a\right) b_{2} b_{0}^{3}+(-4 \beta-16 \delta-8 \gamma-24 \sigma) b_{1}^{4} \\
+\left(\left(-4 \beta k^{2}+4 \gamma k^{2}+12 k^{2} \sigma-2 b c+4 \zeta+2 a\right) b_{1}^{2}+(-16 \beta-120 \sigma) b_{2}^{2}\right) b_{0}^{2}+(16 \beta+24 \gamma+120 \sigma) b_{2} b_{1}^{2} b_{0}
\end{array}\right]=0 .
\end{aligned}
$$

Resolution of the system (4.2) with the help of Maple gives five cases of parametric values as follows:
Case 1:

$$
\begin{align*}
& a=\frac{4 \gamma k^{5}+3 b k w+3 w}{3 k^{2}}, \quad b_{0}=b_{1}=0, \quad \beta=-\frac{12 b \gamma k^{6}-24 b \gamma k^{5}+20 \gamma k^{5}-24 \zeta b k^{3}-24 \gamma k^{4}-24 \zeta k^{2}-3 w}{24 k^{4}(b k+1)}, \\
& c=-\frac{8 \gamma k^{5}-3 b k w-6 w}{3 k(b k+1)}, \quad \delta=\frac{60 b \gamma k^{6}-72 b \gamma k^{5}+68 \gamma k^{5}-24 \zeta b k^{3}-72 \gamma k^{4}-24 \zeta k^{2}-3 w}{384 k^{4}(b k+1)}, \quad \sigma=-\frac{4 \gamma k}{3} . \tag{4.3}
\end{align*}
$$

By using values in (4.3) into (4.1), we have

$$
\begin{equation*}
f=b_{2} z^{2} . \tag{4.4}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{4}{z} . \tag{4.5}
\end{equation*}
$$

By replacing Eq. (2.1) with Eq. (4.5), we obtain the first type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{4 e^{i(k x-w t)}}{-c t+x} . \tag{4.6}
\end{equation*}
$$

(See Figure 4.1)

(a)

(b)

(c)

Figure 4.1: 3D plots of the rational solution (4.6) in Case 1 with the values of $k=2, w=1, b=128, \gamma=-3, b_{1}=0.1$, (a) Real, (b) Imaginary and (c) Complex.

Case 2:

$$
\begin{align*}
& a=\frac{4 \gamma k^{3}}{3}, \quad b=-\frac{1}{k}, \quad b_{0}=b_{1}=0, \quad \beta=\frac{-20 \gamma k^{4}+24 \gamma k^{3}+24 \zeta k+3 c}{24 k^{3}} \\
& \delta=-\frac{-68 \gamma k^{4}+72 \gamma k^{3}+24 \zeta k+3 c}{384 k^{3}}, \quad \sigma=-\frac{4 \gamma k}{3}, \quad w=\frac{8 \gamma k^{5}}{3} \tag{4.7}
\end{align*}
$$

By using values in (4.7) into (4.1), we have

$$
\begin{equation*}
f=b_{2} z^{2} \tag{4.8}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{4}{z} \tag{4.9}
\end{equation*}
$$

By using Eq. (4.9) into Eq. (2.1), we obtain a second type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{4 e^{i\left(-\frac{8}{3} \gamma k^{5} t+k x\right)}}{-c t+x} \tag{4.10}
\end{equation*}
$$

(See Figure 4.2)


Figure 4.2: 3D plots of the rational solution (4.10) in Case 2 with the values of $k=2, c=1, \gamma=3$, (a) Real, (b) Imaginary and (c) Complex.
Case 3:

$$
\begin{align*}
& a=\frac{\gamma k^{5}+3 b k w+3 w}{3 k^{2}}, \quad b_{2}=0, \quad \beta=\frac{-3 b \gamma k^{6}+6 b \gamma k^{5}-5 \gamma k^{5}+6 \zeta b k^{3}+6 \gamma k^{4}+6 \zeta k^{2}+3 w}{6 k^{4}(b k+1)}, \\
& c=\frac{-2 \gamma k^{5}+3 b k w+6 w}{3 k(b k+1)}, \quad \delta=-\frac{-15 b \gamma k^{6}+18 b \gamma k^{5}-17 \gamma k^{5}+6 \zeta b k^{3}+18 \gamma k^{4}+6 \zeta k^{2}+3 w}{24 k^{4}(b k+1)}, \quad \sigma=-\frac{\gamma k}{3} . \tag{4.11}
\end{align*}
$$

By using values in (4.11) into (4.1), we have

$$
\begin{equation*}
f=b_{1} z+b_{0} \tag{4.12}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{2 b_{1}}{b_{1} z+b_{0}} \tag{4.13}
\end{equation*}
$$

By replacing Eq. (2.1) with Eq. (4.13), we obtain a third type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{2 b_{1} e^{i(k x-w t)}}{b_{1}\left(-\frac{\left(-2 \gamma k^{5}+3 b k w+6 w\right) t}{3 k(b k+1)}+x\right)+b_{0}} \tag{4.14}
\end{equation*}
$$

(See Figure 4.3)


Figure 4.3: 3D plots of the rational solution (4.14) in Case 3 with the values of $b_{1}=1, b_{0}=1, k=2, w=1, b=\frac{-5}{12}, \gamma=\frac{1}{64}$, (a) Real, (b) Imaginary and (c) Complex.

Case 4:

$$
\begin{align*}
& a=\frac{k^{3} \gamma}{3}, \quad b=-\frac{1}{k}, \quad b_{2}=0, \quad \beta=\frac{-5 \gamma k^{4}+6 \gamma k^{3}+6 \zeta k+3 c}{6 k^{3}} \\
& \delta=-\frac{-17 \gamma k^{4}+18 \gamma k^{3}+6 \zeta k+3 c}{24 k^{3}}, \quad \sigma=-\frac{k \gamma}{3}, \quad w=\frac{2 k^{5} \gamma}{3} \tag{4.15}
\end{align*}
$$

By using values in (4.15) into (4.1), we have

$$
\begin{equation*}
f=b_{1} z+b_{0} \tag{4.16}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{2 b_{1}}{b_{1} z+b_{0}} \tag{4.17}
\end{equation*}
$$

By using Eq. (4.17) into Eq. (2.1), we obtain a fourth type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{2 b_{1} e^{i\left(\frac{-2}{3} \gamma k^{5} t+k x\right)}}{b_{1}(-c t+x)+b_{0}} \tag{4.18}
\end{equation*}
$$

(See Figure 4.4)
Case 5:

$$
\begin{align*}
& a=\frac{4 \gamma k^{5}+3 b k w+3 w}{3 k^{2}}, \quad b_{2}=\frac{b_{1}^{2}}{4 b_{0}}, \quad \beta=\frac{-12 b \gamma k^{6}+24 b \gamma k^{5}-20 \gamma k^{5}+24 \zeta b k^{3}+24 \gamma k^{4}+24 \zeta k^{2}+3 w}{24 k^{4}(b k+1)} \\
& c=\frac{-8 \gamma k^{5}+3 b k w+6 w}{3 k(b k+1)}, \quad \delta=-\frac{-60 b \gamma k^{6}+72 b \gamma k^{5}-68 \gamma k^{5}+24 \zeta b k^{3}+72 \gamma k^{4}+24 \zeta k^{2}+3 w}{384 k^{4}(b k+1)}, \quad \sigma=-\frac{4 \gamma k}{3} \tag{4.19}
\end{align*}
$$



Figure 4.4: 3D plots of the rational solution (4.18) in Case 4 with the values of $b_{1}=1, b_{0}=1, k=2, w=1, b=\frac{-5}{12}, \gamma=\frac{1}{64}, c=1$, (a) Real, (b) Imaginary and (c) Complex.

By using values in (4.19) into (4.1), we have

$$
\begin{equation*}
f=\frac{b_{1}^{2}}{4 b_{0}} z^{2}+b_{1} z+b_{0} \tag{4.20}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{4 b_{1}}{b_{1} z+2 b_{0}} \tag{4.21}
\end{equation*}
$$

By using Eq. (4.21) into Eq. (2.1), we obtain a fifth type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{4 b_{1} e^{i(k x-w t)}}{b_{1}\left(-\frac{\left(-8 \gamma k^{5}+3 b k w+6 w\right) t}{3 k(b k+1)}+x\right)+2 b_{0}} . \tag{4.22}
\end{equation*}
$$

(See Figure 4.5)


Figure 4.5: 3D plots of the rational solution (4.22) in Case 5 with the values of $b_{1}=2, b_{0}=-1, k=2, w=1, b=\frac{-1}{3}, \gamma=0.1$, (a) Real, (b) Imaginary and (c) Complex.

Case 6:

$$
\begin{align*}
& a=\frac{4 k^{3} \gamma}{3}, \quad b=\frac{-1}{k}, \quad b_{2}=\frac{b_{1}^{2}}{4 b_{0}}, \quad \beta=\frac{-20 \gamma k^{4}+24 \gamma k^{3}+24 \zeta k+3 c}{24 k^{3}} \\
& \delta=-\frac{-68 \gamma k^{4}+72 \gamma k^{3}+24 \zeta k+3 c}{384 k^{3}}, \quad \sigma=-\frac{4 \gamma k}{3}, \quad w=\frac{8 k^{5} \gamma}{3} \tag{4.23}
\end{align*}
$$

By using values in (4.23) into (4.1), we have

$$
\begin{equation*}
f=\frac{b_{1}^{2}}{4 b_{0}} z^{2}+b_{1} z+b_{0} \tag{4.24}
\end{equation*}
$$

By using $y=2(\ln f)_{z}$, we have

$$
\begin{equation*}
y=\frac{4 b_{1}}{b_{1} z+2 b_{0}} \tag{4.25}
\end{equation*}
$$

By using Eq. (4.25) into Eq. (2.1), we obtain a sixth type of rational parabolic solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=\frac{4 b_{1} e^{i\left(k x-\frac{8 k^{5} \gamma}{3} t\right)}}{b_{1}(x-c t)+2 b_{0}} \tag{4.26}
\end{equation*}
$$

(See Figure 4.6)


Figure 4.6: 3D plots of the rational solution (4.26) in Case 5 with the values of $b_{1}=2, b_{0}=-1, k=2, c=1, \gamma=3$, (a) Real, (b) Imaginary and (c) Complex.

## 5. Trigonometric Solutions of Differential Equations Corresponding to Eq. (1.1)

To get trigonometric solution, we suppose a solution of (2.5) in the following form:

$$
\begin{equation*}
f=b_{0} \cos z+b_{1} \sin z \tag{5.1}
\end{equation*}
$$

where $b_{0}, b_{1}$ are all constants. Inserting (5.1) into (2.5) and setting the coefficient terms that are containing independent combinations of cos and sin functions to zero, we get a system of algebraic equations:

$$
\begin{align*}
& 4 b_{1}^{4}\left(4 k^{3} \sigma-b c k+2 a k-b w+8 \gamma k-c+16 \sigma\right)=0 \\
& 2 b_{1}^{4}\left(4 k^{3} \sigma-b c k+2 a k-b w-8 \gamma k-c-32 \sigma\right)=0 \\
& 4 b_{1}^{5}\left(k^{4} \sigma+a k^{2}-b k w-4 \beta k^{2}+4 \gamma k^{2}+4 \zeta+16 \delta-w\right)=0 \\
& 8 b_{1}^{5}\left(k^{4} \sigma+a k^{2}-b k w-2 \beta k^{2}+2 \gamma k^{2}-6 k^{2} \sigma+b c+2 \zeta-a-4 \gamma-4 \sigma-w\right)=0  \tag{5.2}\\
& -8 b_{1}^{5}\left(-2 \beta k^{2}+2 \gamma k^{2}+6 k^{2} \sigma-b c+2 \zeta+a+16 \delta+4 \gamma+4 \sigma\right)=0 \\
& -b_{1}^{5}\left(3 k^{4} \sigma+3 a k^{2}-3 b k w-4 \beta k^{2}+4 \gamma k^{2}-24 k^{2} \sigma+4 b c+4 \zeta-4 a-16 \beta-80 \delta-48 \gamma-112 \sigma-3 w\right)=0 \\
& b_{1}^{5}\left(-k^{4} \sigma-a k^{2}+b k w-4 \beta k^{2}+4 \gamma k^{2}+24 k^{2} \sigma-4 b c+4 \zeta+4 a-16 \beta-16 \delta-16 \gamma-80 \sigma+w\right)=0 .
\end{align*}
$$

After solving the system (5.2) with the help of Maple, we obtain three cases of parametric values as follows:

Case 1:

$$
\begin{align*}
& a=\frac{\left(\begin{array}{c}
3 b \gamma k^{7}+3 b^{2} c k^{4}+18 b \gamma k^{5}+4 \gamma k^{6}-8 b \gamma k^{4}+12 b^{2} c k^{2} \\
+6 b c k^{3}-4 b \gamma k^{3}+8 \gamma k^{4}-4 b \gamma k^{2}-8 \gamma k^{3}+6 b^{2} c \\
+12 b c k+8 b \gamma k+3 c k^{2}+12 \zeta b-24 b \gamma-16 \gamma k+6 c
\end{array}\right)}{3\left(b k^{4}+4 b k^{2}+2 k^{3}+2 b+4 k\right)}, \\
& \beta=-\frac{\left(\begin{array}{c}
3 b \gamma k^{5}-6 b \gamma k^{4}-20 b \gamma k^{3}+8 \gamma k^{4}-6 \zeta b k^{2} \\
+8 b \gamma k^{2}-12 \gamma k^{3}-24 b \gamma k-40 \gamma k^{2}-12 \zeta b \\
-12 \zeta k+24 b \gamma+32 \gamma k-3 c
\end{array}\right)}{6\left(b k^{4}+4 b k^{2}+2 k^{3}+2 b+4 k\right)},  \tag{5.3}\\
& \delta=\frac{\left(\begin{array}{c}
15 b \gamma k^{5}-18 b \gamma k^{4}+28 b \gamma k^{3}+32 \gamma k^{4} \\
-6 \zeta b k^{2}-40 \gamma k^{2}-36 \gamma k^{3}+8 b \gamma k^{2} \\
-12 \zeta b-12 \zeta k-16 \gamma k-3 c
\end{array}\right)}{24\left(b k^{4}+4 b k^{2}+2 k^{3}+2 b+4 k\right)}, \\
& w=\frac{\left(\begin{array}{c}
2 \gamma k^{8}+3 b c k^{5}+20 \gamma k^{6}-8 \gamma k^{5}+12 b c k^{3} \\
+3 c k^{4}-16 \gamma k^{4}+24 \gamma k^{3}+6 b c k+16 \gamma k^{2} \\
+24 \zeta k-32 \gamma k-6 c
\end{array}\right)}{3\left(b k^{4}+4 b k^{2}+2 k^{3}+2 b+4 k\right)}, \\
& \sigma=-\frac{1}{3} \gamma k .
\end{align*}
$$

By utilizing values in (5.3) into (5.1), we have

$$
\begin{equation*}
f=b_{1} \cos z+b_{1} \sin z \tag{5.4}
\end{equation*}
$$

Inserting Eq. (5.4) into (2.4) yields

$$
\begin{equation*}
y=-\frac{2(\sin (z)-\cos (z))}{\sin (z)+\cos (z)} \tag{5.5}
\end{equation*}
$$

By using Eq. (5.5) into Eq. (2.1), we obtain a first type of trigonometric solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=-\frac{2(\sin (x-c t)-\cos (x-c t)) e^{i(k x-w t)}}{\sin (x-c t)+\cos (x-c t)} \tag{5.6}
\end{equation*}
$$

(See Figure 5.1)

(a)

(b)

(c)

Figure 5.1: 3D plots of the rational solution (5.6) in Case 1 with the values of $k=2, c=1, \gamma=3, b=1, \zeta=-1$ (a) Real, (b) Imaginary and (c) Complex.

Case 2:

$$
\begin{equation*}
a=-b^{2} w+2 \zeta-4 \gamma-w, \quad \beta=\zeta-2 \gamma-\frac{1}{4} w, \quad c=-b w, \quad \delta=-\frac{1}{4} \zeta+\frac{1}{16} w, \quad k=0, \quad \sigma=0 \tag{5.7}
\end{equation*}
$$

Using the obtained values in (5.1) gives

$$
\begin{equation*}
f=b_{1} \cos z+b_{1} \sin z \tag{5.8}
\end{equation*}
$$

Inserting Eq. (5.8) into (2.4) yields

$$
\begin{equation*}
y=-\frac{2(\sin (z)-\cos (z))}{\sin (z)+\cos (z)} \tag{5.9}
\end{equation*}
$$

By using Eq. (5.9) into Eq. (2.1), we obtain a second type of trigonometric solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=-\frac{2(\sin (x-c t)-\cos (x-c t)) e^{i(k x-w t)}}{\sin (x-c t)+\cos (x-c t)} \tag{5.10}
\end{equation*}
$$

(See Figure 5.2)


Figure 5.2: 3D plots of the rational solution (5.10) in Case 2 with the values of $b=10, w=1$ (a) Real, (b) Imaginary and (c) Complex.

Case 3:

$$
\begin{align*}
& a=-\frac{\binom{-\gamma k^{11}-6 \gamma k^{9}-50 \gamma k^{7}+44 \gamma k^{6}+3 k^{6} w-44 \gamma k^{5}+12 \zeta k^{4}+72 \gamma k^{4}}{+18 k^{4} w+8 \gamma k^{3}+40 \gamma k^{2}+30 k^{2} w-16 \gamma k+48 \gamma+12 w-24 \zeta}}{3\left(k^{4}+4 k^{2}+2\right)^{2}}, \\
& \beta=\frac{\binom{-\gamma k^{9}+6 \gamma k^{8}+28 \gamma k^{7}+8 \gamma k^{6}+6 k^{6} \zeta+82 \gamma k^{5}-20 \gamma k^{4}-3 k^{4} w+36 k^{4} \zeta}{+152 \gamma k^{3}-144 \gamma k^{2}-12 k^{2} w+84 k^{2} \zeta+48 \gamma k-48 \gamma-6 w+24 \zeta}}{6\left(k^{4}+4 k^{2}+2\right)^{2}},  \tag{5.11}\\
& \delta=-\frac{\binom{-13 \gamma k^{9}+18 \gamma k^{8}-68 \gamma k^{7}+104 \gamma k^{6}+6 k^{6} \zeta-158 \gamma k^{5}+220 \gamma k^{4}}{-3 k^{4} w+36 k^{4} \zeta-40 \gamma k^{3}+48 \gamma k^{2}-12 k^{2} w+84 k^{2} \zeta+24 \zeta-6 w}}{24\left(k^{4}+4 k^{2}+2\right)^{2}},  \tag{5}\\
& c=\frac{2 k\left(\gamma k^{7}+10 \gamma k^{5}-4 \gamma k^{4}-8 \gamma k^{3}+12 \gamma k^{2}+8 \gamma k-16 \gamma+12 \zeta\right)}{3\left(k^{4}+4 k^{2}+2\right)}, \quad b=-\frac{2\left(k^{2}+2\right) k}{k^{4}+4 k^{2}+2}, \quad \sigma=-\frac{1}{3} \gamma k .
\end{align*}
$$

By utilizing values in (5.11) into (5.1), we have

$$
\begin{equation*}
f=b_{1} \cos z+b_{1} \sin z \tag{5.12}
\end{equation*}
$$

Inserting Eq. (5.12) into (2.4) yields

$$
\begin{equation*}
y=-\frac{2(\sin (z)-\cos (z))}{\sin (z)+\cos (z)} \tag{5.13}
\end{equation*}
$$

By using Eq. (5.13) into Eq. (2.1), we obtain a first type of trigonometric solution of Eq. (1.1):

$$
\begin{equation*}
q(x, t)=-\frac{2(\sin (x-c t)-\cos (x-c t)) e^{i(k x-w t)}}{\sin (x-c t)+\cos (x-c t)} \tag{5.14}
\end{equation*}
$$

(See Figure 5.3)


Figure 5.3: 3D plots of the rational solution (5.14) in Case 3 with the values of $k=5, b=10, \gamma=0.5, w=1, \zeta=-1$ (a) Real, (b) Imaginary and (c) Complex.

## 6. Conclusion

New soliton solutions of the LPD equation were obtained with three different current, systematic and powerful methods. In order to understand how the obtained solutions change under different conditions, the solutions obtained by appropriate selection of some parameters affecting the shape and velocity of the solitons are observed. In this context, to understand the mechanism of the original equation (1.1) real, imaginary and complex three-dimensional plots have been drawn for each case of solitons. This paper presents novel solutions of LPD equation that have not been reported in the literature before. Also, comparing with the existing literature, our result is complete and our method is simple and direct. By providing novel solutions, this study contributes to the knowledge of the dynamical aspects of various physical phenomena that are modeled by the LPD equation.

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# Exact Traveling Wave Solutions of the Schamel-KdV Equation with Two Different Methods 

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#### Abstract

The Schamel-Korteweg-de Vries (S-KdV) equation including a square root nonlinearity is very important pattern for the research of ion-acoustic waves in plasma and dusty plasma. As known, it is significant to discover the traveling wave solutions of such equations. Therefore, in this paper, some new traveling wave solutions of the $\mathrm{S}-\mathrm{KdV}$ equation, which arises in plasma physics in the study of ion acoustic solitons when electron trapping is present and also it governs the electrostatic potential for a certain electron distribution in velocity space, are constructed. For this purpose, the Bernoulli Sub-ODE and modified auxiliary equation methods are used. It has been shown that the suggested methods are effective and give different types of function solutions as: hyperbolic, trigonometric, power, exponential, and rational functions. The applied computational strategies are direct, efficient, concise and can be implemented in more complex phenomena with the assistant of symbolic computations. The results found in the paper are of great interest and may also be used to discover the wave sorts and specialities in several plasma systems.


## 1. Introduction

Nonlinear partial differential equations (NPDEs) are used to describe complex problems with numerous phenomena in different fields, including engineering, chemical kinematics, biology, wave theory, optics, physics, fluid mechanics, biomedical science, and others [1]- [4]. S-KdV equation, based upon both usual KdV equation (when $\alpha=0$ ) [5]- [10],

$$
u_{t}+\beta u u_{x}+\delta u_{x x x}=0
$$

and Schamel equation (when $\beta=0$ )

$$
u_{t}+\alpha u^{1 / 2} u_{x}+\delta u_{x x x}=0
$$

which was derived a German scientist Hans Schamel in 1973 has the form [11]- [13]

$$
\begin{equation*}
u_{t}+\left(\alpha u^{1 / 2}+\beta u\right)_{x}+\delta u_{x x x}=0, \quad \alpha \beta \neq 0 \tag{1.1}
\end{equation*}
$$

where $\alpha, \beta$ and $\delta$ are constants which they are refer to the activation trapping, the convection and the dispersion coefficients, respectively. The advantage of implementing the nonlinear $\mathrm{S}-\mathrm{KdV}$ equation to analyze dynamics of modulated waves in dispersive media lies in the diversity of its solutions [14]- [17]. Here we should point out that the $\mathrm{S}-\mathrm{KdV}$ equation has a stronger nonlinearity than the usual KdV

[^1]equation in that the single soliton solution possesses a smaller width and higher velocity [18]. This equation is contained in many physical phenomena involving electromagnetic theory, physical chemistry, geophysics and other fields are examples [19,20]. The square root in the nonlinear term then translates to lowest order some of the kinetic effects, associated with electron trapping [21]. Schamel [22] stated that when $u u_{x}$ is replaced by $\left(|u|^{3 / 2}\right)_{x}$, compared to the classical $K d V$ equation, the Schamel equation possesses a stronger nonlinearity, which reveals that the wave has a smaller width and higher velocity and exact traveling wave solutions for the regularized Schamel equation [23].To create different exact solutions and to notice their properties, various significant methods have been developed [19]- [21], [24]- [29].
The implementations of the Sub-ODE and modified auxiliary equation methods in this paper highlight our main motivation and indicates its capacity to handle nonlinear equations, permitting it to be utilized to solve many types of nonlinearity models. The body of our paper is structured as follows: Methodologies of the Sub-ODE and modified auxiliary equation methods and their detailed structures are given in Section 2. In Section 3, we apply the different methods, introduced in Section 2, to the Schamel-KdV equation to find the exact solutions. Different forms of the exact solutions are derived from these methods. Section 4 is devoted to graphical illustrations of the methods. A discussion section is presented in Section 5. Finally, Section 6 provides conclusions stemming from the results of our work.

## 2. The Bernoulli Sub-ODE Method [30]- [33]

Herein, we introduce the steps of the Bernoulli Sub-ODE method. Suppose, nonlinear partial differential equation is given by

$$
\begin{equation*}
F\left(v, v_{t}, v_{x}, v_{x x}, v_{x x x}, \ldots\right)=0 \tag{2.1}
\end{equation*}
$$

where $v=v(x, t)$ is wave function to be calculated.
Step 1: Apply the traveling wave transformation,

$$
\begin{equation*}
v(x, t)=v(\eta), \eta=k x-c t \tag{2.2}
\end{equation*}
$$

where $k$ is constant and $c$ is speed of the traveling wave. Substituting (2.2) into (2.1), then (2.1) converted to an ordinary differential equation:

$$
\begin{equation*}
F\left(v^{\prime}, v^{\prime \prime}, v^{\prime \prime \prime}, \ldots\right)=0 \tag{2.3}
\end{equation*}
$$

where $F$ is a polynomial in $v(\eta)$ and its derivatives.
Step 2: Presume solutions of (2.3) presented by a series in $G$ :

$$
\begin{equation*}
v(\eta)=\sum_{i=0}^{N} b_{i}(G(\eta))^{i} \tag{2.4}
\end{equation*}
$$

where $b_{i}(0 \leq i \leq N)$ are constants to be calculated, $b_{N} \neq 0$, and $G(\eta)$ satisfies the next ODE,

$$
\begin{equation*}
G^{\prime}(\eta)+\lambda G(\eta)=\mu G(\eta)^{2} \tag{2.5}
\end{equation*}
$$

which has the following solution:

$$
G(\eta)=\frac{1}{\frac{\mu}{\lambda}+d e^{\lambda \eta}}
$$

where $\lambda, \mu \neq 0$ are arbitrary constants.
Step 3: The positive integer $N$ determined by balancing the highest order derivative term with the highest power nonlinear term in (2.3).
Step 4: Replacing (2.4) into (2.3), we acquire a polynomial in $G(\eta)$. Gathering all terms with the same power and equating each one to zero. We get a system of equations which can be solved by using Mathematica program.

### 2.1. The modified auxiliary equation method (MAE) [34]- [37]

Main steps of the modified auxiliary equation method are explained as follows:
Step 1: Solution of (2.3) is given by:

$$
\begin{equation*}
v(\eta)=\sum_{i=0}^{N}\left(b_{i} a^{i f(\eta)}\right. \tag{2.6}
\end{equation*}
$$

where $f(\eta)$ satisfies the following ODE:

$$
\begin{equation*}
f^{\prime}=\frac{1}{\ln a}\left(\mu a^{-f(\eta)}+\sigma+\lambda a^{f(\eta)}\right) \tag{2.7}
\end{equation*}
$$

where $b_{i}(i=0,1,2, \ldots, N), b_{N} \neq 0, \lambda, \sigma$ and $\mu$, are constants to be calculated.
Step 2: In (2.3), $N$ is a positive integer determined via the homogeneous balance principle as illustrated before.
Step 3: Substituting (2.6) and (2.7) in (2.3), and gathering the terms which had like powers of $\left(a^{f(\eta)}\right)$ and putting their coffecients equal to zero, we obtain a set of algebraic equations, which can be solved by the aid of Mathematica program.

Step 4: There is various sets of solutions of (2.7):
Set 1: $\sigma^{2}-4 \lambda \mu<0$ and $\lambda \neq 0$,

$$
a^{f(\eta)}=\frac{-\sigma}{2 \lambda}+\frac{\sqrt{4 \mu \lambda-\sigma^{2}}}{2 \lambda} \tan \left(\frac{\sqrt{4 \mu \lambda-\sigma^{2}} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{-\sigma}{2 \lambda}+\frac{\sqrt{4 \mu \lambda-\sigma^{2}}}{2 \lambda} \cot \left(\frac{\sqrt{4 \mu \lambda-\sigma^{2}} \eta}{2}\right)
$$

Set 2: $\sigma^{2}-4 \lambda \mu>0$ and $\lambda \neq 0$,

$$
a^{f(\eta)}=\frac{-\sigma}{2 \lambda}-\frac{\sqrt{\sigma^{2}-4 \mu \lambda}}{2 \lambda} \tanh \left(\frac{\sqrt{\sigma^{2}-4 \mu \lambda} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{-\sigma}{2 \lambda}-\frac{\sqrt{\sigma^{2}-4 \mu \lambda}}{2 \lambda} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}-4 \mu \lambda} \eta}{2}\right)
$$

Set 3: $\sigma^{2}+4 \mu^{2}<0, \lambda \neq 0 \operatorname{and} \lambda=-\mu$,

$$
a^{f(\eta)}=\frac{\sigma}{2 \mu}-\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}}{2 \mu} \tan \left(\frac{\sqrt{-\sigma^{2}-4 \mu^{2}} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{\sigma}{2 \mu}-\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}}{2 \mu} \cot \left(\frac{\sqrt{-\sigma^{2}-4 \mu^{2}} \eta}{2}\right)
$$

Set 4: $\sigma^{2}+4 \mu^{2}>0, \lambda \neq 0$ and $\lambda=-\mu$,

$$
a^{f(\eta)}=\frac{\sigma}{2 \mu}+\frac{\sqrt{\sigma^{2}+4 \mu^{2}}}{2 \mu} \tanh \left(\frac{\sqrt{\sigma^{2}+4 \mu^{2}} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{\sigma}{2 \mu}+\frac{\sqrt{\sigma^{2}+4 \mu^{2}}}{2 \mu} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}+4 \mu^{2}} \eta}{2}\right)
$$

Set 5: $\sigma^{2}-4 \mu^{2}<0$ and $\lambda=\mu$,

$$
a^{f(\eta)}=\frac{-\sigma}{2 \mu}+\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}}{2 \mu} \tan \left(\frac{\sqrt{-\sigma^{2}+4 \mu^{2}} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{-\sigma}{2 \mu}+\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}}{2 \mu} \cot \left(\frac{\sqrt{-\sigma^{2}+4 \mu^{2}} \eta}{2}\right)
$$

Set 6: $\sigma^{2}-4 \mu^{2}>0$ and $\lambda=\mu$,

$$
a^{f(\eta)}=\frac{-\sigma}{2 \mu}-\frac{\sqrt{\sigma^{2}-4 \mu^{2}}}{2 \mu} \tanh \left(\frac{\sqrt{\sigma^{2}-4 \mu^{2}} \eta}{2}\right)
$$

or

$$
a^{f(\eta)}=\frac{-\sigma}{2 \mu}-\frac{\sqrt{\sigma^{2}-4 \mu^{2}}}{2 \mu} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}-4 \mu^{2}} \eta}{2}\right)
$$

Set 7: $\sigma^{2}=4 \lambda \mu$ and $\lambda=\mu$,

$$
a^{f(\eta)}=-\frac{2+\sigma \eta}{2 \lambda \eta}
$$

Set 8: $\lambda \mu<0, \sigma=0$ and $\lambda \neq 0$,

$$
\begin{aligned}
& a^{f(\eta)}=-\sqrt{\frac{-\mu}{\lambda}} \tanh (\sqrt{-\mu \lambda} \eta), \\
& \text { or } \\
& a^{f(\eta)}=-\sqrt{\frac{-\mu}{\lambda}} \operatorname{coth}(\sqrt{-\mu \lambda} \eta) .
\end{aligned}
$$

Set 9: $\sigma=0$ and $\mu=-\lambda$,

$$
a^{f(\eta)}=\frac{1+e^{-2 \lambda \eta}}{-1+e^{-2 \lambda \eta}}
$$

Set 10: $\mu=\lambda=0$,

$$
a^{f(\eta)}=\cosh (\sigma \eta)+\sinh (\sigma \eta)
$$

Set 11: $\mu=\sigma=h$ and $\lambda=0$,

$$
a^{f(\eta)}=e^{h \eta}-1
$$

Set 12: $\lambda=\sigma=h$ and $\mu=0$,

$$
a^{f(\eta)}=\frac{e^{h \eta}}{1-e^{h \eta}}
$$

Set 13: $\sigma=\lambda+\mu$,

$$
a^{f(\eta)}=-\frac{1-\mu e^{(\mu-\lambda) \eta}}{1-\lambda e^{(\mu-\lambda) \eta}}
$$

Set 14: $\sigma=-(\lambda+\mu)$,

$$
a^{f(\eta)}=\frac{\mu-e^{(\mu-\lambda) \eta}}{\lambda-e^{(\mu-\lambda) \eta}}
$$

Set 15: $\mu=0$,

$$
a^{f(\eta)}=\frac{\sigma e^{\sigma \eta}}{1-\lambda e^{\sigma \eta}}
$$

Set 16: $\lambda=\mu=\sigma \neq 0$,

$$
a^{f(\eta)}=\sqrt{3} \tan \left(\frac{\sqrt{3}}{2} \mu \eta\right)-1
$$

Set 17: $\lambda=\sigma=0$,

$$
a^{f(\eta)}=\mu \eta
$$

Set 18: $\mu=\sigma=0$,

$$
a^{f(\eta)}=\frac{-1}{\lambda \eta}
$$

Set 19: $\lambda=\mu$ and $\sigma=0$,

$$
a^{f(\eta)}=\tan (\mu \eta)
$$

Set 20: $\lambda=0$,

$$
a^{f(\eta)}=e^{\sigma \eta}-\frac{\mu}{\sigma}
$$

## 3. Applications of the Methods

Begin with the following transformation:

$$
v(x, t)=u(x, t)^{2}
$$

with wave transformation (2.2) into (1.1), we obtain following ordinary differential equation:

$$
2 k \beta u^{3} u^{\prime}+6 k^{3} \delta u^{\prime} u^{\prime \prime}+2 u\left((-c+k \alpha u) u^{\prime}+k^{3} \delta u^{\prime \prime \prime}\right)=0 .
$$

Integrating once with respect to $\eta$, we get

$$
\begin{equation*}
-c u^{2}+\frac{2}{3} k \alpha u^{3}+\frac{1}{2} k \beta u^{4}+2 k^{3} \delta\left(u^{\prime}\right)^{2}+2 k^{3} \delta u u^{\prime \prime}=0 \tag{3.1}
\end{equation*}
$$

Balancing $u u^{\prime \prime}$ with $u^{4}$ in (3.1), we get $4 N=2 N+2$, then $N=1$.

### 3.1. The Bernoulli sub-ODE method

Using (2.4), solution of (3.1) is given by

$$
\begin{equation*}
u(\eta)=b_{0}+b_{1} G(\eta) \tag{3.2}
\end{equation*}
$$

Substituting (3.2) in (3.1), then collecting terms of the same powers and putting their coefficients equal to zero, next system of equations are acquired :

$$
\begin{aligned}
& -c b_{0}^{2}+\frac{2}{3} k \alpha b_{0}^{3}+\frac{1}{2} k \beta b_{0}^{4}=0 \\
& -2 c b_{0} b_{1}+2 k^{3} \delta \lambda^{2} b_{0} b_{1}+2 k \alpha b_{0}^{2} b_{1}+2 k \beta b_{0}^{3} b_{1}=0 \\
& -6 k^{3} \delta \lambda \mu b_{0} b_{1}-c b_{1}^{2}+4 k^{3} \delta \lambda^{2} b_{1}^{2}+2 k \alpha b_{0} b_{1}^{2}+3 k \beta b_{0}^{2} b_{1}^{2}=0 \\
& 4 k^{3} \delta \mu^{2} b_{0} b_{1}-10 k^{3} \delta \lambda \mu b_{1}^{2}+\frac{2}{3} k \alpha b_{1}^{3}+2 k \beta b_{0} b_{1}^{3}=0 \\
& 6 k^{3} \delta \mu^{2} b_{1}^{2}+\frac{1}{2} k \beta b_{1}^{4}=0
\end{aligned}
$$

In what follows, we present the two sets of solution:
Set 1:

$$
\begin{equation*}
\delta=-\frac{4 \alpha^{2}}{75 k^{2} \beta \lambda^{2}}, \quad b_{0}=-\frac{4 \alpha}{5 \beta}, \quad b_{1}=\frac{4 \alpha \mu}{5 \beta \lambda}, \quad c=-\frac{16 k \alpha^{2}}{75 \beta}, \quad v(x, t)=\left(-\frac{4 \alpha}{5 \beta}+\frac{4 \alpha \mu}{5 \beta \lambda}\left(\frac{1}{\left.\frac{\mu}{\lambda}+d e^{\lambda\left(k x+\left(\frac{16 k \alpha^{2}}{75 \beta}\right) t\right)}\right)}\right)\right. \tag{3.3}
\end{equation*}
$$

Set 2:

$$
\begin{equation*}
\delta=-\frac{4 \alpha^{2}}{75 k^{2} \beta \lambda^{2}}, \quad b_{0}=0, \quad b_{1}=-\frac{4 \alpha \mu}{5 \beta \lambda}, \quad c=-\frac{16 k \alpha^{2}}{75 \beta}, \quad v(x, t)=\left(-\frac{4 \alpha \mu}{5 \beta \lambda}\left(\frac{1}{\left.\frac{\mu}{\lambda}+d e^{\lambda\left(k x+\left(\frac{16 k \alpha^{2}}{75 \beta}\right) t\right.}\right)}\right)\right)^{2} \tag{3.4}
\end{equation*}
$$

### 3.2. The modified auxiliary equation method (MAE)

(2.6) presents the solution in the form:

$$
\begin{equation*}
u(\eta)=b_{0}+b_{1} a^{f(\eta)} \tag{3.5}
\end{equation*}
$$

Substituting (3.5)in (3.1), then summing terms of like powers and setting their coefficients equal to zero, the next system of equations are obtained:

$$
\begin{aligned}
& -c b_{0}^{2}+\frac{2}{3} k \alpha b_{0}^{3}+\frac{1}{2} k \beta b_{0}^{4}+2 k^{3} \delta \mu \sigma b_{0} b_{1}+2 k^{3} \delta \mu^{2} b_{1}^{2}=0 \\
& -2 c b_{0} b_{1}+4 k^{3} \delta \lambda \mu b_{0} b_{1}+2 k^{3} \delta \sigma^{2} b_{0} b_{1}+2 k \alpha b_{0}^{2} b_{1}+2 k \beta b_{0}^{3} b_{1}+6 k^{3} \delta \mu \sigma b_{1}^{2}=0 \\
& 6 k^{3} \delta \lambda \sigma b_{0} b_{1}-c b_{1}^{2}+8 k^{3} \delta \lambda \mu b_{1}^{2}+4 k^{3} \delta \sigma^{2} b_{1}^{2}+2 k \alpha b_{0} b_{1}^{2}+3 k \beta b_{0}^{2} b_{1}^{2}=0 \\
& 4 k^{3} \delta \lambda^{2} b_{0} b_{1}+10 k^{3} \delta \lambda \sigma b_{1}^{2}+\frac{2}{3} k \alpha b_{1}^{3}+2 k \beta b_{0} b_{1}^{3}=0 \\
& 6 k^{3} \delta \lambda^{2} b_{1}^{2}+\frac{1}{2} k \beta b_{1}^{4}=0
\end{aligned}
$$

Solving the previous system yields two sets of solutions:

$$
c=4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right), \quad \alpha= \pm \frac{15 k^{2} \delta \lambda \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{b_{1}^{2}}, \quad \beta=-\frac{12 k^{2} \delta \lambda^{2}}{b_{1}^{2}}, \quad b_{0}=\frac{\sigma b_{1} \pm \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{2 \lambda} .
$$

Therefore, using the above sets gives the solitary wave solutions to (2.5) in the following formulas:

$$
\left(k x-\left(4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right)\right) t\right)
$$

Set 1: $\sigma^{2}-4 \lambda \mu<0$ and $\lambda \neq 0$,

$$
v_{1,2}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{2 \lambda}+b_{1}\left(\frac{-\sigma}{2 \lambda}+\frac{\sqrt{4 \mu \lambda-\sigma^{2}}}{2 \lambda} \tan \left(\frac{\sqrt{4 \mu \lambda-\sigma^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

or

$$
v_{3,4}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{2 \lambda}+b_{1}\left(\frac{-\sigma}{2 \lambda}+\frac{\sqrt{4 \mu \lambda-\sigma^{2}}}{2 \lambda} \cot \left(\frac{\sqrt{4 \mu \lambda-\sigma^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

Set 2: $\sigma^{2}-4 \lambda \mu>0$ and $\lambda \neq 0$,

$$
v_{5,6}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{2 \lambda}+b_{1}\left(\frac{-\sigma}{2 \lambda}-\frac{\sqrt{\sigma^{2}-4 \mu \lambda}}{2 \lambda} \tanh \left(\frac{\sqrt{\sigma^{2}-4 \mu \lambda}\left(k x-\left(4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

or

$$
\begin{equation*}
v_{7,8}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \lambda \mu+\sigma^{2}\right) b_{1}^{2}}}{2 \lambda}+b_{1}\left(\frac{-\sigma}{2 \lambda}-\frac{\sqrt{\sigma^{2}-4 \mu \lambda}}{2 \lambda} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}-4 \mu \lambda}\left(k x-\left(4 k^{3} \delta\left(-4 \lambda \mu+\sigma^{2}\right)\right) t\right)}{2}\right)\right)^{2}\right. \tag{3.6}
\end{equation*}
$$

Set 3: $\sigma^{2}+4 \mu^{2}<0, \lambda \neq 0$ and $\lambda=-\mu$,

$$
v_{9,10}(x, t)=\left(-\frac{\sigma b_{1} \pm \sqrt{\left(4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(\frac{\sigma}{2 \mu}-\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}}{2 \mu} \tan \left(\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

or

$$
v_{11,12}(x, t)=\left(-\frac{\sigma b_{1} \pm \sqrt{\left(4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(\frac{\sigma}{2 \mu}-\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}}{2 \mu} \cot \left(\frac{\sqrt{-\sigma^{2}-4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

Set 4: $\sigma^{2}+4 \mu^{2}>0, \lambda \neq 0$ and $\lambda=-\mu$,

$$
v_{13,14}(x, t)=\left(-\frac{\sigma b_{1} \pm \sqrt{\left(4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(\frac{\sigma}{2 \mu}+\frac{\sqrt{\sigma^{2}+4 \mu^{2}}}{2 \mu} \tanh \left(\frac{\sqrt{\sigma^{2}+4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

or

$$
v_{15,16}(x, t)=\left(-\frac{\sigma b_{1} \pm \sqrt{\left(4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(\frac{\sigma}{2 \mu}+\frac{\sqrt{\sigma^{2}+4 \mu^{2}}}{2 \mu} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}+4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

Set 5: $\sigma^{2}-4 \mu^{2}<0$ and $\lambda=\mu$,
$v_{17,18}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(-\frac{\sigma}{2 \mu}+\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}}{2 \mu} \tan \left(\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}$,
or

$$
\begin{equation*}
v_{19,20}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(-\frac{\sigma}{2 \mu}+\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}}{2 \mu} \cot \left(\frac{\sqrt{-\sigma^{2}+4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2} . \tag{3.7}
\end{equation*}
$$

Set 6: $\sigma^{2}-4 \mu^{2}>0$ and $\lambda=\mu$,

$$
v_{21,22}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(-\frac{\sigma}{2 \mu}-\frac{\sqrt{\sigma^{2}-4 \mu^{2}}}{2 \mu} \tanh \left(\frac{\sqrt{\sigma^{2}-4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

or

$$
v_{23,24}(x, t)=\left(\frac{\sigma b_{1} \pm \sqrt{\left(-4 \mu^{2}+\sigma^{2}\right) b_{1}^{2}}}{2 \mu}+b_{1}\left(-\frac{\sigma}{2 \mu}-\frac{\sqrt{\sigma^{2}-4 \mu^{2}}}{2 \mu} \operatorname{coth}\left(\frac{\sqrt{\sigma^{2}-4 \mu^{2}}\left(k x-\left(4 k^{3} \delta\left(-4 \mu^{2}+\sigma^{2}\right)\right) t\right)}{2}\right)\right)\right)^{2}
$$

Set 7: $\sigma^{2}=4 \lambda \mu$ and $\lambda=\mu$,

$$
v_{25,26}(x, t)=\left(\frac{\sqrt{\lambda \mu} b_{1}}{\lambda}-\frac{(2+2 k x \sqrt{\lambda \mu}) b_{1}}{2 k x \lambda}\right)^{2}
$$

Set 8: $\lambda \mu<0, \sigma=0$ and $\lambda \neq 0$,

$$
\begin{equation*}
v_{27,28}(x, t)=\left( \pm \frac{\sqrt{-\lambda \mu b_{1}^{2}}}{\lambda}-\sqrt{-\frac{\mu}{\lambda}} b_{1} \tanh \left(\sqrt{-\lambda \mu}\left(k x+16 k^{3} t \delta \lambda \mu\right)\right)\right)^{2} \tag{3.8}
\end{equation*}
$$

or

$$
v_{29,30}(x, t)=\left( \pm \frac{\sqrt{-\lambda \mu b_{1}^{2}}}{\lambda}-\sqrt{-\frac{\mu}{\lambda}} b_{1} \operatorname{coth}\left(\sqrt{-\lambda \mu}\left(k x+16 k^{3} t \delta \lambda \mu\right)\right)\right)^{2}
$$

Set 9: $\sigma=0$ and $\mu=-\lambda$,

$$
v_{31,32}(x, t)=\left( \pm \frac{\sqrt{\lambda^{2} b_{1}^{2}}}{\lambda}+b_{1}\left(\frac{1+e^{-2 \lambda\left(k x-16 k^{3} \delta \lambda^{2} t\right)}}{-1+e^{-2 \lambda\left(k x-16 k^{3} \delta \lambda^{2} t\right)}}\right)\right)^{2}
$$

Set 10: $\lambda=\sigma=h$ and $\mu=0$,

$$
v_{33,34}(x, t)=\left(\frac{h b_{1} \pm \sqrt{h^{2} b_{1}^{2}}}{2 h}+b_{1}\left(\frac{e^{h\left(k x-4 h^{2} k^{3} \delta t\right)}}{1-e^{h\left(k x-4 h^{2} k^{3} \delta t\right)}}\right)\right)^{2}
$$

Set 11: $\sigma=\lambda+\mu$,

$$
v_{35,36}(x, t)=\left(\frac{b_{1}(\lambda+\mu) \pm b_{1}(\lambda-\mu)}{2 \lambda}-b_{1}\left(\frac{1-\mu e^{(\mu-\lambda)\left(k x-4 k^{3} \delta(\lambda-\mu)^{2} t\right)}}{1-\lambda e^{(\mu-\lambda)\left(k x-4 k^{3} \delta(\lambda-\mu)^{2} t\right)}}\right)\right)^{2}
$$

Set 12: $\sigma=-(\lambda+\mu)$,

$$
v_{37,38}(x, t)=\left(\frac{b_{1}(-\lambda-\mu) \pm b_{1}(\lambda-\mu)}{2 \lambda}+b_{1}\left(\frac{\mu-e^{(\mu-\lambda)\left(k x-4 k^{3} \delta(\lambda-\mu)^{2} t\right)}}{\lambda-e^{(\mu-\lambda)\left(k x-4 k^{3} \delta(\lambda-\mu)^{2} t\right)}}\right)\right)^{2}
$$

Set 13: $\mu=0$,

$$
v_{39,40}(x, t)=\left(\frac{\sigma b_{1} \pm \sigma b_{1}}{2 \lambda}+b_{1}\left(\frac{\sigma e^{\sigma\left(k x-4 k^{3} \delta \sigma^{2} t\right)}}{1-\lambda e^{\sigma\left(k x-4 k^{3} \delta \sigma^{2} t\right)}}\right)\right)^{2}
$$

Set 14: $\lambda=\mu=\sigma \neq 0$,

$$
v_{41,42}(x, t)=\left(\frac{b_{1} \pm \sqrt{-3 b_{1}^{2}}}{2}+\frac{1}{2} b_{1}\left(-1+\sqrt{3} \tan \left(\frac{\sqrt{3}}{2} \sigma\left(k x+12 k^{3 \delta} \sigma^{2} t\right)\right)\right)\right)^{2}
$$

Set 15: $\mu=\sigma=0$,

$$
v_{43,44}(x, t)=\frac{b_{1}^{2}}{k^{2} \lambda^{2} x^{2}} .
$$

Set 16: $\lambda=\mu$ and $\sigma=0$,

$$
v_{45,46}(x, t)=\left( \pm \sqrt{-b_{1}^{2}}+b_{1} \tan \left(\mu\left(k x+16 k^{3} \delta \mu^{2} t\right)\right)\right)^{2}
$$

## 4. Graphical Illustrations

The majority of our solutions are presented in the following graphs to illustrate solutions.
In Figure 4.1, we present graph of (3.3) using the Bernoulli Sub-ODE method at $k=2, \alpha=0.5, \beta=0.3, \mu=0.3, \lambda=0.3, d=2$. Figure 4.2 shows graph of (3.4) using the Bernoulli Sub-ODE method at $k=2, \alpha=0.5, \beta=0.3, \mu=0.3, \lambda=0.3, d=2$. Graph of (3.6) using the modified auxiliary equation method at $k=2, b_{1}=0.3, \mu=0.02, \lambda=0.1, \delta=0.1, \sigma=0.3$ is presented in Figure 4.3. Graph of (3.7) using the modified auxiliary equation method at $k=0.6, b_{1}=0.1, \mu=0.03, \delta=1.6, \sigma=0.04$ is given in Figure 4.4. Lastly, Figure 4.5 presents graph of (3.8) using the modified auxiliary equation method at $k=0.7, b_{1}=0.1, \mu=0.3, \delta=0.3, \sigma=0, \lambda=-0.5$.



Figure 4.1: Graph of Eq. (3.3) using the Bernoulli Sub-ODE method at $k=2, \alpha=0.5, \beta=0.3, \mu=0.3, \lambda=0.3, d=2$.


Figure 4.2: Graph of Eq. (3.4) using the Bernoulli Sub-ODE method at $k=2, \alpha=0.5, \beta=0.3, \mu=0.3, \lambda=0.3, d=2$.


Figure 4.3: Graph of Eq. (3.6) using the modified auxiliary equation method at $k=2, b_{1}=0.3, \mu=0.02, \lambda=0.1, \delta=0.1, \sigma=0.3$.


Figure 4.4: Graph of Eq. (3.7) using the modified auxiliary equation method at $k=0.6, b_{1}=0.1, \mu=0.03, \delta=1.6, \sigma=0.04$.


Figure 4.5: Graph of Eq. (3.8) using the modified auxiliary equation method at $k=0.7, b_{1}=0.1, \mu=0.3, \delta=0.3, \sigma=0, \lambda=-0.5$.

## 5. Discussion

The graph is one of the best tools for describing and presenting solutions. In the following, we review the behavior of the wave in the solutions presented: In Figures 4.1-4.2 the wave travels to the left with increasing time $t=0,5,10$. Contrarily, in Figures 4.3-4.5 the wave moves towards left as time passes $t=0,5,10$. The flipped wave is presented in Figure 4.4 as time goes on.

## 6. Conclusion

In this work, a class of some new travelling wave solutions of the Schamel-Korteweg-de Vries equation are successfully found out by using the Bernoulli Sub-ODE and modified auxiliary equation methods. The Bernoulli Sub-ODE is a simple and straightforward method and is applicable to a wide range of problems in science and engineering, but it can be time-consuming to apply the method if the equation involves complex functions. The modified auxiliary equation method can be used to solve a wide range of differential equations, also it can provide closed-form solutions, but it can be difficult to determine the appropriate auxiliary equation to use for a given differential equation. The presented exact solutions provided here may describe various new characteristics of waves and then may be useful in the theoretical and numerical studies of the considered equation. A graphical representation of newly discovered solutions are also shown to explain the dynamics of soliton profiles. The found new soliton solutions of the $\mathrm{S}-\mathrm{KdV}$ equation are of significant importance and can be used in other areas of physics such as plasma physics.

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# Characterizations of Matrix and Compact Operators on BK Spaces 

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#### Abstract

In the present paper, by estimating operator norms, we give some characterizations of infinite matrix classes $\left(\left|E_{\mu}^{r}\right|_{q}, \Lambda\right)$ and $\left(\left|E_{\mu}^{r}\right|_{\infty}, \Lambda\right)$, where the absolute spaces $\left|E_{\mu}^{r}\right|_{q},\left|E_{\mu}^{r}\right|_{\infty}$ have been recently studied by Gökçe and Sarıgöl [1] and $\Lambda$ is one of the well-known spaces $c_{0}, c, l_{\infty}, l_{q}(q \geq 1)$. Also, we obtain necessary and sufficient conditions for each matrix in these classes to be compact establishing their identities or estimates for the Hausdorff measures of noncompactness.


## 1. Introduction

The summability theory is one of the most important field in mathematics specially analysis, applied mathematics, engineering sciences, quantum mechanics and probability theory, therefore, it has been chosen as the subject of study by many authors. The theory of sequence space, which is one of the main topics of the summability theory, is mainly about generalizing the concepts of convergence-divergence for sequences and series. In this context, the primary aim is to assign a limit value for non-convergent sequences or series by using a transformation given by the most general linear mappings of infinite matrices. So, several studies can be traced in the literature dealing with characterization of matrix transformation between special sequence spaces. To mention few of them are [2-6]. On the one hand, from a different perspective, using the notion of absolute summability, a lot of new spaces of series summable by the absolute summability methods have studied and introduced by authors (see [1,7-14]). In recent paper [1], the infinite matrix classes $\left(\left|E_{\mu}^{r}\right|,\left|E_{\mu}^{r}\right|_{q}\right)$ and $\left(\left|E_{\mu}^{r}\right|_{q},\left|E_{\mu}^{r}\right|\right)$ have been investigated. In the present paper, the matrix classes $\left(\left|E_{\mu}^{r}\right|_{q}, \Lambda\right)$ and $\left(\left|E_{\mu}^{r}\right|_{\infty}, \Lambda\right)$ have been characterized with operator norms, where $1 \leq q<\infty$ and $\Lambda \in\left\{c, c_{0}, l_{\infty}, l_{q}\right\}$. Besides, establishing their identities or estimates for the Hausdorff measures of noncompactness, the necessary and sufficient conditions for each matrix in these classes to be compact have been investigated.

A linear subspaces of $\omega$, the set of all sequences of complex numbers, is called a sequence space. Let $\Delta, \Gamma$ be any subspaces of $\omega$ and $U=\left(u_{n v}\right)$ be any infinite matrix of complex components. The transform of a sequence $\delta=\left(\delta_{v}\right) \in \omega$ is the sequence $U(\boldsymbol{\delta})$ deduced by the usual matrix product and its terms are written as

$$
U_{n}(\boldsymbol{\delta})=\sum_{j=0}^{\infty} u_{n j} \delta_{j}
$$

provided that the series converges for all $n \geq 0$. Then, $U$ is called a matrix mapping from the space $\Delta$ into the another spaces $\Gamma$, if the sequence $U(\delta)$ exists and $U(\delta) \in \Gamma$ for all $\delta \in \Delta$. The collection, containing all such infinite matrices, is denoted by $(\Delta, \Gamma)$. A triangle matrix $U$ is given as $u_{n n} \neq 0$ for all $n$ and $u_{n j}=0$ for $n>j$.
The concept of domain of an infinite matrix $U$ in the $\Delta$ is described by

$$
\begin{equation*}
\Lambda_{U}=\left\{\delta=\left(\delta_{n}\right) \in \omega: U(\delta) \in \Delta\right\} \tag{1.1}
\end{equation*}
$$

and also the $\beta$-dual of the sequence space $\Lambda$ is given by the set

$$
\Delta^{\beta}=\left\{y: \sum_{v=0}^{\infty} y_{v} \delta_{v} \text { converges for all } \delta \in \Delta\right\} .
$$

If $\Delta \subset \omega$ is a Frechet space that is a complete locally convex linear metric space, on which all coordinate functionals $r_{n}(\delta)=\delta_{n}$ are continuous for all $n$, then it is said to be an FK space; an FK space whose metric is given by a norm is called a BK space.
$B K$-spaces have a significant role in summability theory. For instance, the matrix operators between $B K$-spaces are continuous and when $\Delta$ is a $B K$-space, the matrix domain $\Delta_{U}$ is also a $B K$-space, and also its norm is given by

$$
\|\delta\|_{\Delta_{U}}=\|U(\delta)\|_{\Delta}
$$

A $B K$-space $\Lambda \supset \phi$ is said to have $A K$ property if, for all sequence $\delta=\left(\delta_{v}\right) \in \Delta$, there is a unique representation $\delta=\sum_{v=0}^{\infty} \delta_{v} e^{(v)}$ where $\left(e^{(v)}\right)$ is the sequence whose only nonzero term is 1 in $v$-th place for $v \geq 0$ and $\phi$ is the set of all finite sequences. For example, while the space $l_{\infty}$ does not have $A K$ property, the sequence space $l_{q}$ has $A K$ property in respect to its natural norm where $q \geq 1$.

Let $\Delta$ and $\Gamma$ be two Banach spaces. The set of all continuous linear operators from $\Delta$ into $\Gamma$ is represented by $\mathscr{B}(\Delta, \Gamma)$ and, for $U \in \mathscr{B}(\Delta, \Gamma)$, the norm of $U$ is stated by

$$
\|U\|=\sup _{\delta \in S_{\Delta}}\|U(\delta)\|_{\Gamma} .
$$

If $y \in \omega$ and $\Delta \supset \phi$ is a $B K$-space, then

$$
\|y\|_{\Delta}^{*}=\sup _{\delta \in S_{\Delta}}\left|\sum_{k=0}^{\infty} y_{k} \delta_{k}\right|,
$$

and it is finite for $y \in \Delta^{\beta}$. Here, $S_{\Delta}^{\prime}$ is the unit sphere in $\Delta$.
Throughout this study, $\mu=\left(\mu_{n}\right)$ is any sequence of positive real numbers, $U=\left(u_{n j}\right)$ be an infinite matrix of complex components for all $n, j \geq 0$ and $q^{*}$ is the conjugate of $q$, that is $1 / q+1 / q^{*}=1$ for $q>1$, and $1 / q^{*}=0$ for $q=1$.
Let $\sum \delta_{k}$ be an infinite series with partial sums $s_{n}$, and $\left(\mu_{n}\right)$ be a sequence of positive terms. The series $\sum \delta_{v}$ is said to be summable $\left|U, \mu_{n}\right|_{q}$, $1 \leq q<\infty$, if (see [15])

$$
\begin{equation*}
\sum_{n=0}^{\infty} \mu_{n}^{q-1}\left|U_{n}(s)-U_{n-1}(s)\right|^{q}<\infty, \tag{1.2}
\end{equation*}
$$

where $U_{-1}(s)=0$.
Point out that the method includes certain well known methods. For instance, for Cesàro matrix with $\mu_{n}=n$ and the weighted mean matrix, it reduces to the absolute Cesàro summability due to Flett [7] and the absolute weighted summability given by Sulaiman [6], respectively. For more applications, we refer readers to ( $[1,8-10,12]$ ).
Also, if we choose the Euler matrix $E^{r}=\left(e_{n i}^{r}\right)$ instead of $U$, the summability $\left|U, \mu_{n}\right|_{q}$ is reduced to the absolute Euler summability $\left|E^{r}, \mu_{n}\right|_{q}$ of order $r$. Here the terms of the matrix $E^{r}=\left(e_{n i}^{r}\right)$ is given by

$$
e_{n i}^{r}=\left\{\begin{array}{l}
\binom{n}{i}(1-r)^{n-i} r^{i}, \quad 0 \leq i \leq n \\
0, \quad i>n
\end{array}\right.
$$

for all $n, i \geq 0$ and $0<r \leq 1$, [1].
The spaces of all series summable by the methods $\left|E^{r}, \mu_{n}\right|_{q}, 1 \leq q<\infty$, and $\left|E^{r}, \mu_{n}\right|_{\infty}$ have recently been introduced by Gökçe and Sarigöl [1] as follows:

$$
\begin{aligned}
& \left|E_{\mu}^{r}\right|_{q}=\left\{\delta=\left(\delta_{v}\right): \sum_{n=1}^{\infty}\left|T_{n}^{r}(q)(\delta)\right|^{q}<\infty\right\} \\
& \left|E_{\mu, q}^{r}\right|_{\infty}=\left\{\delta=\left(\delta_{v}\right): \sup _{n}\left|T_{n}^{r}(q)(\delta)\right|<\infty\right\}
\end{aligned}
$$

where $T_{0}^{r}(q)(\boldsymbol{\delta})=\delta_{0}$ and

$$
\begin{equation*}
T_{n}^{r}(q)(\delta)=\mu_{n}^{1 / q^{*}} \sum_{i=1}^{n}\binom{n-1}{i-1}(1-r)^{n-i} r^{i} \delta_{i} . \tag{1.3}
\end{equation*}
$$

Also, with the notation of domain, we can state $\left|E_{\mu}^{r}\right|_{q}=\left(l_{q}\right)_{T^{r}(q)}$ and $\left|E_{\mu, q}^{r}\right|_{\infty}=\left(l_{\infty}\right)_{T^{r}(q)}$, if we define the matrix $T^{r}(q)=\left(t_{n j}^{r}(q)\right)$ by

$$
t_{n i}^{r}(q)=\left\{\begin{array}{l}
\left.\mu_{n}^{1 / q^{*}} \begin{array}{c}
n-1 \\
i-1 \\
i-1
\end{array}\right)(1-r)^{n-i} r^{i}, \quad 1 \leq i \leq n \\
0, \quad i>n .
\end{array}\right.
$$

The inverse transformation of $T_{n}^{r}(q)$ can be written as

$$
\begin{equation*}
\delta_{n}=\sum_{i=1}^{n} \mu_{i}^{-1 / q^{*}}\binom{n-1}{i-1}(r-1)^{n-i} r^{-n} T_{i}^{r}(q)(\delta), \tag{1.4}
\end{equation*}
$$

[1].
Now, we list some known lemmas:

Lemma 1.1 ([1]). Let $1 \leq q<\infty$. The spaces $\left|E_{\mu}^{r}\right|_{q}$ and $\left|E_{\mu, q}^{r}\right|_{\infty}$ are BK-spaces with the norms $\|\delta\|_{\left|E_{\mu}^{r}\right|_{q}}=\left\|T^{r}(q)(\delta)\right\|_{l_{q}}$ and $\|\delta\|_{\left|E_{\mu, q}^{r}\right|_{\infty}}=$ $\left\|T^{r}(q)(\delta)\right\|_{\infty}$. Also, these are linearly isomorphic to the space $l_{q}$ and $l_{\infty}$, respectively.

Lemma 1.2 ( [16]). The following statements hold:

1. $U \in(l, c)$ iff (i) $\lim _{n} u_{n j}$ exists for all $j \geq 0,(i i) \sup _{n, j}\left|u_{n j}\right|<\infty$, $U \in\left(l, l_{\infty}\right)$ iff (ii) holds.
2. If $1<q<\infty$, then, $U \in\left(l_{q}, c\right)$ if and only if (i)holds, (iii) $\sup _{n} \sum_{j=0}^{\infty}\left|u_{n j}\right|^{q^{*}}<\infty$, $U \in\left(l_{q}, l_{\infty}\right)$ iff $(i i i)$ holds.
3. $U \in\left(l, c_{0}\right)$ iff (iv) $\lim _{n} u_{n j}=0$ for all $j \geq 0$, (ii) hold.
4. If $1<q<\infty$, then, $U \in\left(l_{q}, c_{0}\right)$ iff (iii) and (iv) hold.
5. $U \in\left(l_{\infty}, c\right)$ iff (i) holds, (v) $\sum_{j=0}^{\infty}\left|u_{n j}\right|<\infty$ uniformly in $n$,

$$
U \in\left(l_{\infty}, l_{\infty}\right) \text { iff }(v i) \sup _{n} \sum_{j=0}^{\infty}\left|u_{n j}\right|<\infty .
$$

6. $U \in\left(l_{\infty}, c_{0}\right)$ iff $(v i i) \lim _{n} \sum_{j=0}^{\infty}\left|u_{n j}\right|=0$.
7. If $1 \leq p<\infty$, then $U \in\left(l_{\infty}, l_{q}\right)$ iff (viii) $\sup _{K} \sum_{n=0}^{\infty}\left|\sum_{k \in K}^{\infty} u_{n j}\right|^{q}<\infty$.

Lemma 1.3 ([17]). Let $1 \leq q<\infty$. Then, $U \in\left(l, l_{q}\right)$ iff

$$
\|U\|_{\left(l, l_{q}\right)}=\sup _{j}\left\{\sum_{n=0}^{\infty}\left|u_{n j}\right|^{q}\right\}^{1 / p}<\infty .
$$

Lemma 1.4 ([16]). Let $1<q<\infty$. Then, $U \in\left(l_{q}, l\right)$ iff

$$
\|U\|_{\left(l_{q}, l\right)}=\sup _{N \in \mathfrak{T}}\left\{\sum_{j=0}^{\infty}\left|\sum_{n \in N}^{\infty} u_{n j}\right|^{q^{*}}\right\}^{1 / q^{*}}<\infty
$$

where $\mathfrak{T}$ stands for the collection of all finite subsets of $\mathbb{N}$.
It is difficult to apply Lemma 1.4 in applications. The following lemma presents to us an equivalent applicable norm.
Lemma 1.5 ([18]). Let $1<q<\infty$. Then, $U \in\left(l_{q}, l\right)$ iff

$$
\|U\|_{(l q, l)}^{\prime}=\left\{\sum_{j=0}^{\infty}\left(\sum_{n=0}^{\infty}\left|u_{n j}\right|\right)^{q^{*}}\right\}^{1 / q^{*}}<\infty .
$$

Since $\|U\|_{\left(l_{q}, l\right)} \leq\|U\|_{\left(l_{q}, l\right)}^{\prime} \leq 4\|U\|_{\left(l_{q}, l\right)}$, there exists $\zeta \in[1,4]$ such that $\|U\|_{\left(l_{q}, l\right)}^{\prime}=\zeta\|U\|_{\left(l_{q}, l\right)}$.
Using the Hausdorff measure of noncompactness $\chi$ introduced in [19], characterizations of compact operators on great number of sequence spaces are investigated by many researchers. For instance, to characterize the class of compact operators on several spaces, the Hausdorff measure of noncompactness method have been used by Malkowsky and Rakocevic in [20], Mursaleen and Noman in [21, 22], (see also [1,23-26]).
Let $(\Delta, d)$ be a metric space and $Q$ be a bounded subset of $\Delta$. Then, $\chi$ and the number

$$
\chi(Q)=\inf \{\varepsilon>0: Q \text { has a finite } \varepsilon-\text { net in } \Delta\}
$$

are called the Hausdorff measure of noncompactness and the Hausdorff measure of noncompactness of $Q$, respectively.
Suppose that $S$ is a linear operator between the Banach spaces $\Delta$ and $\Gamma$ such that $S: \Delta \rightarrow \Gamma$. Then, it is said that $S$ is compact if its domain is all of $\Delta$ and, for every bounded sequence $\left(\delta_{n}\right)$ in $\Delta$, the sequence $\left(S\left(\delta_{n}\right)\right)$ has a convergent subsequence in $\Gamma$.
Lemma 1.6 ( [27]). Let $Q \subset \Delta$ be a bounded set where $\Delta$ is one of the normed spaces $c_{0}$ or $l_{q}$ for $1 \leq q<\infty$. If $R_{r}: \Delta \rightarrow \Delta$ is the operator defined by $R_{r}(y)=\left(y_{0}, y_{1}, \ldots y_{r}, 0,0, \ldots\right)$ for all $y \in \Delta$, then

$$
\chi(Q)=\lim _{r \rightarrow \infty}\left(\sup _{\delta \in Q}\left\|\left(I-R_{r}\right)(y)\right\|\right)
$$

Let $\chi_{1}, \chi_{2}$ be Hausdorff measures on $\Delta$ and $\Gamma$. If $S(Q)$ is a bounded subset of $\Gamma$ and there exists $M>0$ such that $\chi_{2}(S(Q)) \leq M \chi_{1}(Q)$ for each bounded subset $Q$ of $\Delta$, then the linear operator $S: \Delta \rightarrow \Gamma$ is called ( $\chi_{1}, \chi_{2}$ )- bounded. If an operator $S$ is $\left(\chi_{1}, \chi_{2}\right)$ - bounded, then the number

$$
\|S\|_{\left(\chi_{1}, \chi_{2}\right)}=\inf \left\{M>0: \chi_{2}(S(Q)) \leq M \chi_{1}(Q) \text { for all bounded sets } Q \subset \Delta\right\}
$$

is called the $\left(\chi_{1}, \chi_{2}\right)$-measure noncompactness of $L$. Also, in case of $\chi_{1}=\chi_{2}=\chi$, it is written by $\|S\|_{(\chi, \chi)}=\|S\|_{\chi}$.

Lemma 1.7 ([28]). $L \in \mathscr{B}(\Delta, \Gamma)$ and $S_{\delta}^{\prime}=\{\delta \in \Delta:\|\delta\| \leq 1\}$ be the unit ball in $\Delta$. Then,

$$
\|S\|_{\chi}=\chi\left(S\left(S_{\delta}^{\prime}\right)\right)
$$

and

$$
\text { S is compact } \Leftrightarrow\|S\|_{\chi}=0 .
$$

Lemma 1.8 ([29]). Let $T=\left(t_{n v}\right)$ be an infinite triangle matrix, $\Delta$ be a normed sequence space and $\chi_{T}$ and $\chi$ stand for the Hausdorff measures of noncompactness on $M_{\Delta_{T}}$ and $M_{\Delta}$, the collections of all bounded sets in $\Delta_{T}$ and $\Delta$, respectively. Then, $\chi_{T}(Q)=\chi(T(Q))$ for each $Q \in M_{\Delta_{T}}$.
Lemma 1.9 ([22]). Let $\Delta=l_{\infty}$ or $\Delta \supset \phi$ be any BK-space with $A K$ property. If $U \in(\Delta, c)$, then

$$
\begin{aligned}
& \lim _{n \rightarrow \infty} u_{n k}=\lambda_{k} \text { exists for all } k, \\
& \lambda=\left(\lambda_{k}\right) \in \Delta^{\beta}, \\
& \sup _{n}\left\|U_{n}-\lambda\right\|_{X}^{*}<\infty, \\
& \lim _{n \rightarrow \infty} U_{n}(\delta)=\sum_{k=0}^{\infty} \lambda_{k} \delta_{k} \text { for each } \delta=\left(\delta_{k}\right) \in \Delta .
\end{aligned}
$$

Lemma 1.10 ([22]). Let $\Delta \supset \phi$ be a BK-space. Then,
(a) If $U \in\left(\Delta, c_{0}\right)$, then

$$
\left\|L_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty}\left(\sup _{n>i}\left\|U_{n}\right\|_{\Delta}^{*}\right) .
$$

(b) If the space $\Delta$ has $A K$ or $\Delta=l_{\infty}$ and $U \in(\Delta, c)$, then

$$
\frac{1}{2} \lim _{i \rightarrow \infty}\left(\sup _{n \geq i}\left\|U_{n}-\lambda\right\|_{\Delta}^{*}\right) \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{i \rightarrow \infty}\left(\sup _{n \geq i}\left\|U_{n}-\lambda\right\|_{\Delta}^{*}\right)
$$

where $\lambda=\left(\lambda_{k}\right)$ defined by $\lambda_{k}=\lim _{n \rightarrow \infty} u_{n k}$, for all $n \in \mathbb{N}$.
(c) If $U \in\left(\Delta, l_{\infty}\right)$, then

$$
0 \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{i \rightarrow \infty}\left(\sup _{n>i}\left\|U_{n}\right\|_{\Delta}^{*}\right) .
$$

## 2. Matrix and Compact Operators on the Spaces $\left|E_{\mu}^{r}\right|_{q}$ and $\left|E_{\mu, q}^{r}\right|_{\infty}$

In this part of the study, firstly, by computing operator norms we obtain some characterizations of infinite matrix classes $\left(\left|E_{\mu}^{r}\right|_{q}, \Lambda\right)$ and $\left(\left|E_{\mu}^{r}\right|_{\infty}, \Lambda\right)$, where $\Lambda$ is one of the spaces $c, c_{0}, l_{\infty}, l_{q}$ and $1 \leq q<\infty$. Moreover, we search the necessary and sufficient conditions for each matrix in these classes to be compact establishing their estimates or identities for the Hausdorff measures of noncompactness.

Lemma 2.1. Let $1 \leq q<\infty$. Then,
(i) If $u=\left(u_{v}\right) \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$, then, $\tilde{u}^{(q)}=\left(\tilde{u}_{v}^{(q)}\right) \in l_{q^{*}}$ for all $\delta \in\left|E_{\mu}^{r}\right|_{q}$
(ii) If $u=\left(u_{v}\right) \in\left\{\left|E_{\mu}^{r}\right|\right\}^{\beta}$, then, $\tilde{u}^{(1)}=\left(\tilde{u}_{v}^{(1)}\right) \in l_{\infty}$ for all $\delta \in\left|E_{\mu}^{r}\right|$
(iii) If $u=\left(u_{v}\right) \in\left\{\left|E_{\mu, q}^{r}\right|_{\infty}\right\}^{\beta}$, then, $\tilde{u}^{(q)}=\left(\tilde{u}_{v}^{(q)}\right) \in l$ for all $\delta \in\left|E_{\mu, q}^{r}\right|_{\infty}$
and the equality

$$
\begin{equation*}
\sum_{v=0}^{\infty} u_{v} \delta v=\sum_{v=0}^{\infty} \tilde{u}_{v}^{(q)} y_{v} \tag{2.1}
\end{equation*}
$$

holds, where $y=T^{r}(q)(\delta)$ is $T^{r}(q)$-transformation sequence of the sequence $\delta=\left(\delta_{v}\right)$ and

$$
\tilde{u}_{v}^{(q)}=\mu_{v}^{-1 / q^{*}} \sum_{n=v}^{\infty}\binom{n-1}{v-1}(r-1)^{n-v} r^{-n} u_{n}, \tilde{u}_{0}^{(q)}=u_{0} .
$$

Proof. (i) Let $u=\left(u_{v}\right) \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$. Considering (1.4) the equation (2.1) is obtained immediately. Also, it follows from Theorem 1.29 in [30] that $\tilde{u}^{(q)} \in l_{q^{*}}$ whenever $u \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$.
As (ii) and (iii) can be proved with similar lines, these parts are left to reader.

Lemma 2.2. Let $1<q<\infty$. Then, we have $\|u\|_{\left|E_{\mu}^{r}\right|_{q}}^{*}=\left\|\tilde{u}^{(q)}\right\|_{l_{q^{*}}}$ for all $u \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta},\|u\|_{\left|E_{\mu}^{r}\right|}^{*}=\left\|\tilde{u}^{(1)}\right\|_{l_{\infty}}$ for all $u \in\left\{\left|E_{\mu}^{r}\right|\right\}^{\beta}$ and $\|u\|_{\left.\right|_{\left.E_{\mu, q}^{r}\right|_{\infty}} ^{*}}=\left\|\tilde{u}^{(q)}\right\|_{l}$ for all $u \in\left\{\left|E_{\mu, q}^{r}\right|_{\infty}\right\}^{\beta}$.

Proof. Take $u \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$. Since $l_{q}^{\beta}=l_{q^{*}}$, we get $\tilde{u}^{(q)} \in l_{q^{*}}$. Also, it follows from Theorem 1.29 in [30] and Lemma 2.1 that

$$
\|u\|_{\left.E_{\mu}^{r}\right|_{q}}^{*}=\sup _{\delta \in S_{\left|E_{\mu}^{r}\right|_{q}}}\left|\sum_{v=0}^{\infty} u_{v} \delta_{v}\right|=\sup _{y \in S_{S_{q}}}\left|\sum_{v=0}^{\infty} \tilde{u}_{v}^{(q)} y_{v}\right|=\left\|\tilde{u}^{(q)}\right\|_{l_{q}}^{*}=\left\|\tilde{u}^{(q)}\right\|_{l_{q^{*}}} .
$$

For $u \in\left\{\left|E_{\mu}^{r}\right|\right\}^{\beta}$ and $u \in\left\{\left|E_{\mu, q}^{r}\right|_{\infty}\right\}^{\beta}$, the proofs are similar, so the proofs are omitted.
Theorem 2.3. Let $1 \leq q<\infty$. Further, let $W=\left(w_{n j}\right)$ be a matrix satisfying

$$
\begin{equation*}
w_{n j}=\mu_{n}^{1 / q^{*}} \sum_{i=1}^{n}\binom{n-1}{i-1}(1-r)^{n-i} r^{i} u_{i j} \tag{2.2}
\end{equation*}
$$

Then, $U \in\left(\Delta,\left|E_{\mu}^{r}\right|_{q}\right)$ equals to $W \in\left(\Delta, l_{q}\right)$, and $U \in\left(\Delta,\left|E_{\mu, q}^{r}\right|_{\infty}\right)$ if and only if $W \in\left(\Delta, l_{\infty}\right)$.
Proof. Let take $\lambda \in \Delta$. Then, considering (2.2) it can be written that

$$
\sum_{j=0}^{\infty} w_{n j} \delta_{j}=\mu_{n}^{1 / q^{*}} \sum_{v=1}^{n}\binom{n-1}{v-1}(1-r)^{n-v} r^{v} \sum_{j=0}^{\infty} u_{j v} \delta_{j}
$$

which implies that $W_{n}(\boldsymbol{\delta})=T_{n}^{r}(q)(U(\boldsymbol{\delta}))$. This shows that $U_{n}(\boldsymbol{\delta}) \in\left|E_{\mu}^{r}\right|_{q}$ when $\delta \in \Delta$ if and only if $W(\boldsymbol{\delta}) \in l_{q}$ when $\delta \in \Delta$, which completes the first part of the proof of the theorem.
The remaining part of the proof is omitted, as it is similar.
Theorem 2.4. Assume that $1 \leq q<\infty$ and $\Delta$ is arbitrary sequence space. Then, $U \in\left(\left|E_{\mu}^{r}\right|_{q}, \Delta\right)$ if and only iffor all $n \geq 0$

$$
V^{(n)} \in\left(l_{q}, c\right) \text { and } \tilde{U}^{(q)} \in\left(l_{q}, \Delta\right)
$$

$U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, \Delta\right)$ if and only if for all $n \geq 0$

$$
V^{(n)} \in\left(l_{\infty}, c\right) \text { and } \tilde{U}^{(q)} \in\left(l_{\infty}, \Delta\right)
$$

Here the matrices $\tilde{U}$ and $V^{(n)}$ are described as

$$
\tilde{u}_{n k}^{(q)}=\mu_{k}^{-1 / p^{*}} \sum_{v=k}^{\infty}\binom{v-1}{k-1}(r-1)^{v-k} r^{-v} u_{n v}
$$

and

$$
v_{m k}^{(n)}=\left\{\begin{array}{lr}
u_{n 0}, & k=0 \\
\mu_{k}^{-1 / q^{*}} \sum_{v=k}^{m}\binom{v-1}{k-1}(r-1)^{v-k} r^{-v} u_{n v}, & 1 \leq k \leq m \\
0, & k>m .
\end{array}\right.
$$

Proof. We only demonstrate for $U \in\left(\left|E_{\mu}^{r}\right|_{q}, \Delta\right)$ to avoid repetition. Assume that $U \in\left(\left|E_{\mu}^{r}\right|_{q}, \Delta\right)$. Given $\delta \in\left|E_{\mu}^{r}\right|_{q}$. Since $\left|E_{\mu}^{r}\right|_{q}=$ $\left(l_{q}\right)_{T^{(r)}(q)}$, it follows from (1.4) that, for $n, m \geq 0$,

$$
\begin{equation*}
\sum_{k=0}^{m} u_{n k} \delta_{k}=\sum_{k=0}^{m} v_{m k}^{(n)} y_{k} . \tag{2.3}
\end{equation*}
$$

So, we get that, for all $\delta \in\left|E_{\mu}^{r}\right|_{q}, U \delta$ is well defined iff $V^{(n)} \in\left(l_{q}, c\right)$. Also, letting $m \rightarrow \infty$, gives (2.3) that $U \delta=\tilde{U}^{(q)} y$. Since $U \delta \in \Delta$, $\tilde{U}^{(q)} y$ is also in $\Delta$, and so $\tilde{U} \in\left(l_{q}, \Delta\right)$.
On the contrary, let $V^{(n)} \in\left(l_{q}, c\right)$ and $\tilde{U}^{(q)} \in\left(l_{q}, \Delta\right)$. Then, by (2.3), we have $U_{n} \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$ for all $n$, which gives that $U \delta$ exists. Also, by $\tilde{U}^{(q)} \in\left(l_{q}, \Delta\right)$ and (2.3), by letting $m \rightarrow \infty$, we get $U \in\left(\left|E_{\mu}^{r}\right|_{q}, \Delta\right)$.
We present the following tables and conditions:

| From To | c | $c_{0}$ | $l_{\infty}$ | 1 | $l_{p}(p>1)$ |
| :--- | :---: | ---: | ---: | ---: | ---: |
| $\left\|E_{\mu}^{r}\right\|_{q}$ | $1,3,12,14$ | $2,3,12,14$ | $3,12,14$ | $4,12,14$ | - |
| $\left\|E_{\mu}^{r}\right\|^{1,6,11,14}$ | $2,6,11,14$ | $6,11,14$ | $5,11,14$ | $5,11,14$ |  |
| $\left\|E_{\mu, q}^{r}\right\|_{\infty}$ | $1,7,13,14$ | $8,13,14$ | $10,13,14$ | $9,13,14$ | $9,13,14$ |

Table 1: From Absolute Euler spaces to $\left\{l_{\infty}, c_{0}, c, l, l_{p}\right\}$

| From To | $c_{s}$ | $b_{s}$ |
| :--- | :---: | :---: |
| $\left\|E_{\mu}^{r}\right\|_{q}$ | $1,3,12,14$ | $3,12,14$ |
| $\left\|E_{\mu}^{r}\right\|^{1,6,11,14}$ | $6,11,14$ |  |
| $\left\|E_{\mu, q}^{r}\right\|_{\infty}$ | $1,7,13,14$ | $10,13,14$ |

Table 2: From Absolute Euler spaces to $\left\{c_{s}, b_{s}\right\}$

1. $\lim _{n \rightarrow \infty} \tilde{u}_{n j}^{(q)}$ exists for all $j \in \mathbb{N}$
2. $\lim _{n \rightarrow \infty} \tilde{u}_{n j}^{(q)}=0$ for all $j \in \mathbb{N}$
3. $\sup _{n} \sum_{j=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}<\infty$
4. $\sup _{N} \sum_{v}\left|\sum_{n \in N} \tilde{n}_{n j}^{(q)}\right|^{q^{*}}<\infty$
5. $\sup _{j} \sum_{n}\left|\tilde{u}_{n j}^{(q)}\right|^{p}<\infty,(1 \leq p<\infty)$
6. $\sup _{n, j}\left|\tilde{u}_{n j}^{(q)}\right|<\infty$
7. $\sum_{j=0}^{\infty}\left|\tilde{u}_{n j}^{q}\right|<\infty$ uniformly in $n$
8. $\lim _{n} \sum_{j=0}^{\infty}\left|\tilde{u}_{n j}^{q}\right|=0$
9. $\sup _{K} \sum_{n=0}^{\infty}\left|\sum_{k \in K}^{\infty} \tilde{u}_{n j}^{q}\right|^{p}<\infty,(1 \leq p<\infty)$
10. $\sup _{n} \sum_{j=0}^{\infty}\left|\tilde{u}_{n j}^{q}\right|<\infty$
11. $\sup _{m, j}\left|v_{m j}^{(n)}\right|<\infty$
12. $\sup _{m} \sum_{j=0}^{\infty}\left|v_{m j}^{(n)}\right|^{q^{*}}<\infty$
13. $\sum_{j=0}^{\infty}\left|v_{m j}^{(n)}\right|<\infty$ uniformly in $m$
14. $\lim _{m \rightarrow \infty} v_{m j}^{(n)}$ exists for all $j, n \in \mathbb{N}$

We obtain following by Theorem 2.4.
Theorem 2.5. Let $1<p, q<\infty$. Then, Table 1 presents us the necessary and sufficient conditions for $U \in(\eta, \Lambda)$, where $\eta$ is one of absolute Euler spaces and $\Lambda \in\left\{c, c_{0}, l_{\infty}, l, l_{p}\right\}$.
Take the matrices $T_{1}=\left(t_{n j}^{1}\right)$ and $T_{2}=\left(t_{n j}^{2}\right)$ as

$$
t_{n j}^{1}=\left\{\begin{array}{c}
1,0 \leq j \leq n \\
0, \quad j>n
\end{array}\right.
$$

and

$$
t_{n j}^{2}=\left\{\begin{array}{c}
1, n=j \\
-1, n=j+1 \\
0, \text { otherwise } .
\end{array}\right.
$$

Then, since $b_{s}=\left\{l_{\infty}\right\}_{T_{1}}, c_{s}=\{c\}_{T_{1}}$ and $b v_{q}=\left\{l_{q}\right\}_{T_{2}}$, characterization of the matrix classes $(\eta, \Theta)$ can be obtained immediately as follows, where $\Theta \in\left\{c_{s}, b_{s}, b v_{q}\right\}$ and $\eta$ is one of the any absolute Euler spaces.

Corollary 2.6. Let's take $u(n, j)=\sum_{i=0}^{n} u_{i j}$ instead of $u_{n j}$ in the matrices $V^{(n)}=\left(v_{m v}^{(n)}\right)$ and $\tilde{U}^{(p)}=\left(\tilde{u}_{n v}^{(p)}\right)$ for all $n, j \geq 0$. Then, Table 2 presents us the necessary and sufficient conditions for $U \in(\eta, \Theta)$, where $\Theta \in\left\{c_{s}, b_{s}\right\}$ and $\eta$ is one of the absolute Euler spaces.
Corollary 2.7. Put $b_{n j}=u_{n j}-u_{n+1, j}$ instead of $u_{n j}$ in the matrices $V^{(n)}$ and $\tilde{U}^{(q)}$ for all $n, j \geq 0$. Then,

$$
\begin{aligned}
& U \in\left(\left|E_{\mu}^{r}\right|, b v_{p}\right) \text { iff the conditions 5, 11, } 14 \text { hold, } \\
& U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, b v_{p}\right) \text { iff the conditions 9,13, } 14 \text { hold. }
\end{aligned}
$$

Theorem 2.8. (i) Let $1<q<\infty$ and $\Lambda \in\left\{c_{0}, c, l_{\infty}\right\}$. Then,

$$
\begin{aligned}
& U \in\left(\left|E_{\mu}^{r}\right|_{q}, \Lambda\right) \Rightarrow\left\|S_{U}\right\|=\sup _{n}\left\|\tilde{U}_{n}^{(q)}\right\|_{l_{q^{*}}}=\sup _{n}\left(\sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}\right)^{1 / q^{*}} \\
& U \in\left(\left|E_{\mu}^{r}\right|, \Lambda\right) \Rightarrow\left\|S_{U}\right\|=\sup _{n}\left\|\tilde{U}_{n}^{(1)}\right\|_{l_{\infty}}=\sup _{n, v}\left|\tilde{u}_{n v}^{(1)}\right| \\
& U \in\left(\left|E_{\mu,}^{r}\right|_{\infty}, \Lambda\right) \Rightarrow\left\|S_{U}\right\|=\sup _{n}\left\|\tilde{U}_{n}^{(q)}\right\|_{l}=\sup _{n} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right| .
\end{aligned}
$$

(ii) Let $1<q<\infty$. Then, there exists $\zeta \in[1,4]$ such that

$$
\begin{aligned}
& U \in\left(\left|E_{\mu}^{r}\right|_{q}, l\right) \Rightarrow\left\|S_{U}\right\|=\frac{1}{\zeta}\left\|\tilde{U}^{(q)}\right\|_{\left(l_{q}, l\right)}^{\prime}=\frac{1}{\zeta}\left\{\sum_{v=0}^{\infty}\left(\sum_{n=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|\right)^{q^{*}}\right\}^{1 / q^{*}} \\
& U \in\left(\left|E_{\mu}^{r}\right|^{*}, l_{q}\right) \Rightarrow\left\|S_{U}\right\|=\left\|\tilde{U}^{(1)}\right\|_{\left(l, l_{q}\right)}=\sup _{v}\left\{\sum_{n=0}^{\infty}\left|\tilde{u}_{n v}^{(1)}\right|^{q}\right\}^{\frac{1}{q}}, \\
& U \in\left(\left|E_{\mu}^{r}\right|^{\prime}, l\right) \Rightarrow\left\|S_{U}\right\|=\left\|\tilde{U}_{n}^{(1)}\right\|_{(l, l)}=\sup _{v} \sum_{n=0}^{\infty}\left|\tilde{u}_{n v}^{(1)}\right| \\
& U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, l_{q}\right) \Rightarrow\left\|S_{U}\right\|=\left\|\tilde{U}^{(q)}\right\|_{\left(l_{\infty}, l_{q)}\right)} \\
& U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, l\right) \Rightarrow\left\|S_{U}\right\|=\left\|\tilde{U}^{(q)}\right\|_{(l, l)}
\end{aligned}
$$

Proof. The theorem can be easily proved by using Lemma 1.3, Lemma 1.5, Lemma 2.2 and Theorem 1.23 in [30], so it have left to reader.

Theorem 2.9. Let $1<q<\infty$.
(a) If $U \in\left(\left|E_{\mu}^{r}\right|_{q}, c_{0}\right)$, then

$$
\left\|S_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty} \sup _{n>r}\left\|\tilde{U}_{n}^{(q)}\right\|_{l_{q^{*}}}=\lim _{r \rightarrow \infty} \sup _{n>r}\left(\sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}\right)^{1 / q^{*}},
$$

and
$L_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}=0$.
(b) If $U \in\left(\left|E_{\mu}^{r}\right|_{q}, c\right)$, then

$$
\frac{1}{2} \lim _{r \rightarrow \infty} \sup _{n>r}\left(\sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right|^{q^{*}}\right)^{1 / q^{*}} \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r}\left(\sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right|^{q^{*}}\right)^{1 / q^{*}}
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right|^{q^{*}}=0$, where $\tilde{u}_{v}=\lim _{n \rightarrow \infty} \tilde{u}_{n v}$, for all $n \in \mathbb{N}$.
(c) If $U \in\left(\left|E_{\mu}^{r}\right|_{q}, l_{\infty}\right)$, then

$$
0 \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r}\left(\sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}\right)^{1 / q^{*}},
$$

and
if $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|^{q^{*}}=0, S_{U}$ is compact.

Proof. (a) Let $U \in\left(\left|E_{\mu}^{r}\right|_{q}, c_{0}\right)$. Then, the series $\sum_{n=0}^{\infty} u_{n v} \lambda_{v}$ converges for all $\lambda \in\left|E_{\mu}^{r}\right|_{q}$, or, equivalently $U_{n}=\left\{u_{n v}\right\}_{v=0}^{\infty} \in\left\{\left|E_{\mu}^{r}\right|_{q}\right\}^{\beta}$. So, it follows from Lemma 2.2 that $\left\|U_{n}\right\|_{\left.E_{\mu}^{r}\right|_{q}}^{*}=\left\|\tilde{U}_{n}\right\|_{l_{q^{*}}}$. Also, by Lemma 1.10 (a), we have

$$
\left\|S_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty} \sup _{n>r}\left\|\tilde{U}_{n}\right\|_{l_{q^{*}}} .
$$

Hence, the compactness of $S_{U}$ is immediate by Lemma 1.7, which completes the proof of (a).
(b) Let take the unit sphere $S_{\left|E_{\mu}^{r}\right|_{q}}^{\prime}$ in $\left|E_{\mu}^{r}\right|_{q}$. From Lemma 1.7 it follows that

$$
\left\|S_{U}\right\|_{\chi}=\chi\left(U\left(S_{\left|E_{\mu}^{r}\right|_{q}}^{\prime}\right)\right)
$$

Further, since $\left|E_{\mu}^{r}\right|_{q} \cong l_{q}, U \in\left(\left|E_{\mu}^{r}\right|_{q}, c\right)$ equals to $\tilde{U} \in\left(l_{q}, c\right)$, and

$$
\left\|S_{U}\right\|_{\chi}=\chi\left(U\left(S_{\left|E_{\mu}^{r}\right|_{q}}^{\prime}\right)\right)=\chi\left(\tilde{U}\left(T\left(S_{\left|E_{\mu}^{r}\right|_{q}}^{\prime}\right)\right)\right)=\left\|S_{\tilde{U}}\right\|_{\chi} .
$$

which implies, by Lemma 1.10 (b),

$$
\begin{equation*}
\frac{1}{2} \lim _{r \rightarrow \infty}\left(\sup _{n \geq r}\left\|\tilde{U}_{n}-\tilde{u}\right\|_{l_{q}}^{*}\right) \leq\left\|L_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty}\left(\sup _{n \geq r}\left\|\tilde{U}_{n}-\tilde{u}\right\|_{l_{q}}^{*}\right), \tag{2.4}
\end{equation*}
$$

where $\tilde{u}_{k}=\lim _{n \rightarrow \infty} \tilde{u}_{n k}$, for all $k \geq 0$.
Considering Theorem 1.29 in [30], it can be easily written that $\left\|\tilde{U}_{n}-\tilde{u}\right\|_{l_{q}}^{*}=\left\|\tilde{U}_{n}-\tilde{u}\right\|_{l_{p^{*}}}$. The last equality and (2.4) complete the first part of the proof of $(b)$. Also, the compactness of $S_{U}$ is concluded by Lemma 1.7.
(c) can be proved by similar way, so it is omitted.

By following the above lines, the proof of the following theorems also can be obtained immediately. Therefore, we just give the statement of the theorems.
Theorem 2.10. (a) If $U \in\left(\left|E_{\mu}^{r}\right|, c_{0}\right)$. Then

$$
\left\|S_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty} \sup _{n>r}\left\|\tilde{U}_{n}^{(1)}\right\|_{l_{\infty}}=\lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}\right|,
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{n}_{n v}^{(1)}\right|=0$.
(b) If $U \in\left(\left|E_{\mu}^{r}\right|, c\right)$, then

$$
\frac{1}{2} \lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}-\tilde{u}_{v}\right| \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}-\tilde{u}_{v}\right|
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}-\tilde{u}_{v}\right|=0$
where $\tilde{u}_{v}=\lim _{n \rightarrow \infty} \tilde{u}_{n v}$, for all $v \in \mathbb{N}$.
(c) If $U \in\left(\left|E_{\mu}^{r}\right|, l_{\infty}\right)$, then

$$
0 \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}\right|,
$$

and
$S_{U}$ is compact if $\lim _{r \rightarrow \infty} \sup _{n>r} \sup _{v}\left|\tilde{u}_{n v}^{(1)}\right|=0$.
Theorem 2.11. Let $1<q<\infty$.
(a) If $U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, c_{0}\right)$, then

$$
\left\|S_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty} \sup _{n>r}\left\|\tilde{U}_{n}^{(q)}\right\|_{l}=\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|,
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|=0$.
(b) If $U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, c\right)$, then

$$
\frac{1}{2} \lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right| \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right|
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}-\tilde{u}_{v}\right|=0$
where $\tilde{u}_{v}=\lim _{n \rightarrow \infty} \tilde{u}_{n v}$, for all $v \in \mathbb{N}$.
(c) If $U \in\left(\left|E_{\mu, q}^{r}\right|_{\infty}, l_{\infty}\right)$, then

$$
0 \leq\left\|S_{U}\right\|_{\chi} \leq \lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|,
$$

and
$S_{U}$ is compact if $\lim _{r \rightarrow \infty} \sup _{n>r} \sum_{v=0}^{\infty}\left|\tilde{u}_{n v}^{(q)}\right|=0$.
Theorem 2.12. (a) If $U \in\left(\left|E_{\mu}^{r}\right|, l_{q}\right), 1 \leq q<\infty$, then

$$
\left\|S_{U}\right\|_{\chi}=\lim _{r \rightarrow \infty}\left\{\sup _{v}\left(\sum_{n=r+1}^{\infty}\left|\tilde{u}_{n v}^{(1)}\right|^{q}\right)^{1 / q}\right\}
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sup _{v} \sum_{n=r+1}^{\infty}\left|\tilde{u}_{n v}^{(1)}\right|^{q}=0$.
(b) If $U \in\left(\left|E_{\mu}^{r}\right|_{q}, l\right), 1<q<\infty$, then there exists $\zeta \in[1,4]$ such that

$$
\left\|S_{U}\right\|_{\chi}=\frac{1}{\zeta} \lim _{r \rightarrow \infty}\left\{\sum_{v=0}^{\infty}\left(\sum_{n=r+1}^{\infty}\left|\tilde{u}_{n v}^{q)}\right|\right)^{q^{*}}\right\}^{1 / q},
$$

and
$S_{U}$ is compact iff $\lim _{r \rightarrow \infty} \sum_{v=0}^{\infty}\left(\sum_{n=r+1}\left|\tilde{u}_{n v}^{(q)}\right|\right)^{q^{*}}=0$.

## 3. Conclusion

One of the most important subjects in summability theory is the theory of sequence spaces which concerns with the generalization of the concept of convergence for series and sequences. In this sense, the primary aim is to assign a limit value for divergent sequences or series by using transformation which is given by the most general linear mappings of infinite special matrices. So, there has been a large literature, concerned with characterizing completely all matrices which transform one given sequence space into another. Besides this, the literature has been also grown up in terms of the studies of many sequence spaces defined as domain of special matrices and related matrix operators (see, for instance, [1-4, 6-12]). For a recent paper [1], the infinite matrix classes $\left(\left|E_{\mu}^{r}\right|,\left|E_{\mu}^{r}\right|_{q}\right)$ and $\left(\left|E_{\mu}^{r}\right|_{q},\left|E_{\mu}^{r}\right|\right)$ have been introduced. In this study, estimating the operator norms, the classes $\left(\left|E_{\mu}^{r}\right|_{q}, \Lambda\right)$ and $\left(\left|E_{\mu}^{r}\right|_{\infty}, \Lambda\right)$ have been characterized where $1 \leq q<\infty$. Also, in case $\Lambda$ is one of the spaces $c_{0}, c, l_{\infty}, l_{q}$, the necessary and sufficient conditions for each matrix in these classes to be compact have been obtained and certain identities or estimates for the Hausdorff measures of noncompactness have been established.

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# On Rough $\mathscr{I}$-Convergence and $\mathscr{I}$-Cauchy Sequence for Functions Defined on Amenable Semigroups 

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#### Abstract

In this paper, firstly we introduced the concepts of rough $\mathscr{I}$-convergence, rough $\mathscr{I}^{*}$ convergence, rough $\mathscr{I}$-Cauchy sequence, and rough $\mathscr{I}^{*}$-Cauchy sequence of a function defined on discrete countable amenable semigroups. Then, we investigated the relations between them.


## 1. Introduction

Throughout the paper, $\mathbb{N}$ denotes the set of all positive integers and $\mathbb{R}$ the set of all real numbers. The idea of $\mathscr{I}$-convergence was introduced by Kostyrko et al. [1] as a generalization of statistical convergence which is based on the structure of the ideal $\mathscr{I}$ of subset of $\mathbb{N}$.
Phu [2] introduced, firstly, the notion of rough convergence in finite-dimensional normed spaces. In [2], he investigated some properties of $\operatorname{LIM}^{r} x$ such as boundedness, closedness and convexity, and also he defined the notion of rough Cauchy sequence. Then, Phu [3] studied on rough convergence and some important properties of this concept. Furthermore, recently some authors [4-8] investigated the rough convergence types in some normed spaces.
In [9], Day studied on the concept of amenable semigroups (or briefly ASG). Then, some authors [10-12] studied the notions of summability in ASG. Douglas [13] extended the notion of arithmetic mean to ASG and obtained a characterization for almost convergence in ASG. In [14], Nuray and Rhoades presented the concepts of convergence and statistical convergence in ASG. Dündar et al. [15] and Dündar, Ulusu [16] introduced rough convergence and investigated some properties of rough convergence in ASG. Dündar, Ulusu [17] studied rough statistical convergence in ASG. Also, Dündar et al. [18] defined rough ideal convergence and some properties in ASG. Recently, some authors studied on the new concepts in ASG (see [19-22]).
First of all, we remember the basic definitions and concepts that we will use in our study such as amenable semigroups, rough convergence, rough ideal convergence, etc. (see [2,3,8-16, 18-24, 26, 27]).
Let a real number $r \geq 0$ and $\mathbb{R}^{n}$ (the real $n$-dimensional space) with the norm $\|\cdot\|$, and a sequence $x=\left(x_{k}\right)_{k=0}^{n} \subset \mathbb{R}^{n}$.
A sequence $\left(x_{k}\right)$ is said to be $r$-convergent to $L$, denoted by $x_{k} \xrightarrow{r} L$, provided that

$$
\forall \varepsilon>0 \exists k_{\varepsilon} \in \mathbb{N}: k \geq k_{\varepsilon} \Rightarrow\left\|x_{k}-L\right\|<r+\varepsilon .
$$

The rough limit set of the sequence $x=\left(x_{k}\right)$ is showed by $\operatorname{LIM}^{r} x=\left\{L \in \mathbb{R}^{n}: x_{k} \xrightarrow{r} L\right\}$.
A sequence $x=\left(x_{k}\right)$ is said to be $r$-convergent if $\operatorname{LIM}^{r} x \neq \emptyset$ and $r$ is called the convergence degree of the sequence $\left(x_{k}\right)$. For $r=0$, we get the ordinary convergence.
Let $G$ be a discrete countable amenable semigroups (or briefly DCASG) with identity in which both left and right cancelation laws hold, and $w(G)$ denotes the space of all real valued functions on $G$.
If $G$ is a countable amenable group, there exists a sequence $\left\{S_{n}\right\}$ of finite subsets of $G$ such that

[^2]
(i) $G=\bigcup_{n=1}^{\infty} S_{n}$,
(ii) $S_{n} \subset S_{n+1}(n=1,2, \ldots)$,
(iii) $\lim _{n \rightarrow \infty} \frac{\left|S_{n} g S_{n}\right|}{\left|S_{n}\right|}=1, \lim _{n \rightarrow \infty} \frac{\left|g S_{n} \cap S_{n}\right|}{\left|S_{n}\right|}=1$, for all $g \in G$.

If a sequence of finite subsets of $G$ satisfy (i)-(iii), then it is called a Folner sequence (or briefly FS) of $G$.
Throughout the paper, we take $G$ be a DCASG with identity in which both left and right cancelation laws hold.
For any FS $\left\{S_{n}\right\}$ of $G$, a function $f \in w(G)$ is said to be convergent to $t$ if for every $\varepsilon>0$ there exists a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that $|f(g)-t|<\varepsilon$, for all $m>k_{0}$ and $g \in G \backslash S_{m}$.
Let $X \neq \emptyset$. A class $\mathscr{I}$ of subsets of $X$ is said to be an ideal in $X$ provided:
i) $\emptyset \in \mathscr{I}$,
ii) $A, B \in \mathscr{I}$ implies $A \cup B \in \mathscr{I}$,
iii) $A \in \mathscr{I}, B \subset A$ implies $B \in \mathscr{I}$.
$\mathscr{I}$ is called a nontrivial ideal if $X \notin \mathscr{I}$. A nontrivial ideal $\mathscr{I}$ in $X$ is called admissible if $\{x\} \in \mathscr{I}$, for each $x \in X$.
Throughout the paper, we take $\mathscr{I}$ as an admissible ideal in $\mathbb{N}$.
Let $X \neq \emptyset$. A class $\emptyset \neq \mathscr{F}$ of subsets of $X$ is said to be a filter in $X$ provided:
i) $\emptyset \notin \mathscr{F}$,
ii) $A, B \in \mathscr{F}$ implies $A \cap B \in \mathscr{F}$,
iii) $A \in \mathscr{F}, A \subset B$ implies $B \in \mathscr{F}$.

If $\mathscr{I}$ is a nontrivial ideal in $X, X \neq \emptyset$, then the class

$$
\mathscr{F}(\mathscr{I})=\{M \subset X:(\exists A \in \mathscr{I})(M=X \backslash A)\}
$$

is a filter on $X$, called the filter associated with $\mathscr{I}$.
An admissible ideal $\mathscr{I} \subset 2^{\mathbb{N}}$ satisfies the property $(A P)$, if for every countable family of mutually disjoint sets $\left\{A_{1}, A_{2}, \ldots\right\}$ belonging to $\mathscr{I}$, there exists a countable family of sets $\left\{B_{1}, B_{2}, \ldots\right\}$ such that $A_{j} \Delta B_{j}$ is a finite set for $j \in \mathbb{N}$ and $B=\bigcup_{j=1}^{\infty} B_{j} \in \mathscr{I}$ (hence $B_{j} \in \mathscr{I}$ for each $j \in \mathbb{N}$ ).
After then, we let $\mathscr{I} \subseteq 2^{G}$ be an admissible ideal for amenable semigroup $G$.
A function $f \in w(G)$ is said to be $\mathscr{I}$-convergent to $s$ for any FS $\left\{S_{n}\right\}$ for $G$, if for every $\varepsilon>0$

$$
\{g \in G:|f(g)-s| \geq \varepsilon\} \in \mathscr{I} .
$$

In this case, we write $\mathscr{I}-\lim f(g)=s$.
A function $f \in w(G)$ is said to be $\mathscr{I}^{*}$-convergent to $s$, for any FS $\left\{S_{n}\right\}$ for $G$ if there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $\left.G \backslash M \in \mathscr{I}\right)$ and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0|f(g)-s|<\varepsilon$, for all $n>k_{0}$ and all $g \in M \backslash S_{n}$. In this case, we write $\mathscr{I}^{*}-\lim f(g)=s$.
A function $f \in w(G)$ is said to be $\mathscr{I}$-Cauchy sequence, for any $\operatorname{FS}\left\{S_{n}\right\}$ for $G$ if for every $\varepsilon>0$, there exists an $h=h(\varepsilon) \in G$ such that

$$
\{g \in G:|f(g)-f(h)| \geq \varepsilon\} \in \mathscr{I} .
$$

A function $f \in w(G)$ is said to be $\mathscr{I}^{*}$-Cauchy sequence, for any FS $\left\{S_{n}\right\}$ for $G$ if there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $G \backslash M \in \mathscr{I}$ ) and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0|f(g)-f(h)|<\varepsilon$, for all $n>k_{0}$ and $g, h \in M \backslash S_{n}$.
For any FS $\left\{S_{n}\right\}$ of $G$, a function $f \in w(G)$ is said to be rough convergent ( $r$-convergent) to $t$ if

$$
\begin{equation*}
\forall \varepsilon>0 \exists k_{\varepsilon} \in \mathbb{N}: m \geq k_{\varepsilon} \Rightarrow|f(g)-t|<r+\varepsilon, \tag{1.1}
\end{equation*}
$$

for all $g \in G \backslash S_{m}$ or equivalently if limsup $|f(g)-t| \leq r$, for all $g \in G \backslash S_{m}$. In this instance, we write $r-\lim f(g)=t$ or $f(g) \xrightarrow{r} t$. If (1.1) holds, then $t$ is an $r$-limit point of the function $f \in w(G)$, which is usually no longer unique (for $r>0$ ). Hence, we have to think the so-called rough limit set ( $r$-limit set) of the function $f \in w(G)$ defined by $\operatorname{LIM}^{r} f:=\{t: f(g) \xrightarrow{r} t\}$.
For any FS $\left\{S_{n}\right\}$ for $G$, the function $f \in w(G)$ is said to be $r$-convergent if $\operatorname{LIM}^{r} f \neq \emptyset$. In this instance, $r$ is called the convergence degree of the $f \in w(G)$.
For any FS $\left\{S_{n}\right\}$ of $G$, a function $f \in w(G)$ is said to be a rough Cauchy sequence with roughness degree $\wp$, if $\forall \varepsilon>0 \exists k_{\varepsilon}: m \geq k_{\varepsilon} \Rightarrow$ $|f(g)-f(h)| \leq \wp+\varepsilon$ is hold for $\wp>0$ and all $g, h \in G \backslash S_{m}$. $\wp$ is also said to be Cauchy degree of $f \in w(G)$.

## 2. Main Results

In this section, we introduced the concepts of rough $\mathscr{I}$-convergence, rough $\mathscr{I}^{*}$-convergence, rough $\mathscr{I}$-Cauchy sequence and rough $\mathscr{I}^{*}$-Cauchy sequence of a function defined on discrete countable amenable semigroups. Then, we investigated relations between them.
Definition 2.1. For any $F S\left\{S_{n}\right\}$ of $G$, a function $f \in w(G)$ is said to be rough $\mathscr{I}$-convergent ( $r$ - $\mathscr{I}$-convergent) to $s$ iffor every $\varepsilon>0$

$$
\begin{equation*}
\{g \in G:|f(g)-s| \geq r+\varepsilon\} \in \mathscr{I} \tag{2.1}
\end{equation*}
$$

or equivalently if

$$
\mathscr{I}-\limsup |f(g)-s| \leq r
$$

is satisfied. In this instance, we write

$$
r-\mathscr{I}-\lim f(g)=s \text { or } f(g) \xrightarrow{r-\mathscr{g}} s .
$$

On the other hand, we say that $f(g) \xrightarrow{r-\mathscr{G}} s$ if and only if the condition

$$
|f(g)-s| \leq r+\varepsilon
$$

holds for every $\varepsilon>0$ and almost $g \in G$.

In this convergence $r$ is named the roughness degree. For $r=0$, we get the $\mathscr{I}$-convergence.
If (2.1) holds, then $s$ is an $r-\mathscr{I}$-limit point of the function $f \in w(G)$, which is usually no longer unique (for $r>0$ ). Hence, we have to think the so-called rough $\mathscr{I}$-limit set of the function $f \in w(G)$ defined by

$$
\mathscr{I}-\operatorname{LIM}^{r} f:=\{s: f(g) \xrightarrow{r-\mathscr{g}} s\} .
$$

For any FS $\left\{S_{n}\right\}$ for $G$, the function $f \in w(G)$ is said to be $r$ - $\mathscr{I}$-convergent if

$$
\mathscr{I}-\operatorname{LIM}^{r} f \neq \emptyset .
$$

If $\mathscr{I}-\operatorname{LIM}^{r} f \neq \emptyset$ for a function $f \in w(G)$, then we have

$$
\mathscr{I}-\operatorname{LIM}^{r} f=[\mathscr{I}-\limsup f-r, \mathscr{I}-\liminf f+r] .
$$

Remark 2.2. If $\mathscr{I}$ is an admissible ideal, then for a function $f \in w(G)$, usual rough convergence implies rough $\mathscr{I}$-convergence for any $F S$ $\left\{S_{n}\right\}$ of $G$.

Definition 2.3. A function $f \in w(G)$ is said to be rough $\mathscr{I}$-Cauchy sequence, for any $F S\left\{S_{n}\right\}$ for $G$ iffor every $\varepsilon>0$, there exists an $h=h(\varepsilon) \in G$ such that

$$
\{g \in G:|f(g)-f(h)| \geq r+\varepsilon\} \in \mathscr{I} .
$$

Theorem 2.4. If $f \in w(G)$ is rough $\mathscr{I}$-convergent for any $F S\left\{S_{n}\right\}$ for $G$, then it is rough $\mathscr{I}$-Cauchy for same sequence.
Proof. For any Folner sequence $\left\{S_{n}\right\}$ for $G$, let

$$
r-\mathscr{I}-\lim f(g)=s .
$$

Then, for every $\varepsilon>0$, we have

$$
A_{\varepsilon}=\{g \in G:|f(g)-s| \geq r+\varepsilon\} \in \mathscr{I} .
$$

Since $\mathscr{I}$ is an admissible ideal there exists an $h \in G$ such that $h \notin A_{\mathcal{\varepsilon}}$. Now, let

$$
B_{\varepsilon}=\{g \in G:|f(g)-f(h)| \geq 2(r+\varepsilon)\} .
$$

Taking into account the inequality

$$
|f(g)-f(h)| \leq|f(g)-s|+|f(h)-s|,
$$

we observe that if $g \in B_{\varepsilon}$, then

$$
|f(g)-s|+|f(h)-s| \geq 2(r+\varepsilon) .
$$

On the other hand, since $h \notin A_{\varepsilon}$ we have

$$
|f(h)-s|<r+\varepsilon
$$

and so

$$
|f(g)-s|>r+\varepsilon .
$$

Hence, $g \in A_{\varepsilon}$ and so we have

$$
B_{\varepsilon} \subset A_{\varepsilon} \in \mathscr{I} .
$$

Thus, $B_{\varepsilon} \in \mathscr{I}$ that is, $f$ is rough $\mathscr{I}$-Cauchy sequence.
Definition 2.5. A function $f \in w(G)$ is said to be rough $\mathscr{I}^{*}$-convergent to $s$, for any $F S\left\{S_{n}\right\}$ for $G$ if there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $G \backslash M \in \mathscr{I})$ and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0$

$$
\begin{equation*}
|f(g)-s|<r+\varepsilon, \tag{2.2}
\end{equation*}
$$

for all $n>k_{0}$ and all $g \in M \backslash S_{n}$. In this case, we write

$$
r-\mathscr{I}^{*}-\lim f(g)=s .
$$

In this convergence $r$ is named the roughness degree. For $r=0$, we get the $\mathscr{I}^{*}$-convergence.
If (2.2) holds, then $s$ is an $r-\mathscr{I}^{*}$-limit point of the function $f \in w(G)$, which is usually no longer unique (for $r>0$ ).
Hence, we have to think the so-called rough $\mathscr{I}^{*}$-limit set of the function $f \in w(G)$ defined by

$$
\mathscr{I}^{*}-\operatorname{LIM}^{r} f:=\left\{s: f(g) \xrightarrow{r-\mathscr{I}^{*}} s\right\} .
$$

For any FS $\left\{S_{n}\right\}$ for $G$, the function $f \in w(G)$ is said to be $r-\mathscr{I}^{*}$-convergent if

$$
\mathscr{I}^{*}-\operatorname{LIM}^{r} f \neq \emptyset .
$$

Theorem 2.6. If $f \in w(G)$ is rough $\mathscr{I}^{*}$-convergent to $s$, then $f$ is rough $\mathscr{I}$-convergent to $s$ for any $F S\left\{S_{n}\right\}$ for $G$.
Proof. For any FS $\left\{S_{n}\right\}$ for $G$, let

$$
r-\mathscr{I}^{*}-\lim f(g)=s
$$

Then, there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $H=G \backslash M \in \mathscr{I})$ and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0$

$$
|f(g)-s|<r+\varepsilon,
$$

for all $n>k_{0}$ and all $g \in M \backslash S_{n}$. Therefore obviously,

$$
A(\varepsilon)=\{g \in G:|f(g)-s| \geq r+\varepsilon\} \subset H \cup S_{k_{0}} .
$$

Since $\mathscr{I}$ is admissible,

$$
H \cup S_{k_{0}} \in \mathscr{I}
$$

and so

$$
A(\varepsilon) \in \mathscr{I} .
$$

Hence,

$$
r-\mathscr{I}-\lim f(g)=s .
$$

Definition 2.7. A function $f \in w(G)$ is said to be rough $\mathscr{I}^{*}$-Cauchy sequence, for any $F S\left\{S_{n}\right\}$ for $G$ if there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $G \backslash M \in \mathscr{I}$ ) and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0$

$$
|f(g)-f(h)|<r+\varepsilon
$$

for all $n>k_{0}$ and $g, h \in M \backslash S_{n}$.
Theorem 2.8. If $f \in w(G)$ is rough $\mathscr{I}^{*}$-Cauchy for any $F S\left\{S_{n}\right\}$ for $G$, then it is rough $\mathscr{I}$-Cauchy for same sequence.
Proof. Let $f \in w(G)$ be an rough $\mathscr{I}^{*}$-Cauchy for any FS $\left\{S_{n}\right\}$ for $G$. Then by definition, there exists $M \subset G, M \in \mathscr{F}(\mathscr{I})$ (i.e., $\left.G \backslash M \in \mathscr{I}\right)$ and a $k_{0}=k_{0}(\varepsilon) \in \mathbb{N}$ such that for every $\varepsilon>0$

$$
|f(g)-f(h)|<r+\varepsilon,
$$

for all $n>k_{0}$ and $g, h \in M \backslash S_{n}$. Let $H=G \backslash M$. It is clearly $H \in \mathscr{I}$ and

$$
A(\varepsilon)=\{g \in G:|f(g)-f(h)| \geq r+\varepsilon\} \subset H \cup S_{k_{0}}
$$

Since $\mathscr{I}$ is admissible,

$$
H \cup S_{k_{0}} \in \mathscr{I}
$$

and so

$$
A(\varepsilon) \in \mathscr{I} .
$$

Consequently, $f$ is rough $\mathscr{I}$-Cauchy for same sequence.
Following theorems show relationships between $\mathscr{I}$-convergence and $\mathscr{I}^{*}$-convergence, between $\mathscr{I}$-Cauchy sequence and $\mathscr{I}^{*}$-Cauchy sequence. These theorems can be proved like in [19,25], these theorems are given without the proof.
Theorem 2.9. Let $\mathscr{I} \subset 2^{G}$ be an admissible ideal with the property $(A P)$. If $f(g) \in w(G)$ is rough $\mathscr{I}$-convergent to $s$, then $f$ is rough $\mathscr{I}^{*}$-convergent to sfor any $F S\left\{S_{n}\right\}$ for $G$.
Theorem 2.10. Let $\mathscr{I} \subset 2^{G}$ be an admissible ideal with the property $(A P)$. If $f \in w(G)$ is rough $\mathscr{I}$-Cauchy for any $F S\left\{S_{n}\right\}$ for $G$, then it is rough $\mathscr{I}^{*}$-Cauchy for same sequence.

## 3. Conclusion

In this paper, we introduced the concepts of rough $\mathscr{I}$-convergence, rough $\mathscr{I}^{*}$-convergence, rough $\mathscr{I}$-Cauchy sequence and rough $\mathscr{I}^{*}$ Cauchy sequence of a function defined on discrete countable amenable semigroups. Also, we investigated relations between them. Then after, The concepts given here can also be studied for double sequences.

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