RESEARCH ARTICLE

Monitoring the Period Variation of KR Cyg Eclipsing Binary

M. Geldi^{1*} , K. Çalışkan¹, and E. Sipahi¹

ABSTRACT

This study examines the period variation of the eclipsing binary system KR Cyg, a near-contact binary characterized by its short orbital period and significant stellar interactions. The precise orbital parameters of the system have been determined through light curve data collected periodically since 1999 and were published in our previous studies. Additionally, minimum light observations have been continuously monitored to analyze the period variation of the system. In our earlier work, the masses of the primary and secondary components were determined as $2.88 \pm 0.20 M_{\odot}$ and $1.26 \pm 0.07 M_{\odot}$, with corresponding radii of $2.59 \pm 0.06 R_{\odot}$ and $1.80 \pm 0.04R_{\odot}$. The bolometric albedo and effective temperature of the less massive star were also investigated, and deviations caused by mutual illumination effects were identified. These results provide valuable insights into the evolutionary status of KR Cyg and shed light on the dynamic processes that occur in near-contact binary systems. Using updated eclipse timings, an (O-C) diagram was constructed, revealing both long-term trends and periodic oscillations. These cyclical variations are thought to indicate the presence of a tertiary component affecting the system via the light-time effect. Additional observations and modeling are recommended to confirm the tertiary hypothesis and further refine the parameters of the system.

Keywords: Techniques: CCD photometry; stars: KR Cyg; period variation

1. INTRODUCTION

Near-contact binary systems (NCBs) represent a fascinating class of stellar binaries where at least one component is close to filling its Roche lobe, often leading to significant tidal interactions and mutual distortion. These systems are typically characterized by short orbital periods and exhibit complex light curve variations due to effects such as ellipsoidal modulation, reflection, and sometimes partial eclipses. NCBs serve as crucial laboratories for studying stellar evolution, mass transfer processes, and the influence of proximity on stellar structure and behavior. Their diverse configurations often position them as precursors to more evolved systems, such as contact binaries or cataclysmic variables. Understanding the physical and dynamical properties of NCBs provides valuable insights into binary evolution scenarios and their role in broader astrophysical contexts.

KR Cyg is an EB-type eclipsing binary system initially identified by Schneller (1931) and Schneller (1932) in the early 1930s. Visual light curves were later documented by Lassovszky (1936), Gaposchkin (1953), and Tsesevich (1954). Photographic light curves were produced by Wachmann (1948) and Nekrasova (1945), whose studies indicated that KR Cyg is an Algol-type binary. Subsequent observations by Vetesnik (1965) introduced the first photoelectric light curves for the system, determining an orbital period of 0.8451538 days. Based on these light curves, Vetešnik categorized the system as β Lyrae-type, comprising a B9-type primary star and an F5-type secondary star. However, Koch (1973) assigned the secondary star a spectral type of G4, while Hill et al. (1975) conducted the first spectroscopic analysis, suggesting the system is closer to type A3. In 1980, Wilson & Rafert (1980) analyzed Vetešnik's light curves and calculated the first photometric solution, yielding a mass ratio of q = 0.478. A statistical analysis by Giuricin et al. (1983) grouped KR Cyg among 'a' class Algol systems, which includes classical semi-detached binaries and Algols with subgiant secondaries. Al-Naimiy et al. (1985) applied Fourier analysis to derive the geometric and physical properties of the system, reporting a slightly adjusted photometric mass ratio of q = 0.51. Later, Kholopov (1985) classified the system as A2V. In 2004, Sipahi & Gülmen (2004) presented three-colour photometric data, further refining the system's parameters. Their WD-based analysis revealed a photometric mass ratio of q = 0.43, suggesting that the cooler, less massive secondary nearly fills its Roche lobe. Shaw (1990) classified KR Cyg as a near-contact binary (NCB), a group where both stars either fill or nearly fill their Roche lobes. At a later time, Sipahi (2012) compiled all known light minimum times and analysed the system's orbital period. Her findings revealed periodic oscillations in the (O-C) data, with an amplitude of 0.001 days and a cycle lasting approximately 76 years,

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Table 1. Basic parameters of the observed stars.

Star	α_{2025}	δ_{2025}	V	Spt Typ
	(hh:mm:ss.ss)	(dd:mm:ss.ss)	(mag)	
KR Cyg	20:10:05.83	+30:37:29.60	9.34	A0
HD 191398	20:09:39.73	+30:24:42.96	9.00	A0V
HD 333664	20:10:18.54	+30:18:07.56	9.63	A0

which she attributed to a hypothetical third component. Shortly after Sipahi et al. (2013) presented multi-colour, five-year photometric light curves and radial velocity measurements for the near-contact binary system KR Cyg.

In this study, the component masses were determined as $2.88 \pm 0.20 M_{\odot}$ and $1.26 \pm 0.07 M_{\odot}$, with corresponding radii of $2.59 \pm 0.06R_{\odot}$ and $1.80 \pm 0.04R_{\odot}$. Additionally, they investigated the empirical determination of the albedo and effective temperature of the cooler, less massive star in KR Cyg, alongside a comparison with two similar near-contact binaries, AK CMi and DO Cas. Discrepancies between the observed and computed fluxes are attributed to mutual illumination effects, where the heated surface layers of the illuminated star experience variations in bolometric albedo, limb-darkening coefficients, and gravity-brightening exponents. Interestingly, the derived effective albedos are generally lower than those predicted for stars with convective envelopes. It is important to note that these findings are preliminary and warrant further investigation. Tvardovskyi & Marsakova (2015) have also published a study on the period variation of the system. In this study, the light curve variations and period changes of the KR Cyg eclipsing binary system were analysed using the photometric observations from 2024 and the updated minimum time. This study aims to investigate the presence of the third component in the system through (O-C) analysis using approximately 95 years of minimum time data.

2. OBSERVATIONS

Observations were acquired with a thermoelectrically cooled ALTA U +47 1024 × 1024 pixel CCD camera attached to a 40 cm Schmidt-Cassegrains type MEADE telescope at Ege University Observatory. The observations made in the *V* band were continued on September 27, 28, and 29, in the season of 2024. In Table 1, the coordinates, apparent visual magnitudes, and spectral types of the variable and the comparison stars are given. Although the variable and comparison stars are very close in the sky, differential atmospheric extinction corrections were applied. The atmospheric extinction coefficients were obtained from observations of the comparison stars on each night. Heliocentric corrections were also applied to the times of the observations. The mean averages of the standard deviations are 0.025 mag for observations acquired in the V band. To compute the standard deviations of observations, we used the standard deviations of the reduced differential magnitudes in the sense comparisons minus the check stars for each night. All KR Cyg data was phased using the minimum time and period taken from Sipahi (2012). Then, the V-band light curve shown in Figure 1 was derived.

KR Cyg eclipsing binary has been observed at different inter-

vals over 26 years at the Ege University Observatory. Discussions on the light variations of the KR Cyg system obtained in different years have been provided by Sipahi & Gülmen (2004), Sipahi & Tas (2006), Sipahi (2012), and Sipahi et al. (2013). Observations obtained in the *V* band in 2024 demonstrated that the system's light curves exhibit similar variations. The depths of the minima do not change. Both minima are sufficiently symmetrical. The secondary minimum lies just at the phase 0.5. Since the radial velocity of the system, the solution of the light curve, and the absolute parameters are provided in Sipahi et al. (2013), the parameters of that study have been adopted in this work.

3. ORBITAL PERIOD VARIATION

To investigate the orbital period variation of KR Cyg, all accessible minimum times were gathered from the literature. In addition, the system was observed for three nights in 2024. From these observations, a primary minimum time was determined. Additionally, the primary and secondary minimum times of the system have been determined from the Transiting Exoplanet Survey Satellite (*TESS*) data (Ricker et al. 2015). All other minimum times were obtained from the database of (O-C) Gateway (Paschke & Brat 2006). All minimum timings of KR Cyg eclipsing binary are provided in the appendix. In total, 179 visual, 29 photographic, and 219 photoelectric minimum times of KR Cyg were used in the period analysis. The (O-C) residuals of KR Cyg were calculated using the following linear ephemeris and plotted in Figure 2.

$$JD$$
 (Hel.) = 2455036.5310(14) + 0^d.84515279(6) × E. (1)

Subsequently, the (O-C)I values and (O-C)II residuals were plotted against the epoch number (*E*) in Figure 2 and Figure 3, respectively. The residuals (O-C) II suggest that the orbital period of KR Cyg exhibits a long-term sinusoidal variation. It is hypothesized that the most plausible explanation for the cyclic variation in the (O-C)II residuals is the light-travel-time effect (LTTE) caused by an undetected third body in the system. To model this effect, the LTTE equation, presented as expression (2) by Irwin (1959), was applied to the (O-C)II residuals of the system's eclipse timings.

$$\Delta t = \frac{a_{12} \sin i'}{c} \left\{ \frac{1 - e'^2}{1 + e' \cos \nu'} \sin(\nu' + \omega') + e' \cos \omega' \right\}. \tag{2}$$

Here, Δt represents the time shift caused by the LTTE, c is the speed of light, and a_{12} , i', e' and ω' denote the semi-major axis, inclination, eccentricity, and longitude of periastron of the absolute orbit of the eclipsing binary's center of mass around the triple system's center of mass. Additionally, ν' refers to the true anomaly of the eclipsing binary's center of mass along this orbit

A weighted least-squares analysis was performed to determine two parameters (T_0 and P) for the linear ephemeris of KR Cyg, along with five parameters ($a_{12} \sin i'$, e', ω' , T' and P') for the LTTE. The results of this analysis are summarized in Table 2. The observed data points, along with the theoretical curve providing the best fit, are plotted against the epoch number in Figure 3. Assuming that the orbit of the presumed third body is

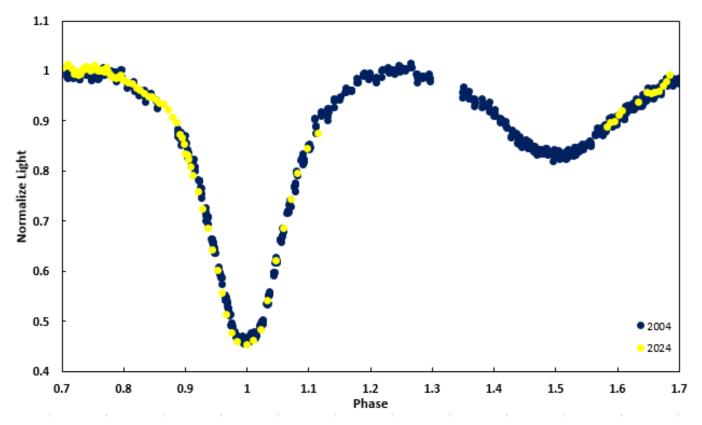


Figure 1. The *V*-band normalized light curves of the KR Cyg from 2004 (blue-coloured points) and 2024 (yellow-coloured points) are presented (brightness values were converted to intensity to compare the light curve variations over a 20-year time interval).

Table 2. The parameters derived from (O-C) analysis of KR Cyg.

Parameter	Value
T_0 (HJD)	2455036.5310±0.0014
P (day)	0.84515279±0.00000006
K (day)	0.012±0.006
P' (year)	70.9±1.9
e'	0.93 ± 0.12
ω' (deg)	26.70±0.02
$a_{12}\sin i'$ (AU)	1.99±0.22
$f(m) (M_{\odot})$	0.00157±0.00010
$M_3 (i'=15^{\circ})$	$1.41 M_{\odot}$
$M_3 (i'=30^{\circ})$	$0.66 M_{\odot}$
$M_3 (i'=45^{\circ})$	$0.45 M_{\odot}$
$M_3 (i'=60^{\circ})$	$0.36M_{\odot}$
$M_3 (i'=75^{\circ})$	$0.33M_{\odot}$
$M_3 \ (i'=90^\circ)$	$0.32M_{\odot}$

circular, we obtained the mass function as $f(m) = 0.00157 M_{\odot}$ for the third body, using the Equation:

$$f(m) = \frac{4\pi^2}{GT^2} \times (a_{12}\sin i')^3 = \frac{(M_3 \sin i')^3}{(M_1 + M_2 + M_3)^2}.$$
 (3)

where M_1 , M_2 , and M_3 are the masses of the binary's components and the third body, respectively.

The mass of the third component can be determined using Equation (3), which varies based on the orbital inclination. For instance, when $i' = 90^{\circ}$, the minimum mass $M_{3,\text{min}}$ is calculated

to be $0.32~M_{\odot}$. In this calculation, the total mass of the eclipsing binary system is assumed to be $4.14~M_{\odot}$. The masses of the third body for various inclination angles i' have been computed and are presented in Table 3. The data in Table 3 also indicate that KR Cyg follows an eccentric orbit around the center of mass of the triple system, with a period of approximately ~ 71 years.

4. CONCLUSION

In this study, the orbital period variation of the KR Cyg system was reanalysed using data based on its 2024 photometry. The system has been observed at Ege University observatory during certain years and monitored for approximately 26 years. During this time, photometric and spectroscopic studies of the system were obtained, and the absolute parameters of the components were determined by Sipahi et al. (2013). The orbital period variation caused by a third body was proposed by Sipahi (2012). All minimum time observations conducted after the period variation study proposed by Sipahi (2012) have supported this suggestion. The (O-C) variation of the KR Cyg appears scattered. Both the primary and secondary minima exhibit the same variation in the (O-C) diagram. With our 2024 observations, the system's period variation has been updated under the assumption that the orbital period change is caused by a third body. Tvardovskyi & Marsakova (2015) also examined the period variation of the KR Cyg system and determined the orbital period of the third body to be approximately 80 years. The minimum timings to be obtained in the next approximately five years are crucial for refining the orbital parameters of the third body of the system.

Table 3. Times of minima of KR Cyg.

HID (+24 00000)	Vaan	Cvala	(0,0)			HID (+24 00000)		Cyclo	(0,0)	Mathad	Deferences
HJD (+24 00000)	Year	Cycle	(O-C)	Method	References	HJD (+24 00000)	Year	Cycle	(O-C)	Method	References
25700.4230	1929.3	-34711.0	-0.04594	pg	O-C Gateway	41234.3540	1971.9	-16331.0	-0.00319	vis	O-C Gateway
26120.4860	1930.5	-34214.0	-0.02334	pg	O-C Gateway	41490.4170	1972.6	-16028.0	-0.02115	vis	O-C Gateway
26298.7740	1931.0	-34003.0	-0.06234	pg	O-C Gateway	41490.4230	1972.6	-16028.0	-0.01515	vis	O-C Gateway
26447.5310 26586.1400	1931.4 1931.8	-33827.0 -33663.0	-0.05204 -0.04792	pg vis	O-C Gateway O-C Gateway	41506.4740 41511.5550	1972.6 1972.6	-16009.0 -16003.0	-0.02203 -0.01194	vis vis	O-C Gateway O-C Gateway
26590.3700	1931.8	-33658.0	-0.04792	vis	O-C Gateway	41522.5430	1972.0	-15990.0	-0.01194	vis	O-C Gateway O-C Gateway
26591.2090	1931.8	-33657.0	-0.04308	vis	O-C Gateway	41528.4524	1972.7	-15983.0	-0.01092	vis	O-C Gateway
26596.2800	1931.8	-33651.0	-0.04983	vis	O-C Gateway	41556.3550	1972.7	-15950.0	-0.01738	vis	O-C Gateway
26601.3550	1931.8	-33645.0	-0.04565	vis	O-C Gateway	41583.3900	1972.8	-15918.0	-0.01484	vis	O-C Gateway
26602.1910	1931.8	-33644.0	-0.05481	vis	O-C Gateway	41589.3080	1972.8	-15911.0	-0.01290	vis	O-C Gateway
26607.2700	1931.8	-33638.0	-0.04672	vis	O-C Gateway	41594.4080	1972.8	-15905.0	0.01619	vis	O-C Gateway
27245.3740	1933.6	-32883.0	-0.03225	vis	O-C Gateway	41599.4410	1972.9	-15899.0	-0.02172	vis	O-C Gateway
27553.0000	1934.4	-32519.0	-0.04147	V	O-C Gateway	41622.2820	1972.9	-15872.0	0.00018	vis	O-C Gateway
28074.4660	1935.8	-31902.0	-0.03407	pg	O-C Gateway	41627.3310	1972.9	-15866.0	-0.02173	В	O-C Gateway
28333.9050	1936.5	-31595.0	-0.05664	V	O-C Gateway	41627.3330	1972.9	-15866.0	-0.01973	vis	O-C Gateway
28380.4170	1936.7	-31540.0	-0.02798	pg	O-C Gateway	41812.4120	1973.4	-15647.0	-0.02895	vis	O-C Gateway
28545.2100	1937.1	-31345.0	-0.03956	pg	O-C Gateway	41823.4030	1973.5	-15634.0	-0.02492	vis	O-C Gateway
28773.3910	1937.7	-31075.0	-0.04952	pg	O-C Gateway	41850.4440	1973.5	-15602.0	-0.02878	vis	O-C Gateway
28806.3640	1937.8	-31036.0	-0.03744	pg	O-C Gateway	41850.4510	1973.5	-15602.0	-0.02178	vis	O-C Gateway
28811.4430	1937.8	-31030.0	-0.02935	pg	O-C Gateway	41871.5850	1973.6	-15577.0	-0.01657	vis	O-C Gateway
29106.4027	1938.7	-30681.0	-0.02759	pg	O-C Gateway	41904.5430	1973.7	-15538.0	-0.01949 -0.02534	vis	O-C Gateway
29106.4130 29143.5780	1938.7 1938.8	-30681.0 -30637.0	-0.01729 -0.03897	pg	O-C Gateway	41931.5820	1973.8 1973.8	-15506.0 -15505.0	-0.02534	vis	O-C Gateway
29230.6040	1938.8	-30534.0	-0.05897	pg ng	O-C Gateway O-C Gateway	41932.4330 41938.3480	1973.8	-15303.0	-0.01949	vis vis	O-C Gateway O-C Gateway
30259.2040	1939.0	-30334.0	-0.00339	pg pg	O-C Gateway	41938.3510	1973.8	-15498.0	-0.02033	vis	O-C Gateway
30531.3270	1941.6	-28995.0	-0.01321	pg pg	O-C Gateway	41943.4220	1973.8	-15492.0	-0.01733	vis	O-C Gateway
32296.8470	1947.4	-26906.0	-0.02000	pg pg	O-C Gateway	42144.5560	1974.4	-15254.0	-0.01740	vis	O-C Gateway
32821.6840	1948.8	-26285.0	-0.03317	pg	O-C Gateway	42193.5810	1974.5	-15196.0	-0.02337	vis	O-C Gateway
33220.6060	1949.9	-25813.0	-0.02277	pg	O-C Gateway	42221.4741	1974.6	-15163.0	-0.02027	vis	O-C Gateway
33226.5160	1949.9	-25806.0	-0.02883	pg	O-C Gateway	42254.4490	1974.7	-15124.0	-0.00629	vis	O-C Gateway
33536.6850	1950.8	-25439.0	-0.03050	pg	O-C Gateway	42337.2410	1974.9	-15026.0	-0.03916	vis	O-C Gateway
33567.9700	1950.9	-25402.0	-0.01612	pg	O-C Gateway	42419.2410	1975.1	-14929.0	-0.01887	vis	O-C Gateway
33923.7660	1951.8	-24981.0	-0.02898	pg	O-C Gateway	42570.5200	1975.5	-14750.0	-0.02203	vis	O-C Gateway
34213.6610	1952.6	-24638.0	-0.02102	pg	O-C Gateway	42576.4480	1975.5	-14743.0	-0.01009	vis	O-C Gateway
34512.8460	1953.5	-24284.0	-0.01972	pe	O-C Gateway	42708.2930	1975.9	-14587.0	-0.00875	vis	O-C Gateway
34596.5280	1953.7	-24185.0	-0.00774	vis	O-C Gateway	42958.4560	1976.6	-14291.0	-0.01066	vis	O-C Gateway
34602.4390	1953.7 1953.7	-24178.0	-0.01280 -0.00771	vis	O-C Gateway	42980.4260	1976.6	-14265.0 -14246.0	-0.01460	vis	O-C Gateway O-C Gateway
34607.5150 34650.6050	1953.7	-24172.0 -24121.0	-0.00771	vis	O-C Gateway O-C Gateway	42996.4850 43013.3900	1976.7 1976.7	-14246.0	-0.01348 -0.01152	vis vis	O-C Gateway O-C Gateway
35221.9170	1955.4	-23445.0	-0.02044	pg	O-C Gateway	43046.3560	1976.7	-14220.0	-0.001132	vis	O-C Gateway
35774.6500	1956.9	-22791.0	-0.02721	pg pg	O-C Gateway	43280.4500	1977.5	-13910.0	-0.01945	vis	O-C Gateway
36818.4010	1959.8	-21556.0	-0.03855	pg pg	O-C Gateway	43673.4540	1978.5	-13445.0	-0.01099	vis	O-C Gateway
36868.3020	1959.9	-21497.0	-0.00151	pg	O-C Gateway	43689.5150	1978.6	-13426.0	-0.00788	vis	O-C Gateway
38336.3120	1963.9	-19760.0	-0.02001	vis	O-C Gateway	43706.4260	1978.6	-13406.0	0.00009	vis	O-C Gateway
38558.5900	1964.5	-19497.0	-0.01691	pe	O-C Gateway	43739.3680	1978.7	-13367.0	-0.01883	vis	O-C Gateway
38559.4352	1964.5	-19496.0	-0.01686	pe	O-C Gateway	43744.4350	1978.7	-13361.0	-0.02274	vis	O-C Gateway
38559.4360	1964.5	-19496.0	-0.01606	pe	O-C Gateway	43766.4210	1978.8	-13335.0	-0.01068	vis	O-C Gateway
38580.5649	1964.6	-19471.0	-0.01595	pe	O-C Gateway	43777.4150	1978.8	-13322.0	-0.00365	vis	O-C Gateway
38580.5657	1964.6	-19471.0	-0.01515	pe	O-C Gateway	44165.3320	1979.9	-12863.0	-0.01128	vis	O-C Gateway
38597.4671	1964.6	-19451.0	-0.01678	pe	O-C Gateway	44426.4810	1980.6	-12554.0	-0.01416	vis	O-C Gateway
38608.4535	1964.7	-19438.0	-0.01736	pe	O-C Gateway	44437.4880	1980.6	-12541.0	0.00587	vis	O-C Gateway
38614.3693	1964.7	-19431.0	-0.01762	pe	O-C Gateway	44443.3700	1980.6 1980.7	-12534.0	-0.02819	vis vis	O-C Gateway
38614.3702 38652.4050	1964.7 1964.8	-19431.0 -19386.0	-0.01672 -0.01374	pe vis	O-C Gateway O-C Gateway	44454.3700 44498.3490	1980.7	-12521.0 -12469.0	-0.01516 0.01595	vis	O-C Gateway O-C Gateway
38675.2210	1964.9	-19359.0	-0.01574		O-C Gateway	44503.3870	1980.8	-12463.0	-0.01696	vis	O-C Gateway
38675.2229	1964.9	-19359.0	-0.01084	pe pe	O-C Gateway	44514.3870	1980.8	-12450.0	-0.01090	vis	O-C Gateway
38941.4520	1965.6	-19044.0	-0.00863	vis	O-C Gateway	44569.2970	1981.0	-12385.0	-0.00394	vis	O-C Gateway
39040.3170	1965.9	-18927.0	-0.02637	vis	O-C Gateway	44569.2990	1981.0	-12385.0	-0.02680	vis	O-C Gateway
39040.3300	1965.9	-18927.0	-0.01337	vis	O-C Gateway	44913.2800	1981.9	-11978.0	-0.02254	vis	O-C Gateway
39051.3050	1965.9	-18914.0	-0.02535	vis	O-C Gateway	44913.2920	1981.9	-11978.0	-0.01054	vis	O-C Gateway
39051.3060	1965.9	-18914.0	-0.02435	vis	O-C Gateway	45130.4907	1982.5	-11721.0	-0.01582	vis	O-C Gateway
39389.3770	1966.8	-18514.0	-0.01403	vis	O-C Gateway	45600.3920	1983.8	-11165.0	-0.01887	vis	O-C Gateway
40420.4578	1969.6	-17294.0	-0.01830	vis	O-C Gateway	45644.3340	1983.9	-11113.0	-0.02476	vis	O-C Gateway
40725.5590	1970.5	-16933.0	-0.01686	vis	O-C Gateway	45878.4480	1984.6	-10836.0	-0.01778	vis	O-C Gateway
40731.4690	1970.5	-16926.0	-0.02293	vis	O-C Gateway	45889.4330	1984.6	-10823.0	-0.01975	vis	O-C Gateway
40753.4590	1970.5	-16900.0	-0.00687	vis	O-C Gateway	45911.3980	1984.7	-10797.0	-0.02870	vis	O-C Gateway
40759.3740	1970.6	-16893.0	-0.00793	vis	O-C Gateway	45933.3840	1984.7	-10771.0	-0.01664	vis	O-C Gateway
40780.5000 40785.5700	1970.6 1970.6	-16868.0 -16862.0	-0.01072 -0.01163	vis vis	O-C Gateway O-C Gateway	46004.3820 46271.4480	1984.9 1985.7	-10687.0 -10371.0	-0.01138 -0.01332	vis vis	O-C Gateway O-C Gateway
40786.4100	1970.6	-16861.0	-0.01163	vis	O-C Gateway	46271.4480	1985.7	-10371.0	-0.01332	vis	O-C Gateway
40791.4800	1970.6	-16855.0	-0.01679	vis	O-C Gateway	46271.4390	1985.7	-10371.0	-0.00232	vis	O-C Gateway
40796.5540	1970.7	-16849.0	-0.01770	vis	O-C Gateway	46331.4520	1985.8	-10300.0	-0.01518	vis	O-C Gateway
40830.3670	1970.8	-16809.0	-0.00767	vis	O-C Gateway	46917.5890	1987.4	-9606.5	0.00921	pe	O-C Gateway
40830.3860	1970.8	-16809.0	0.01133	vis	O-C Gateway	46992.3660	1987.6	-9518.0	-0.00972	vis	O-C Gateway
40841.3460	1970.8	-16796.0	-0.01565	pe	O-C Gateway	47003.3520	1987.7	-9505.0	-0.01069	vis	O-C Gateway
40890.3550	1970.9	-16738.0	-0.02545	vis	O-C Gateway	47008.4070	1987.7	-9499.0	-0.02660	vis	O-C Gateway
41080.5150	1971.4	-16513.0	-0.02458	vis	O-C Gateway	47023.6300	1987.7	-9481.0	-0.01633	vis	O-C Gateway
41135.4540	1971.6	-16448.0	-0.02044	vis	O-C Gateway	47039.6920	1987.8	-9462.0	-0.01221	vis	O-C Gateway
41146.4560	1971.6	-16435.0	-0.00541	vis	O-C Gateway	47055.3250	1987.8	-9443.5	-0.01452	vis	O-C Gateway
41162.4970	1971.7	-16416.0	-0.02229	vis	O-C Gateway	47083.6410	1987.9	-9410.0	-0.01110	vis	O-C Gateway
41168.4130	1971.7	-16409.0	-0.02235	vis	O-C Gateway	47111.5250	1988.0	-9377.0	-0.01711	vis	O-C Gateway
					0.00		1000 0	0265.0	0.01202		
41173.4930	1971.7	-16403.0	-0.01326	vis	O-C Gateway	47121.6700	1988.0	-9365.0	-0.01393	vis	O-C Gateway
	1971.7 1971.7 1971.8	-16403.0 -16402.0 -16370.0	-0.01326 -0.01912 -0.00527	vis pe vis	O-C Gateway O-C Gateway O-C Gateway	47121.6700 47151.2420 47300.8430	1988.1 1988.5	-9305.0 -9330.0 -9153.0	-0.01393 -0.02224 -0.01309	vis vis vis	O-C Gateway O-C Gateway O-C Gateway

 Table 3. Continue

Table 3. Continue											
HJD (+24 00000)	Year	Cycle	(O-C)	Method	References	HJD (+24 00000)	Year	Cycle	(O-C)	Method	References
47362.5360	1988.6	-9080.0	-0.01616	pe	O-C Gateway	51691.4093	2000.5	-3958.0	-0.00987	BV	Sipahi & Gülmen (2004)
47368.4310		-9073.0		vis	O-C Gateway	51718.4518			-0.01223		Sipahi & Gülmen (2004)
47374.3630		-9066.0	-0.02129		O-C Gateway	51721.4147			-0.00736		Sipahi & Gülmen (2004)
		-9066.0 -9053.0	-0.00229 0.00874		O-C Gateway O-C Gateway	51726.4804 51737.4680	2000.6		-0.01257 -0.01194		Sipahi & Gülmen (2004) Sipahi & Gülmen (2004)
		-9033.0	-0.00520		O-C Gateway	51797.4700	2000.8		-0.01194	CCD	O-C Gateway
47410.7090					O-C Gateway	51816.4875			-0.01412		O-C Gateway
47412.3997	1988.8	-9021.0	-0.01641	vis	O-C Gateway	51850.2932	2000.9	-3770.0	-0.01449	CCD	O-C Gateway
				vis	O-C Gateway	52411.4782			-0.01022		O-C Gateway
			-0.02364		O-C Gateway	52427.5330			-0.01330		O-C Gateway
47460.5750 47462.2650			-0.01476 -0.01506	vis vis	O-C Gateway O-C Gateway	52531.4839 52576.2812	2002.8		-0.01606 -0.01180		O-C Gateway O-C Gateway
			-0.01300		O-C Gateway	52612.6234			-0.01110		O-C Gateway
47522.2880		-8891.0	0.00216		O-C Gateway	52613.0512	2003.0		-0.00590		O-C Gateway
	1989.6	-8666.0	-0.01244		O-C Gateway	52815.4619	2003.6	-2628.0	-0.00903	V	O-C Gateway
47721.7200					O-C Gateway	52832.3635			-0.01047		O-C Gateway
47734.4230		-8640.0	0.00409	vis	O-C Gateway	52840.3829			-0.02001		O-C Gateway
47805.3980 48012.4530		-8556.0 -8311.0	-0.01365	vis vis	O-C Gateway O-C Gateway	52859.4074 52864.4790	2003.7		-0.01142 -0.01073	pe CCD	O-C Gateway O-C Gateway
48043.7340				vis	O-C Gateway	52899.1337	2003.7			V	O-C Gateway O-C Gateway
				vis	O-C Gateway	52951.5306	2003.9	-2467.0	-0.00976		O-C Gateway
			-0.01045	vis	O-C Gateway	52955.3517	2004.0	-2462.5	0.00816	CCD	O-C Gateway
48143.4660			-0.00633	vis	O-C Gateway	53192.4000			-0.00859	pe	O-C Gateway
					O-C Gateway	53195.3598		-2178.5	-0.00682		O-C Gateway
48180.6430 48191.6310			-0.01601	vis vis	O-C Gateway O-C Gateway	53212.6832 53225.3614			-0.00903 -0.00811	pe	O-C Gateway O-C Gateway
48202.6110			-0.02195	vis	O-C Gateway	53267.6169			-0.01019	CCD	O-C Gateway
48208.5410				vis	O-C Gateway	53269.3084					O-C Gateway
48235.5820				vis	O-C Gateway	53289.5913	2004.9			CCD	O-C Gateway
48469.6810			-0.01989	vis	O-C Gateway	53572.7160			-0.01086	CCD	O-C Gateway
48491.6460					O-C Gateway	53585.3955			-0.00863		O-C Gateway
48507.7090 48535.6160		-7725.0 -7692.0	-0.02372 -0.00672	vis vis	O-C Gateway O-C Gateway	53593.4249 53601.4563			-0.00817 -0.00571	-Ir	O-C Gateway O-C Gateway
			-0.00072		O-C Gateway	53639.4865			-0.00371		O-C Gateway O-C Gateway
48820.4400		-7355.0	0.00115	vis	O-C Gateway	53891.3416				uvby	O-C Gateway
48853.3960	1992.7	-7316.0	-0.00376	vis	O-C Gateway	53907.3985	2006.6	-1336.0	-0.00843	pe	O-C Gateway
48983.5410					O-C Gateway	53937.4060				uvby	O-C Gateway
49164.4000			-0.01559	vis	O-C Gateway	53978.3920			-0.00767	BV	O-C Gateway
49195.6680 49238.7720			-0.01820 -0.01694		O-C Gateway O-C Gateway	53984.3076 53991.4921			-0.00813 -0.00742		O-C Gateway O-C Gateway
49266.6650			-0.01094	vis	O-C Gateway	54198.5565	2000.8	-991.5		BVRI	O-C Gateway O-C Gateway
49271.7350			-0.01485	vis	O-C Gateway	54270.8169	2007.6	-906.0		CCD	O-C Gateway
49562.4500			-0.03204	vis	O-C Gateway	54282.2214	2007.6	-892.5	-0.01031	V	O-C Gateway
49687.5530			-0.01149	vis	O-C Gateway	54313.4927	2007.7	-855.5	-0.00962		O-C Gateway
49926.7130 49948.6950			-0.02942 -0.02137	vis	O-C Gateway O-C Gateway	54338.4286 54614.7946	2007.7 2008.5	-826.0	-0.00570 -0.00430		O-C Gateway O-C Gateway
49953.7740				vis	O-C Gateway	54653.6721	2008.5		-0.00430		O-C Gateway
49958.8400	1995.8		-0.01819		O-C Gateway	54682.4095	2008.7	-419.0		BVR	O-C Gateway
	1995.8		-0.02340		O-C Gateway	54696.3480	2008.7	-402.5	-0.00804	В	O-C Gateway
50226.7610			-0.01028	pe	O-C Gateway	54702.6904	2008.7		-0.00428		O-C Gateway
50266.4764				CCD	O-C Gateway	54762.2706	2008.9	-324.5	-0.00727		O-C Gateway
50308.7380 50325.6340					O-C Gateway O-C Gateway	55018.7782 55036.5296			-0.00321 0.00000		O-C Gateway O-C Gateway
50336.6320					O-C Gateway	55058.4980			-0.00554		O-C Gateway
50402.5540					O-C Gateway	55058.9206	2009.7		-0.00552		O-C Gateway
50542.8470					O-C Gateway	55059.3432		27.0	-0.00550	pe	O-C Gateway
50553.8430			-0.00198	vis	O-C Gateway	55059.7658	2009.7		-0.00547	pe	O-C Gateway
50668.7750 50690.7400				vis	O-C Gateway O-C Gateway	55060.1884	2009.7		-0.00545 -0.00552	pe	O-C Gateway O-C Gateway
50690.7400			-0.01956	vis vis	O-C Gateway O-C Gateway	55060.6109 55061.0335	2009.7		-0.00552 -0.00550	pe pe	O-C Gateway O-C Gateway
					O-C Gateway	55061.4561			-0.00548	pe	O-C Gateway
50700.4609	1997.8	-5130.5	-0.01790	CCD	O-C Gateway	55061.8787	2009.7	30.0	-0.00545	pe	O-C Gateway
			-0.01383	CCD	O-C Gateway	55062.3012	2009.7		-0.00553	pe	O-C Gateway
50750.3193					O-C Gateway	55062.7238					O-C Gateway
50753.2828 50755.3856					O-C Gateway O-C Gateway	55063.1464 55063.5690			-0.00548 -0.00545	pe pe	O-C Gateway O-C Gateway
50772.3013					O-C Gateway	55075.4040				-Ir	O-C Gateway
50948.5166					O-C Gateway	55096.5329			-0.00247		O-C Gateway
51007.6870	1998.6	-4767.0	-0.00445	vis	O-C Gateway	55362.7581			-0.00006		O-C Gateway
51033.4436					O-C Gateway	55379.6589		406.0	-0.00229		O-C Gateway
51036.4121					O-C Gateway	55386.4356		414.0	0.01320		Sipahi (2012)
51040.6410 51083.7520			-0.01136 -0.00310		O-C Gateway O-C Gateway	55397.4081 55419.3832	2010.6		-0.00128 -0.00012	-Ir BV	O-C Gateway Sipahi (2012)
51133.6100					O-C Gateway	55461.6389			-0.00201		O-C Gateway
51325.4531				CCD	O-C Gateway	55465.0120			-0.00951		O-C Gateway
			-0.01371		Sipahi & Gülmen (2004)	55691.5220		775.0	-0.00017	V	Sipahi (2012)
51391.3746					O-C Gateway	55801.3912			-0.00069		O-C Gateway
51393.4932			-0.01000 -0.01264		O-C Gateway	56147.0599 56158.4718		1314.0	0.00097		O-C Gateway
51429.4095 51434.4798					Sipahi & Gülmen (2004) O-C Gateway	56158.4718 56186.3665		1327.5 1360.5	0.00332 0.00801		O-C Gateway O-C Gateway
51443.3488			-0.01323		O-C Gateway	56483.4310		1712.0	0.00301		O-C Gateway O-C Gateway
51454.3437	1999.8	-4238.5	-0.01042	BV	Sipahi & Gülmen (2004)	56487.6559	2013.6	1717.0	0.00083		O-C Gateway
51459.4121					O-C Gateway	56494.4181	2013.6	1725.0	0.00182	-I	O-C Gateway
51468.2868	1999.9	-4222.0	-0.01232	CCD	O-C Gateway	56496.5367	2013.7	1727.5	0.00754	-I	O-C Gateway

Table 3. Continue

HJD (+24 00000)	Year	Cycle	(O-C)	Method	References	HJD (+24 00000)	Year	Cycle	(O-C)	Method	References
56535.4043	2013.8	1773.5	-0.00184	-I	O-C Gateway	59776.5744(3)	2022.6	5608.5	0.01151		TESS
56810.5034	2014.5	2099.0	0.00038	-I	O-C Gateway	59779.5318(2)	2022.6	5612.0	0.01083		TESS
56924.5950	2014.8	2234.0	-0.00350	vis	O-C Gateway	59779.9543(4)	2022.6	5612.5	0.01083		TESS
56924.5997	2014.8	2234.0	0.00120	V	O-C Gateway	59780.3776(2)	2022.6	5613.0	0.01152		TESS
57198.4295	2015.6	2558.0	0.00185	-I	O-C Gateway	59780.8015(7)	2022.6	5613.5	0.01286		TESS
57201.8101	2015.6	2562.0	0.00184	V	O-C Gateway	59781.2233(3)	2022.6	5614.0	0.01206		TESS
57206.4575	2015.6	2567.5	0.00095	Clear	O-C Gateway	59781.6430(12)	2022.6	5614.5	0.00919		TESS
57214.4880	2015.6	2577.0	0.00247	-I	O-C Gateway	59782.0673(3)	2022.6	5615.0	0.01087		TESS
57219.5606	2015.6	2583.0	0.00416	-I	O-C Gateway	59783.7602(9)	2022.6	5617.0	0.01348		TESS
57225.4742	2015.6	2590.0	0.00170	-I	O-C Gateway	59784.1754(7)	2022.6	5617.5	0.00613		TESS
57260.5559	2015.7	2631.5	0.00960	-Ir	O-C Gateway	59784.6029(4)	2022.6	5618.0	0.01105		TESS
57267.7314	2015.8	2640.0	0.00131	V	O-C Gateway	59785.0251(2)	2022.6	5618.5	0.01072		TESS
57574.5243	2016.6	3003.0	0.00414	-I	O-C Gateway	59785.4475	2022.7	5619.0	0.01054	V	O-C Gateway
57657.3476	2016.8	3101.0	0.00254	Clear	O-C Gateway	59785.4519(16)	2022.6	5619.0	0.01486	•	TESS
57924.4177	2017.6	3417.0	0.00474	-Ir	O-C Gateway	59785.8711(3)	2022.7	5619.5	0.01151		TESS
57926.5294	2017.6	3419.5	0.00356	-Ir	O-C Gateway	59786.2934(2)	2022.7	5620.0	0.01128		TESS
57995.4110	2017.8	3501.0	0.00530	CCD	O-C Gateway	59786.7153(10)	2022.7	5620.5	0.01061		TESS
58265.8611	2018.5	3821.0	0.00685	V	O-C Gateway	59787.1391(3)	2022.7	5621.0	0.01183		TESS
58306.4286	2018.6	3869.0	0.00711	Ř	O-C Gateway	59787.5603(6)	2022.7	5621.5	0.01103		TESS
58327.5577	2018.7	3894.0	0.00738	-Ir	O-C Gateway	59787.9834(2)	2022.7	5622.0	0.01017		TESS
58344.4606	2018.7	3914.0	0.00730	BVRI	O-C Gateway	59788.4073(10)	2022.7	5622.5	0.01030		TESS
58414.6086	2018.9	3997.0	0.00722	V	O-C Gateway	59788.8293(3)	2022.7	5623.0	0.01225		TESS
58749.2881	2019.8	4393.0	0.00711	v	O-C Gateway	59789.2493(8)	2022.7	5623.5	0.00916		TESS
58749.2884	2019.8	4393.0	0.00711	v	O-C Gateway	59789.6756(3)	2022.7	5624.0	0.01284		TESS
58749.2884	2019.8	4393.0	0.00737	v	O-C Gateway	59790.0968(1)	2022.7	5624.5	0.01204		TESS
59035.7977	2020.6	4732.0	0.01026	v	O-C Gateway	59790.5193(17)	2022.7	5625.0	0.01137		TESS
59083.1224	2020.7	4788.0	0.00646	v	O-C Gateway	59792.2097(4)	2022.7	5627.0	0.01137		TESS
59337.5131	2021.4	5089.0	0.00647	v	O-C Gateway	59792.6323(4)	2022.7	5627.5	0.01115		TESS
59337.5148	2021.4	5089.0	0.00815	v	O-C Gateway	59793.0543(1)	2022.7	5628.0	0.01130		TESS
59381.4637	2021.5	5141.0	0.00921	-Ir	O-C Gateway	59793.4782(7)	2022.7	5628.5	0.01032		TESS
59770.2357(3)	2022.6	5601.0	0.00521	-11	TESS	59793.8999(2)	2022.7	5629.0	0.01222		TESS
59770.6608(10)	2022.6	5601.5	0.01140		TESS	59794.3236(7)	2022.7	5629.5	0.01155		TESS
59771.0807(2)	2022.6	5602.0	0.01355		TESS	59794.7446(3)	2022.7	5630.0	0.01233		TESS
59771.5033(6)	2022.6	5602.5	0.01123		TESS	59795.1693(8)	2022.7	5630.5	0.01301		TESS
59771.9257(2)	2022.6	5603.0	0.01130		TESS	59795.5905(2)	2022.7	5631.0	0.01361		TESS
59772.3482(4)	2022.6	5603.5	0.01113		TESS	59796.0130(1)	2022.7	5631.5	0.01162		TESS
59772.7705(2)	2022.6	5604.0	0.01101		TESS	59812.4930	2022.7	5651.0	0.01102	V	O-C Gateway
59773.1928(6)	2022.6	5604.5	0.01076		TESS	59839.5376	2022.7	5683.0	0.01114	V	O-C Gateway
59773.6174(3)	2022.6	5605.0	0.01040		TESS	60195.3471	2022.8	6104.0	0.01069	В	O-C Gateway
59774.0410(9)	2022.6	5605.5	0.01247		TESS	60195.3471	2023.8	6104.0	0.01158	В	O-C Gateway
59774.4619(2)	2022.6	5606.0	0.01332		TESS	60195.3472	2023.8	6104.0	0.01158	В	O-C Gateway
59774.8823(8)	2022.6	5606.5	0.01188		TESS	60195.3475	2023.8	6104.0	0.01100	В	O-C Gateway
59775.3012(21)	2022.6	5607.0	0.00907		TESS	60536.3619	2023.8	6507.5	0.01193	В	O-C Gateway
59775.7303(10)	2022.6	5607.5	0.00603		TESS TESS	60583.2694(2)	2024.7	6563.0	0.00763	V V	In this study
59776.1535(15)	2022.6	5608.0	0.01259		TESS TESS	00363.2094(2)	2024.8	0.606.0	0.00919	V	in this study
39110.1333(13)	2022.0	3006.0	0.01515		1 E33						

The latest data obtained for the eclipsing binary KR Cyg indicate that there is no significant change in the light curves. Detailed analyses of the characteristics of the light curves of the system have been provided by Sipahi (2012). According to these authors, comparison of the radii of the components with their corresponding Roche lobes suggests that the KR Cyg system is near-contact but not in contact. By combining the results of light curve and radial velocity analyses, they determined the absolute physical parameters of the system. Both components are located on the main sequence of the Hertzsprung-Russell diagram. The systematic behaviour of the residuals between the observed and fitted light curves is attributed to the mutual heating effect of the components, particularly in the case of the cooler star in this near-contact system. Analyses indicate that the effective albedo and effective temperature of the irradiated star are significantly altered. The empirically derived albedo value is smaller than expected for a convective star. As the albedo decreases, the effective temperature increases.

Near-contact binary systems are of significant interest in astrophysics due to their role in understanding mass transfer, angular momentum evolution, and stellar mergers. Just like KR Cyg eclipsing binary system, V994 Her, RZ Dra, and CN And illustrate the diversity among NCB systems with a third component. In V994 Her eclipsing binary system, the presence of a

tertiary component has been confirmed through spectroscopic analysis, revealing a hierarchical triple system (Zasche et al. 2023). The third body of the system influences the orbital evolution of the binary and contributes to its mass-transfer dynamics. Similarly, RZ Dra, a well-studied eclipsing binary, shows evidence of long-term orbital period variations, attributed to a distant third companion (Erdem et al. 2011). Meanwhile, CN And stands out due to its compact tertiary component, which significantly perturbs the inner binary's orbit, creating detectable eclipsing light curve variations (Cai et al. 2019). These systems provide critical insights into the dynamical stability of multi-star systems and the mechanisms driving their evolution. Observations of such binaries with a third component challenge existing models of stellar evolution, particularly regarding angular momentum redistribution and the eventual fate of such systems, which may evolve into contact binaries or merge into single stars.

From these investigations, we could draw out the following conclusions:

1. The orbital period variation of the system is caused by a third body. According to the results of the period variation analysis, the orbital period of the third body has been updated in \sim 71 years.

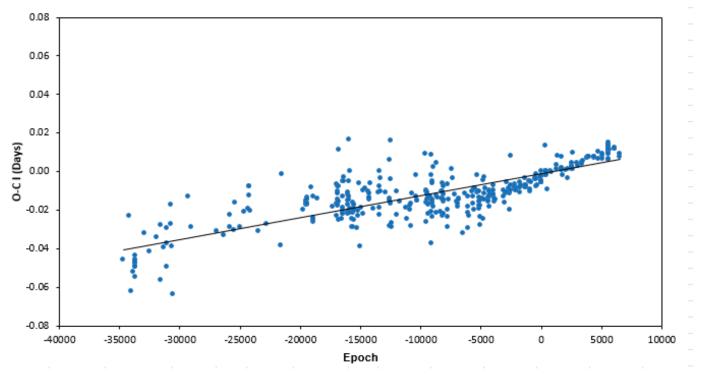


Figure 2. (O-C) diagram for KR Cyg.

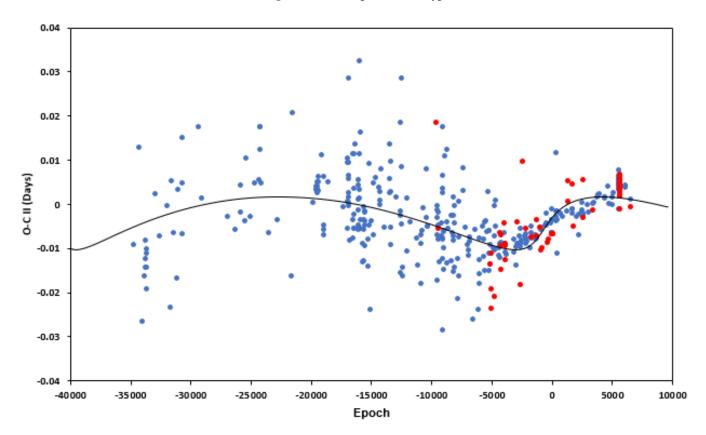


Figure 3. Residuals (O-C II) for KR Cyg. The filled blue and red circles represent the primary and secondary minima, respectively.

- 2. No significant changes have been observed in the light curves of the system over approximately 26 years of observations.
- 3. High-resolution spectroscopic observations are required to test the existence of the third body in the system.
- 4. Observations of minimum timings to be obtained in the coming years are crucial for determining the orbital parameters of the third body more precisely. Observations should continue.

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