

HOMOGENEOUS NON-STATIC SOLUTIONS OF EINSTEIN-MAXWELL FIELD-EQUATIONS FOR AN ISOTROPIC SPACE-TIME

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We consider solutions of the EINSTEIN-MAXWELL field-equations in vacuo for spatially homogeneous electro-magnetic fields. We find that, except for certain times, the metric is regular everywhere, there is no magnetic-field and there is an electric-field of equal strength in the x , y and z directions which increases with time.

1. **Introduction.** The EINSTEIN-MAXWELL field-equations in vacuo are,

$$(1.1) \quad G_{\mu}^{\nu} = 8\pi (F_{\mu\sigma} F^{\nu\sigma} + \frac{1}{4} g_{\mu}^{\nu} F_{\sigma\alpha} F^{\sigma\alpha}),$$

$$(1.2) \quad F_{\mu\nu, \sigma} + F_{\nu\sigma, \mu} + F_{\sigma\mu, \nu} = 0,$$

$$(1.3) \quad (\sqrt{-g} F^{\mu\nu})_{, \nu} = 0,$$

where G_{μ}^{ν} is the EINSTEIN tensor

$$(1.4) \quad G_{\mu}^{\nu} \text{ def } R_{\mu}^{\nu} - \frac{1}{2} g_{\mu}^{\nu} R$$

where R_{μ}^{ν} is the RICCI curvature tensor

$$(1.5) \quad R_{\mu\nu} = \Gamma_{\mu\tau}^{\sigma} \Gamma_{\sigma\nu}^{\tau} - \Gamma_{\mu\nu}^{\sigma} \Gamma_{\sigma\tau}^{\tau} + (\Gamma_{\mu\tau}^{\tau})_{, \nu} - (\Gamma_{\mu\nu}^{\tau})_{, \tau}$$

and

$$(1.6) \quad R \text{ def } R_{\mu\nu} g^{\mu\nu}.$$

$g^{\mu\nu}$ is the fundamental tensor and $\Gamma_{\mu\nu}^{\sigma}$ are the CHRISTOFFEL 3-index symbols. $F_{\mu\alpha}$ is the skew-symmetric electromagnetic field tensor.

The RAINICH conditions for the existence of a non-null electro-magnetic field are [1]:

a) The MAXWELL tensor has zero trace,

$$(1.7) \quad R = 0.$$

b) The square of this tensor is proportional to the unit matrix,

$$(1.8) \quad R_{\mu}^{\sigma} R_{\sigma}^{\nu} = \rho^2 \delta_{\mu}^{\nu}$$

c) The electro-magnetic energy-density is positive definite,

$$(1.9) \quad R_{00} > 0.$$

d) A certain form of the RICCI tensor shall have zero curl,

$$(1.10) \quad \frac{g_{\mu\sigma} \epsilon^{\alpha\beta\gamma\lambda} R_{\alpha}^{\nu} R_{\beta\gamma, \lambda}}{4 e^2 \sqrt{-g}} = a, \mu = 0$$

where $\epsilon^{\alpha\beta\gamma\lambda}$ is the LEVI-CIVITA permutation symbol and $\epsilon^{0123} = 1$.

When the EINSTEIN-MAXWELL equations hold, these RAINICH conditions are completely equivalent to EINSTEIN's description of electro-magnetic radiation and gravitation.

Further, for an electro-magnetic field, the eigen values of the EINSTEIN tensor reduce to the form,

$$(1.11) \quad \varrho, \varrho, -\varrho, -\varrho.$$

2. **Field-Equations.** We consider the isotropic line-element, [2],

$$(2.1) \quad ds^2 = e^{-2\Phi} dt^2 - e^{2\Phi} (dx^2 + dy^2 + dz^2)$$

where we take Φ to be a function of t alone.

The components of the EINSTEIN tensor G_{μ}^{ν} for the space-time (2.1) are found to be,

$$(2.2) \quad \begin{aligned} G_0^0 &= 3 e^{2\Phi} \dot{\Phi}^2 \\ G_1^1 &= G_2^2 = G_3^3 = e^{2\Phi} (5 \dot{\Phi}^2 + 2 \ddot{\Phi}) \end{aligned}$$

where the dot denotes differentiation with respect to time.

From (2.2) it is seen that the RAINICH conditions (1.7) to (1.10) are satisfied.

The eigen-values of the EINSTEIN tensor are given by

$$(2.3) \quad G_{\mu}^{\alpha} G_{\alpha}^{\nu} = e^{2\Phi} g_{\mu}^{\nu}.$$

These are seen to satisfy (1.11), with,

$$(2.4) \quad \varrho = 3 e^{2\Phi} \dot{\Phi}^2.$$

The condition (1.7) gives,

$$(2.5) \quad \Phi = \frac{1}{3} \log (3t + c).$$

where c is some constant.

The metric (2.1) now becomes,

$$(2.6) \quad ds^2 = \frac{1}{(3t + c)^{2/3}} dt^2 - (3t + c)^{2/3} (dx^2 + dy^2 + dz^2)$$

We see that this metric is regular everywhere except for values of $t = -c/3$.

3. **Solutions.** Working directly with the electromagnetic field tensor we find the solutions of the EINSTEIN-MAXWELL field equations.

From (1.1) and (2.2), we have the following equations :

$$(3.1) \quad g^{22} (F_{12}^2 + F_{13}^2 + F_{23}^2) + g^{00} (F_{01}^2 + F_{02}^2 + F_{03}^2) = \frac{e^{4\Phi}}{4\pi} 3 \dot{\Phi}^2$$

$$(3.2) \quad F_{13} F_{03} + F_{12} F_{02} = 0,$$

$$(3.3) \quad F_{12} F_{01} - F_{23} F_{03} = 0,$$

$$(3.4) \quad F_{13} F_{04} + F_{23} F_{04} = 0,$$

$$(3.5) \quad g^{22} F_{12}^2 - g^{00} F_{03}^2 = \frac{e^{4\Phi}}{4\pi} (5\dot{\Phi}^2 + 2\ddot{\Phi}),$$

$$(3.6) \quad g^{22} F_{13}^2 - g^{00} F_{01}^2 = \frac{e^{4\Phi}}{4\pi} (5\dot{\Phi}^2 + 2\ddot{\Phi}),$$

$$(3.7) \quad g_{22}^2 F_{13}^2 - g^{00} F_{02}^2 = \frac{e^{4\Phi}}{4\pi} (5\dot{\Phi}^2 + 2\ddot{\Phi}).$$

Equations (3.2) and (3.5) give,

$$(3.8) \quad F_{03}^2 (g^{22} F_{23}^2 - g^{00} F_{01}^2) = F_{01}^2 \frac{e^{4\Phi}}{4\pi} (5\dot{\Phi}^2 + 2\ddot{\Phi}).$$

From (3.6) and (3.8), we have,

$$(3.9) \quad F_{03}^2 = F_{01}^2.$$

Also, from (3.4), (3.6) and (3.7),

$$(3.10) \quad F_{02}^2 = F_{01}^2.$$

From (3.9) and (3.10),

$$(3.11) \quad F_{01}^2 = F_{02}^2 = F_{03}^2.$$

Similarly, (3.5), (3.6) and (3.7) give,

$$(3.12) \quad F_{12}^2 = F_{13}^2 = F_{23}^2.$$

From (3.11) and (3.12), using (3.1), we have,

$$(3.13) \quad g^{22} F_{12}^2 + g^{00} F_{03}^2 = \frac{e^{2\Phi}}{4\pi} \dot{\Phi}^2.$$

(1.7), (3.5) and (3.13) now give,

$$(3.14) \quad F_{12} = 0$$

and

$$(3.15) \quad F_{03} = \pm \frac{\dot{\Phi} e^{\Phi}}{\sqrt{\pi}}.$$

Hence we have,

$$(3.16) \quad F_{12} = F_{13} = F_{23} = 0$$

and

$$(3.17) \quad F_{01} = F_{02} = F_{03} = \pm (3t + c)^{1/3}.$$

It is seen that, with these values, the equations (1.2) and (1.3) hold.

We find that, if ϕ is taken as a function of t alone, there is no magnetic induction but there is an electric field of equal strength in the x , y and z directions. As t increases, the electric field also increases.

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ÖZET

EINSTEIN-MAXWELL denklemleri, boşlukta ve uzayca homogen olan elektromagnetik alanlar için göz önüne alınmaktadır. Belirli bazı zaman anlar hariç, metriğin her yerde regüler olduğu, magnelik alanın olmadığı ve x , y ve z doğrultularında eşil kuvveti bulunan ve zaman ile birlikte artan bir elektrik alanının varlığı gösteriliyor.