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# **Research Article**

# Effect of aperture averaging on four petal Gaussian beams in atmospheric turbulence

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#### ARTICLE INFO

# ABSTRACT

Article history: Received 01 October 2020 Revised 06 February 2021 Accepted 06 March 2021 *Keywords:* Aperture averaging Atmospheric turbulence Four petal Gaussian beam Scintillation Aperture averaged scintillation of four petal Gaussian beam is studied in this article. Split step propagation approach which is used in wave propagation applications is selected to model atmospheric turbulence. Results are plotted in two types. First type is the analysis of aperture averaged scintillation versus propagation distance for constant receiver aperture. Second ones involve scintillation performance applying aperture averaging at constant distance. All results are compared with Gauss beam since commercial lasers generally radiates in Gaussian distribution. We observe that four petal Gaussian beam becomes more advantageous under moderate turbulence than weak one. In other point of view, it is possible to obtain less scintillation index by increasing beam order. Our results are applicable optical applications operating in atmosphere.

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# 1. Introduction

Because of the recent developments in technology, communication systems with low latency and high data rate need to arise. Related with this, 5G and beyond technologies play a vital role to meet this demand. Free space optics systems are the one of the components of 5G systems. Free space optics systems provide low latency and higher bit rate but performance of these systems is dependent on the atmospheric conditions. One way to overcome the negative effects of atmosphere is beam shaping. Utilizing non-conventional beams, performance of free space optics systems can be improved.

Propagation of different types of beams through random media attracts the attentions of scientists in different points of views. Intensity profiles, scintillation behavior, beam size, and coherency are some of these points [1]. Wave propagation lies in the background of this analysis. Solution of Huygens-Fresnel integral gives the received field through free space or turbulence [2]. For complex source field expressions, split step propagation methods are also used [3]. In some studies, accuracy of this method is also increased [4]. Benefiting from these methods scintillation index which is the most effective noise factor [5] is measured for untraditional beams. While scintillation index of sine hollow beam is less than Gauss beam [6], less scintillation index than sine hollow beam can be provided by Mathieu-Gauss beam [7]. In addition, we prove that cylindrical-sinc Gaussian beam has less scintillation index than Gauss beam [8]. Besides this this type of beam has a diverging nature propagating in atmospheric turbulence [9]. Considering the effect of scintillation, authors show how to detect information using Gaussian vortex beam [10]. In addition to above beams, there are some other types of beams that provides less scintillation. Regarding with this, it is shown in [11, 12] that Airy and partially coherent Airy beams have less scintillation than Gauss beam. Similarly, scintillation index of truncated flat-topped beam is low [13]. Poynting vector of Weber beam is calculated in [14]. For high frequency wave propagation, the Eulerian Gaussian beam method is generalized in [15].

In addition to methods and some applications listed above, four petal Gaussian beam is introduced by propagating through ABCD system [16]. If Vortex is added to four petal Gauss beam, it is observed that topological charge is quite dominant on the central hollow in far field [17]. While four petal Gaussian vortex beam

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clockwise phase distribution through the shows propagation in uniaxial crystal [18], four petal Lorentz-Gauss beam turns into elliptical Gaussian beam in the same medium [19].  $M^2$ , propagation factor for paraxial region is derived in [20]. Four petal Gaussian beam is propagated through fractional Fourier system and it is concluded as fraction, aperture, and source parameters are effective in received field [21]. Besides these, free space propagation and propagation in turbulent medium is also another interesting field for wave propagation. In the light of this, raise in topological charge provides larger hollow in the center for four petal Lorentz-Gauss vortex beam [22]. Similarly, larger topological charge reduces the effect of atmospheric turbulence on partially coherent elliptic Gaussian vortex beams [23]. Partially coherent four petal Gaussian beam shows resistance to oceanic turbulence if it is generated with larger beam order [24]. Finally, it is shown in [25] that four petal Gaussian vortex beam having incoherency evolves into Gaussian like shape easier if strength of turbulence increases. Airy transform is applied to four petal Gaussian beam and Airy like beam is obtained at the output [26].

Bearing in mind above literature review, we study the aperture averaged scintillation for four petal Gaussian beam under weak and moderate atmospheric turbulent regime. We use split step propagation to model the atmosphere. This model is used for propagate wave through turbulent medium. We analyzed the results with respect to propagation distance and size of receiver aperture. We anticipate that our results will be useful for optical communication and range measurement systems' designers.

# **2.** Source Field Distribution and Method to Measure Performance

#### 2.1 Four petal gaussian beam

Source field expression is written for four petal Gaussian beam [16] as

$$u_{s}(s_{x}, s_{y}) = (\frac{s_{x}s_{y}}{\alpha^{2}})^{2n} \exp(-\frac{s_{x}^{2} + s_{y}^{2}}{\alpha^{2}})$$
(1)

where n = 0, 1, 2, 3.... refer to beam order,  $s_x, s_y$  are transverse plane source coordinates, and  $\alpha = \sqrt{\alpha_x^2 + \alpha_y^2}$ being the Gaussian source size. Based on equation.1, Figure 1 and 2 are plotted to show the effect of beam order and Gaussian source size on the initial plane. It is seen from Figure 1 that while beam order increases, hollow region in the center gets larger. When beam order is set to 1, effect of Gaussian source size is seen in Figure 2. For asymmetric case  $\alpha_x > \alpha_y$ , beam lies along x-axis. Similar idea is valid for  $\alpha_y > \alpha_x$ . Furthermore, beam spreads on transverse plane symmetrically if  $\alpha_x, \alpha_y$  are increased in the same amount.



Figure 1. Transverse source plane intensity distribution for constant Gaussian source size



Figure 2. Transverse source plane intensity distribution for constant beam order

#### 2.2. Measurement of scintillation in atmosphere

Numerical split step propagation set-up involves special values presented in table 1 below. In this set-up, between the screens, free space conditions are satisfied. We chose modified von-Karman power spectral density since it is the closest model to Hill spectrum. Refractive index structure constant is the parameter to determine the strength of turbulence and it is taken into account in  $r_0$ . To satisfy the averaging in Eq. 4, number of realizations is taken as 500 which corresponds to infinity in reality. To avoid the complexity of computation of Huygens-Fresnel integral, this method benefits from convolution property of Fourier transform as it is seen from Eq. 2.

After applying all settings, received field after one step is found as

$$u_r(r_x, r_y, L) = A \ge F^{-1} \begin{pmatrix} F\left[ \left(\frac{s_x s_y}{\alpha^2} \right)^{2n} \exp\left(-\frac{s_x^2 + s_y^2}{\alpha^2}\right) \right] \\ F\left\{ exp\left[ \frac{jk}{2L} \left(s_x^2 + s_y^2\right) \right] \right\} \end{pmatrix}$$
(2)

where

$$A = \frac{-jk \exp(jkL)}{2\pi L} \exp\left[\frac{jk}{2L}(r_x^2 + r_y^2)\right] \exp[\psi(r)]$$
(3)

here *k* denotes wave number,  $r_x$ ,  $r_y$  are receiver plan coordinates, *F* and  $F^{-1}$  refer to Fourier and inverse Fourier transform, and  $\psi(r)$  indicates phase fluctuations due to atmosphere and it involves power spectral density. Then, aperture averaged scintillation is evaluated benefiting from received field as:

$$m^{2} = \frac{\left\langle P(r_{x}, r_{y})^{2} \right\rangle}{\left\langle P(r_{x}, r_{y}) \right\rangle^{2}} - 1 = \frac{\left( \int u_{r}(r_{x}, r_{y}) u_{r}^{*}(r_{x}, r_{y}) dS \right)^{2} / N}{\left( \int u_{r}(r_{x}, r_{y}) u_{r}^{*}(r_{x}, r_{y}) dS / N \right)^{2}}$$
(4)

where \* indicates complex conjugate. Adapted model of above mathematical equations is presented in Figure 3.

#### 3. Results and Discussions

In this part of the study, we comment on the numerical results.

Table 1. Numerical set-up	parameters
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Source plane dimensions(S)	10cm X 10cm
Propagation distance(L)	Up to 5500m
Power spectral density	$\phi = 0.023 r_0^{-\frac{5}{3}} \frac{\exp\left(-\left(\frac{f}{f_m}\right)^2\right)}{(f^2 + f_0^2)^{\frac{11}{6}}}  [27]$
Refractive index structure constant $(C_n^2)$	For weak turbulence: $10^{-14}m^{-\frac{2}{3}}$ For moderate turbulence : $10^{-13}m^{-\frac{2}{3}}$
Grid size in transverse plane	512 X 512
Inner scale frequency $(f_m)$	$\frac{5.92}{l_0 2\pi}, \ l_0 \to 0$
Outer scale frequency $(f_0)$	$\frac{1}{L_0},  L_0 \to \infty$
Fried parameter	$r_0 = (0.423(\frac{2\pi}{\lambda})^2 C_n^2 L)^{-\frac{3}{5}}$
Operating wavelength $(\lambda)$	1550nm
Number of screens	21: 0-1200m
	51: 1200-3200m
	91: 3200-5500m
Receiver plane dimensions in Figures 4 and 5	25 X 25 grids
Propagation distance in Figures 6 and 7	3300m
Realization amount for averaging (N)	500



Figure 3. Propagation model through the atmosphere



Figure 4. Aperture averaged scintillation of selected beam versus propagation path under weak turbulence



Figure 5. Aperture averaged scintillation of selected beam versus propagation path under moderate turbulence

We show aperture averaged scintillation index variations against propagation distance in Figures 4 and 5. It is seen from these figures that while only second order symmetric beam has advantageous as compared to Gauss beam in weak turbulence, all settings of four petal Gaussian beams have lower scintillation index for stronger turbulence regime.

In moderate turbulence case, scintillation index decreases as beam order raises. Additionally, increase in source size provides similar result with larger beam order. Besides this, asymmetry, without considering the direction, brings scintillation reduction.

In order to show the advance of four petal Gaussian beam, Table 2 is given. As it is observed from this table, however scintillation index of fourth order Four petal Gaussian beam is higher than Gaussian beam at close distance, significant amount of scintillation reduction can be provided by Four petal Gaussian beam at longer distances. At 5 km, scintillation of Gaussian beam is approximately two times higher than Four petal Gaussian beam.



versus receiver aperture under weak turbulence.

Distance(km)	Gauss beam	Fourth order
		Four petal
		Gaussian beam
0.4	0.001086	0.02483
0.6	0.006571	0.02489
0.8	0.01256	0.02586
1	0.01903	0.02767
1.2	0.02592	0.03022
1.4	0.03321	0.03343
1.6	0.04084	0.03723
1.8	0.04879	0.04151
2	0.05071	0.04621
2.6	0.0829	0.06194
3	0.1008	0.07293
3.6	0.1278	0.08853
4	0.1454	0.09734
5	0.1859	0.109

Table 2. Aperture averaged scintillation index under moderate turbulence.



Figure 7. Aperture averaged scintillation of selected beam versus propagation receiver aperture under moderate turbulence

In other point of view, four petal Gaussian beam has lesser scintillation index with versus receiver aperture opening as it is seen from Figures 6 and 7. We see that second and fourth order beam has lesser scintillation index for small receiver apertures. In addition, large source size four petal Gaussian beam has similar performance with Gauss beam under weak turbulence conditions. First order beam has the largest scintillation index. Then asymmetric ones and third order beams follow it.

As it is wanted, scintillation mitigation is seen for moderate turbulence as it is shown in Figure 7. We see that scintillation index of all kinds of four petal beams is less than Gauss beam for small apertures. In advantageous region, from the lowest scintillation index up to the highest one beams are listed as third, fourth order beams, asymmetric and large source size beams, second, and first order beams. We can investigate that four petal beam has less point like scintillation index as compared to Gauss beam since it can be evaluated in small aperture openings. This advantage vanishes when aperture size gets wider.

# 4. Conclusion

Aperture averaged scintillation of four petal Gaussian beam is studied in this article. Numerical results show that beam order and scintillation index is inversely proportional to each other. Additionally, scintillation index of four petal Gaussian beam is quite advantageous as compared to Gaussian beam for small aperture openings under moderate turbulence. In addition, four petal Gaussian beam is resistive to atmospheric turbulence. Because, scintillation index of all selected types four petal Gaussian beam is becomes advantageous when turbulence strength raises. Consequently, in the light of these investigations, four petal Gaussian beam can be selected as source beam for optical wireless communication and LIDAR systems operating in turbulence. In the future, our studies will focus on to measure the scintillation of other beams and generation of four petal beam experimentally.

# Declaration

The author declared no potential conflicts of interest with respect to the research, authorship, and/or publication of this article. The author also declared that this article is original, was prepared in accordance with international publication and research ethics, and ethical committee permission or any special permission is not required.

# **Author Contributions**

M. Bayraktar is responsible for all parts of the study.

#### Nomenclature

$\alpha = \sqrt{\alpha_x^2 + \alpha_y^2}$	:	Gaussian source size(m)		
$C_n^2$	:	Refractive index structure constant		
$f_{\mathrm{m}}$	:	Inner scale frequency		
$f_{\circ}$	:	Outer scale frequency		
F	:	Fourier transform		
$F^{-1}$	:	Inverse Fourier transform		
k	:	Wave number $(m^{-1})$		
L	:	Propagation distance		
$L_s$	:	Receiver plane size		
$m^2$	:	Aperture averaged scintillation index		
n	:	beam order (0,1,2)		
$r_x, r_y$	:	Receiver plane coordinates(m)		
$S_x, S_y$	:	Source plane coordinates(m)		
S	:	Source plane size		
$u_{s}$	:	Source field expression		
u <sub>r</sub>	:	Received field expression		
$\psi(r)$	:	Phase fluctuations		
λ	:	Operating wavelength		

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